



Non-adiabatic Quantum Dynamics of the Dissociative Charge Transfer $\text{He}^+ + \text{H}_2 \rightarrow \text{He} + \text{H} + \text{H}^+$

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We present the non-adiabatic, conical-intersection quantum dynamics of the title collision where reactants and products are in the ground electronic states. Initial-state-resolved reaction probabilities, total integral cross sections, and rate constants of two H_2 vibrational states, $v_0 = 0$ and 1, in the ground rotational state ($j_0 = 0$) are obtained at collision energies $E_{\text{coll}} \leq 3$ eV. We employ the lowest two excited diabatic electronic states of HeH_2^+ and their electronic coupling, a coupled-channel time-dependent real wavepacket method, and a flux analysis. Both probabilities and cross sections present a few groups of resonances at low E_{coll} , whose amplitudes decrease with the energy, due to an ion-induced dipole interaction in the entrance channel. At higher E_{coll} , reaction probabilities and cross sections increase monotonically up to 3 eV, remaining however quite small. When H_2 is in the $v_0 = 1$ state, the reactivity increases by ~ 2 orders of magnitude at the lowest energies and by ~ 1 order at the highest ones. Initial-state resolved rate constants at room temperature are equal to 1.74×10^{-14} and to $1.98 \times 10^{-12} \text{ cm}^3 \text{ s}^{-1}$ at $v_0 = 0$ and 1, respectively. Test calculations for H_2 at $j_0 = 1$ show that the probabilities can be enhanced by a factor of $\sim 1/3$, that is *ortho*- H_2 seems ~ 4 times more reactive than *para*- H_2 .

Keywords: $\text{He}^+ + \text{H}_2$, wavepacket, conical intersection, non-adiabatic, quantum, dynamics

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INTRODUCTION

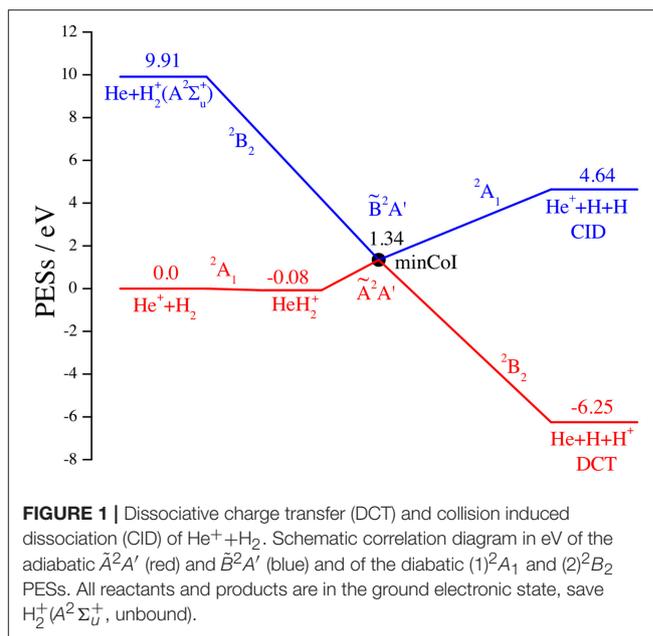
Atomic Hydrogen and Helium are the dominant chemical species of the early Universe (Galli and Palla, 2013) and of the interstellar medium, and are easily ionized by cosmic rays. Therefore, these atoms and their ions are the astrochemical fundamental reactants (Lepp et al., 2002), together with ubiquitous photons, giving first simple diatoms as H_2 , H_2^+ (Stancil et al., 1993), and HeH^+ (Zygelman et al., 1998) and then atom+diatom bimolecular collisions as $\text{He} + \text{H}_2^+$ (De Fazio et al., 2012), $\text{H} + \text{HeH}^+$ (De Fazio, 2014; Gamallo et al., 2015), and $\text{He}^+ + \text{H}_2$.

When all chemical species are in the ground electronic states, the $\text{He} + \text{H}_2^+$ reaction is endothermic and rather slow, but the other two are exothermic by about ~ 0.7 and 6.2 eV, respectively, as many ion+neutral astrochemical reactions (Herbst and Klemperer, 1973). $\text{H} + \text{HeH}^+$ is barrierless and gives quickly the $\text{He} + \text{H}_2^+$ products, but the collision-induced dissociative charge transfer (DCT) $\text{He}^+ (^2S) + \text{H}_2 (X^1\Sigma_g^+) \rightarrow \text{He} (^1S) + \text{H} (^2S) + \text{H}^+$ is very slow at low collision energy E_{coll} and at room temperature, if H_2 is in the ground vibro-rotational state. In fact, $\text{H} + \text{HeH}^+ \leftrightarrow \text{He}^+ + \text{H}_2$ occurs on the ground potential energy surface (PES) \tilde{X}^2A' of HeH_2^+ , which

is well-separated from the excited electronic species, but the low- E_{coll} DCT involves the first two excited adiabatic electronic states \tilde{A}^2A' and \tilde{B}^2A' of HeH_2^+ , which differ by a two-electron excitation and are coupled by a C_{2v} conical intersection (CI) (Preston et al., 1978; McLaughlin and Thompson, 1979). In the latter case the non-adiabatic coupling is weak and the lower, dissociative cone of the intersection seam gives rise to an adiabatic barrier with diabatic character, that is the DCT tends to follow the diabatic PESs without changing the electronic configuration. This strongly inhibits the reactivity that is associated with the tunneling through the barrier (Preston et al., 1978).

This is schematically shown in the correlation diagram of **Figure 1** for the \tilde{A}^2A' and \tilde{B}^2A' adiabatic and $(1)^2A_1$ and $(2)^2B_2$ diabatic PESs V of HeH_2^+ , where all chemical species are in the ground electronic states, save $\text{H}_2^+(A^2\Sigma_u^+)$ that is unbound, and the energy is referred to the reactant minimum. This diagram is obtained from the *ab initio* Multi-Reference Configuration-Interaction results of McLaughlin and Thompson (1979), changing the label of the reactant diabatic state 2A_1 from (5) to (1), and from the analytical fits of the associated diabatic PESs V_{11} and V_{22} by Aguado et al. (1993). In the scheme of **Figure 1** we omit the ground adiabatic PES, well below the excited PESs and with too small non-adiabatic couplings. However, all three adiabatic PESs are strongly coupled by intense laser pulses (Szidarovszky and Yamanouchi, 2016) when the ground PES becomes populated (Schauer et al., 1989). A more complete description of the reaction dynamics in presence of electric and magnetic fields is out of the scope of the present article. Note in the figure the small ion-induced dipole minimum in the C_{2v} entrance channel at $R(\text{He}-\text{H}_2) = 4.45 a_0$, $r(\text{H}-\text{H}) = 1.42 a_0$, and $V = -0.08 \text{ eV}$ and the CI C_{2v} minimum at $R = 4.89 a_0$, $r = 2.18 a_0$, and $V = 1.34 \text{ eV}$, owing to the intersection (Preston et al., 1978) between the $\text{H}_2(X^1\Sigma_g^+)$ and $\text{H}_2^+(A^2\Sigma_u^+)$ curves at $r = 2.19 a_0$ and $V = 1.28 \text{ eV}$.

Experimental studies on the $\text{He}^+ + \text{H}_2 \rightarrow \text{He} + \text{H} + \text{H}^+$ DCT dynamics date back to 1955 (Stedford and Hasted, 1955) and 1961 (Giese and Maier, 1961), when the integral cross sections (ICSS) were measured at $E_{\text{coll}} = 4 \text{ eV}$ and found $< 0.05 \text{ \AA}^2$. ICSS as functions of E_{coll} were then measured in many works up to 1996 (Dhuicq et al., 1996), finding values from 0.01 up to $\sim 2 \text{ \AA}^2$, for E_{coll} between 3 and 100 eV, by considering H_2 in the ground vibro-rotational state and all ground and excited open states of H. Below 3 eV, the H product is in the ground 1s state and the ICSS are so small and so difficult to measure that experimental values present large discrepancies (Reinig et al., 1992). However, the ICSS increase by one/two orders of magnitude if H_2 is excited by one vibrational quantum (Preston et al., 1978) or E_{coll} grows up to $\sim 100 \text{ eV}$. Differential cross sections were also measured (Dhuicq et al., 1998) at $E_{\text{coll}} > 9 \text{ eV}$, that is for the formation of the H^* excited product. Accordingly, small rate constants were observed (Johnsen et al., 1980), with values of $(1.5 \pm 0.15) \times 10^{-13}$ and $(1.1 \pm 0.1) \times 10^{-13} \text{ cm}^3 \text{ s}^{-1}$ at 78 and 330 K, respectively, that is nearly four order of magnitude lower than the Langevin (Gioumousis and Stevenson, 1958) estimates for ion+neutral barrierless and exothermic reactions.



A few works have also theoretically investigated the dynamics of the DCT collision since 1994, when Aguillon (1994) employed a semiclassical coupled wavepacket (WP) method and the analytical diabatic PESs of Aguado et al. (1993) for computing ICSS for ground and excited vibrational states ν_0 of H_2 , up to $\nu_0 = 4$ and in the E_{coll} range from 2 to 10 eV. He found total-ICSS values from $\sim 0.001 \text{ \AA}^2$ ($\nu_0 = 0$, $E_{\text{coll}} = 2 \text{ eV}$) to $\sim 1 \text{ \AA}^2$ ($\nu_0 = 3$, $E_{\text{coll}} = 10 \text{ eV}$), in agreement with the most recent measurements (Dhuicq et al., 1996). In a subsequent work (Aguillon, 1998), an improved version of this semiclassical method was used for obtaining new ICSS values and differential cross sections. Approximated theoretical models were also employed (Dhuicq et al., 1996) for explaining observed cross sections above 9 eV, when excited $\text{H}^*(n=2)$ was produced. Finally, quantum infinite-order-sudden cross sections were calculated (Martinez et al., 2000) for $\text{He}^+ + \text{H}_2 \rightarrow \text{He} + \text{H}^*(n \geq 2) + \text{H}^+$ at $E_{\text{coll}} \geq 10 \text{ eV}$, using the accurate diabatic representation of Aguado et al. (1993) and six more approximated diabatic electronic states (Sidis, 1996).

As far as we know, no further studies of the title reaction have been published and its collision dynamics below $\sim 2 \text{ eV}$ is unknown. In particular, only semiclassical or approximated quantum theoretical studies were carried out, although both conical intersection and barrier tunneling are purely quantum effects. We thus here report a rigorous time-dependent quantum study of the DCT $\text{He}^+ + \text{H}_2(\nu_0 = 0, 1) \rightarrow \text{He} + \text{H} + \text{H}^+$ with all species in the ground electronic state, at thermal and hyperthermal collision energy up to 3 eV, using the diabatic PESs of Aguado et al. (1993) and WP and flux formalisms. In section Theory and Calculations we present the theoretical method and its numerical details. Section Collision Results reports initial-state-resolved total reaction probabilities, ICSS, and thermal rate constants. Finally, we present our conclusions in section Conclusions.

THEORY AND CALCULATIONS

Potential Energy Surfaces and Non-adiabatic Coupling

As we said in the Introduction, the HeH_2^+ adiabatic and diabatic electronic states relevant in the present work were obtained *ab initio* in McLaughlin and Thompson (1979), fitted analytically in Aguado et al. (1993), and they are schematically plotted in **Figure 1**. We label the adiabatic species and PESs by \tilde{A}^2A' and \tilde{B}^2A' and by V_A and V_B , respectively. As already discussed (Aguado et al., 1993), these states belong to the fully symmetric irreducible representation for linear ($C_{\infty v}$) and non-symmetric (C_s) geometries, while **Figure 1** shows that they transform as A_1/B_2 and B_2/A_1 for perpendicular geometries (C_{2v}), before/after the CI, respectively, which rules the title reaction. On the other hand, we label the associated diabatic electronic states and PESs by (1) A_1 and (2) B_2 and by V_{11} , V_{22} , and V_{12} , respectively, where the third surface describes the CI non-adiabatic coupling in the diabatic representation.

In order to simultaneously describe these electronic species and take into account the CI, a fit of the diabatic PESs and coupling V_{11} , V_{22} , and V_{12} was made in Aguado et al. (1993). The description of the adiabatic states is obtained as the eigenvalues V_A and V_B of a 2×2 matrix $\begin{pmatrix} V_{11} & V_{12} \\ V_{12} & V_{22} \end{pmatrix}$, in which the interaction term V_{12} must have the correct symmetry, being anti-symmetric with respect to the permutation of the H atoms and thus vanishing identically for equal He–H distances. This coupling term V_{12} was fitted in Aguado et al. (1993) to reproduce the CI between the \tilde{A}^2A' and \tilde{B}^2A' adiabatic states.

The diabatic surfaces were fitted in Aguado et al. (1993) using the Aguado-Paniagua functional form (Aguado and Paniagua, 1992; Aguado et al., 1998) that expands the energy as a multidimensional permutationally invariant polynomial (Aguado et al., 1994, 2001) in Rydberg type variables (Rydberg, 1931) of the form $\rho_{AB} = R_{AB} \exp(-\beta_{AB}R_{AB})$, where A and B are two nuclei and R_{AB} is their distance. For the interaction term, a simple expansion in Rydberg functions,

$$V_{12}(R_{\text{HeH}}, R_{\text{HeH}'}, R_{\text{HH}'}) = C_{12} \rho_{\text{HH}'} (\rho_{\text{HeH}} - \rho_{\text{HeH}'}), \quad (1)$$

fulfills the anti-symmetric requirement.

The DCT is produced through the CI, as shown in **Figure 2** using nuclear Jacobi coordinates R , $r = R_{\text{HH}'}$, and γ , at $R = 4 a_0$ and $\gamma = 0$. The top panel shows the adiabatic \tilde{A}^2A' and \tilde{B}^2A' PESs of HeH_2^+ in the reactant channel, obtained from the diabatic V_{ij} surfaces, where \tilde{A}^2A' correlates with $\text{He}^+(^2S) + \text{H}_2(X^1\Sigma_g^+)$ for $r < 2 a_0$ and with $\text{He}(^1S) + \text{H}_2^+(A^2\Sigma_u^+)$, unbound for $r > 2 a_0$. To analyze the accuracy of the fit, we compare the analytical non-adiabatic coupling matrix elements (NACMEs) in the adiabatic representation, obtained from the fitted diabatic energies and coupling, with the *ab initio* results calculated using the MOLPRO program package (Werner et al., 2018). As done previously

for H_4^+ and H_5^+ in Sanz-Sanz et al. (2015), the analytical NACMEs can be calculated from the generalized Hellmann-Feynman theorem,

$$\langle \tilde{A}^2A' | \frac{\partial}{\partial R_{AB}} | \tilde{B}^2A' \rangle = \frac{1}{V_B - V_A} \langle \tilde{A}^2A' | \frac{\partial \hat{H}^{el}}{\partial R_{AB}} | \tilde{B}^2A' \rangle, \quad (2)$$

where \hat{H}^{el} is the electronic Hamiltonian and the rhs is obtained from the derivatives of the diabatic energies V_{ij} (Sanz-Sanz et al., 2015). *Ab initio* calculations have been performed using the Multi-Reference Configuration-Interaction method, with the aug-cc-pVTZ basis set of Dunning (1989) and Woon and Dunning (1994). The *ab initio* NACMEs are obtained using a first order difference method with an interval of $0.01 a_0$, and that for $R_{AB} = r$ is compared with the fitted one in the bottom panel of **Figure 2**. The agreement, in form and in position, between both is excellent, indicating that the equation used for the interaction term V_{12} is also appropriate to reproduce the NACMEs. The probability density of the first two vibrational states of $\text{H}_2(X^1\Sigma_g^+)$ has been included in the top panel of **Figure 2**, in order to analyze the reason why the reaction is faster when H_2 is vibrationally excited (Aguillon, 1994, 1998). While the maximum of the probability density of the vibrational state $v_0 = 0$ is found at $1.4 a_0$, those in the first excited vibrational state $v_0 = 1$ are at 1.25 and $1.73 a_0$. The latter is close to the region in which the NACME has its maximum, that explains the enhancement of the reactivity as we shall see in section Reaction Probabilities.

In **Figure 3** the ratio between the interaction term and the energy difference of the diabatic states, $V_{12}/(V_{22}-V_{11})$, is plotted as a function of the R_{HeH} and $R_{\text{HH}'}$ distances, for several values of the angle $\theta \leq 180^\circ$ between these distances. This ratio is important for the calculation of the adiabatic PESs from the fitted diabatic ones, according to:

$$V_{A/B} = \frac{1}{2} \left\{ V_{11} + V_{22} - / + [(V_{11} - V_{22})^2 + 4V_{12}^2]^{1/2} \right\}. \quad (3)$$

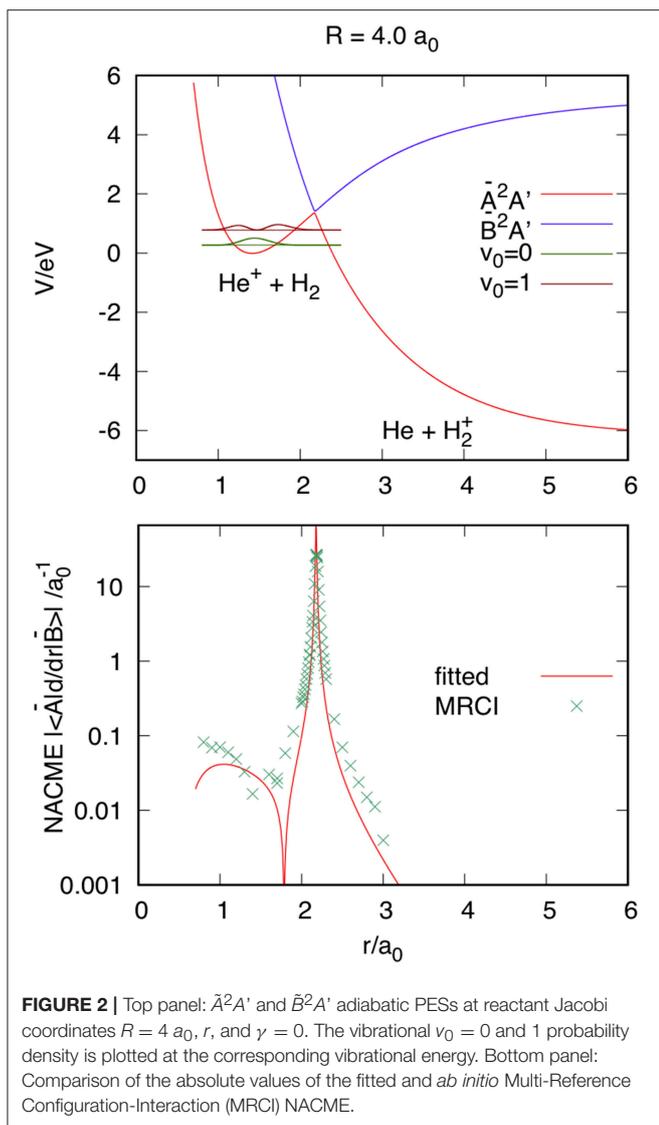
As expected, $V_{12}/(V_{22}-V_{11})$ changes sign and its absolute value is maximum in the region of the diabatic crossing line shown in Figures 2, 3 of Aguado et al. (1993).

In conclusion, these results show that the collisional dynamics of $\text{DCT He}^+(^2S) + \text{H}_2(X^1\Sigma_g^+) \rightarrow \text{He}(^1S) + \text{H}(^2S) + \text{H}^+$ can be investigated with high accuracy in the present diabatic electronic representation, using the V_{11} , V_{22} , and V_{12} PESs.

Collision Dynamics

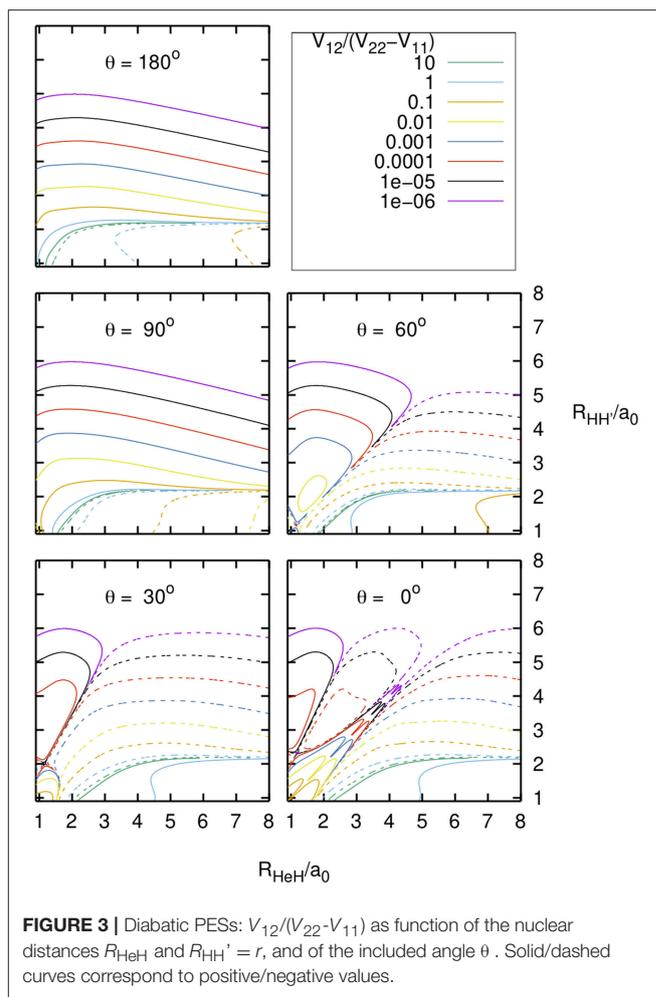
Since many years we are presenting our quantum theory (Petrongolo, 1988) and results of non-adiabatic effects in spectroscopy of triatomics and dynamics of atom+diatom collisions and we here report a brief summary, following our work on the CI dynamics of the $\text{OH}(A^2\Sigma^+) + \text{H}(^2S)$ reaction (Gamallo et al., 2013).

The $\text{He}^+ + \text{H}_2$ collision is described by reactant Jacobi coordinates R , r , and γ , by a body-fixed reference frame with the z axis along \mathbf{R} , and by a HeH_2^+ spinless rovibronic Hamiltonian \hat{H} , which contains the electronic \hat{H}^{el} and the total angular



momentum \hat{J} operators. \hat{H} is represented in an orthonormal basis of diabatic electronic states $(1)^2A_1$ and $(2)^2B_2$, radial grid $|Rr\rangle$, associated Legendre $|jK\rangle$, and symmetry Wigner states $|K+p\rangle$. Here $(1)^2A_1$ and $(2)^2B_2$ are coupled by \hat{H}^{el} owing to the CI, $\hbar K$ is the \hat{J}_z eigenvalue, and we omit J and its space-fixed Z component in $|K+p\rangle$, where the total parity is $p = (-)^{J+K_{\min}}$ with $K_{\min} = 0$ or 1. The $2J+1$ values of K are thus factorized in two non-interacting groups, with $K_{\min} \leq K \leq J$, of dimensions $J+1$ or J according to $K_{\min} = 0$ or 1, respectively.

Initial-state-resolved reaction probabilities are computed through the quantum, real WP formalism of Gray and Balint-Kurti (1998) and Meijer et al. (1998), essentially equal to the Chebyshev approach of Guo (2012). Shortly, an *arccos* mapping of the HeH_2^+ time-dependent Schrödinger equation is solved recursively, using a scaled and shifted Hamiltonian \hat{H}_s and starting from an initial and complex WP $|\psi_0\rangle = |a_0\rangle + i|b_0\rangle$ (Gray and Balint-Kurti, 1998). This initial WP describes



the entrance channel $\text{He}^+(^2S) + \text{H}_2(X^1\Sigma_g^+)$, with the diabatic electronic state $(1)^2A_1$ and the R -dependent term

$$g_0(R) = \pi^{-1/4} \alpha^{-1/2} \exp[-(R - R_0)^2/2\alpha^2] \exp[-i(2\mu_R E_0)^{1/2} (R - R_0)], \quad \text{in atomic units,} \quad (4)$$

where μ_R is the $\text{He}^+ + \text{H}_2$ reduced mass. The r and angular terms of $|\psi_0\rangle$ are the vibrational $|v_0\rangle$ and rotational $|j_0 K_0\rangle$ states of $\text{H}_2(X^1\Sigma_g^+)$, and finally $|K_0+p\rangle$ is the initial overall rotational species. The recursions are

$$|a_1\rangle = \hat{H}_s |a_0\rangle - \left(1 - \hat{H}_s^2\right)^{1/2} |b_0\rangle, \quad \text{first complex propagation,} \quad (5)$$

$$|a_{n+2}\rangle = 2\hat{H}_s |a_{n+1}\rangle - |a_n\rangle, \quad \text{other real propagations,} \quad (6)$$

where the square root in Equation (5) is evaluated with a Chebyshev expansion, and Equation (6) is a standard Chebyshev propagation of just a real WP, which is also absorbed at $R > R_{\text{abs}}$ and $r > r_{\text{abs}}$ by the Gaussians $\exp[-C_{\text{abs}}^R (R - R_{\text{abs}})^2]$ and $\exp[-C_{\text{abs}}^r (r - r_{\text{abs}})^2]$, respectively. At the end of the propagation, we

obtain the probability via a time-to-energy Fourier transform and a flux analysis (Meijer et al., 1998) on the $(2)^2B_2$ PES.

We compute non-adiabatic initial-state-resolved reaction probabilities $P_{v_0}^J(E_{\text{coll}})$, with the initial WP on the HeH_2^+ $(1)^2A_1$ diabatic PES and $\text{H}_2(X^1\Sigma_g^+)$ in the ground and first excited vibrational state, $v_0 = 0$ and 1, and in the ground rotational state $j_0 = 0$. Then the initial K_0 is equal to $K_{\text{min}} = 0$ and the total parity p is $(-)^J$. Here we have

$$P_{v_0}^J(E_{\text{coll}}) = \sum_{v'j'K'} \left| S_{2v'j'K' \leftarrow 1v_000}^J(E_{\text{coll}}) \right|^2, \quad (7)$$

where (1) and (2) are the diabatic electronic states, v' , j' , and K' label open vibrational, rotational, and helicity states of the products, respectively, and S^J is the state-to-state parity adapted S-matrix at J . The total ICS is then defined by

$$\sigma_{v_0}(E_{\text{coll}}) = \frac{\pi}{2\mu_R E_{\text{coll}}} \sum_J (2J+1) P_{v_0}^J(E_{\text{coll}}), \quad (8)$$

and the initial-state-resolved rate constant is

$$k_{v_0}(T) = \left(\frac{8}{\pi\mu_R k_B^3 T^3} \right)^{1/2} \int_0^\infty E_{\text{coll}} \sigma_{v_0} \exp(-E_{\text{coll}}/k_B T) dE_{\text{coll}}, \quad (9)$$

where T is the temperature and k_B is the Boltzmann constant.

The calculations are done up to $J = 150$, using the coupled-channel formalism with Coriolis couplings among the K values, up to K_{max} , that are necessary for converging the probabilities. Considering the H–H permutation symmetry, the numerical parameters of the WP propagations are listed in **Table 1**: they correspond to 7,053,300 basis states at $K = 0$, including two coupled electronic states. These parameters span the whole range of E_{coll} from 0.0002 to 3 eV, with $\Delta E_{\text{coll}} = 0.0001$ or 0.001 eV below or above 0.2 eV, respectively. The most important parameters that affect the convergence of the calculations are the number of the propagation kilo-steps, $kstep$, the maximum K value, K_{max} , and the dimension of the R grid, nR . We shall see in the next sections some convergence results with respect to $kstep$ and K_{max} . The probabilities obtained with $nR = 461$ of **Table 1** are practically indistinguishable by those corresponding to $nR = 559$. All calculations are done with our J - K -parallelized Open MPI codes.

COLLISION RESULTS

Reaction Probabilities

We plot in **Figure 4** the opacities functions $(2J+1)P_{v_0}^J(E_{\text{coll}})$ at $J = 20$ and 40, $v_0 = 0$ and 1, $E_{\text{coll}} \leq 3$ eV, and $K_{\text{max}} = 0, 1$, and 2, with $kstep = 5$, which converges the probabilities within 0.1% above ~ 0.5 eV. The curves at $v_0 = 0$ and 1 present a similar behavior, namely narrow resonance features at $J = 20$ and $E_{\text{coll}} \leq 0.5$ eV and a smooth, fast increase above. From the comparison among lower and upper panels we see that $v_0 = 1$ is more reactive

TABLE 1 | Parameters of the quantum dynamics calculations.

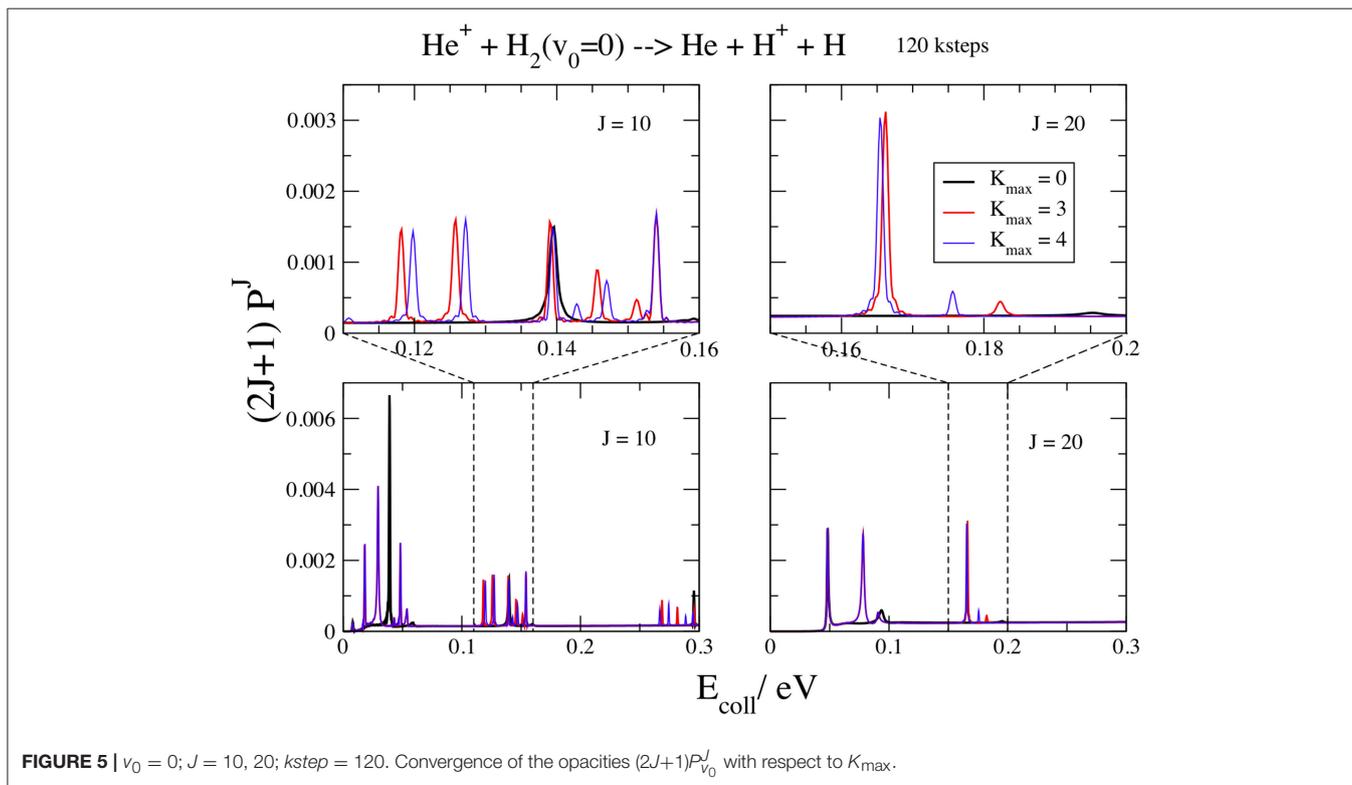
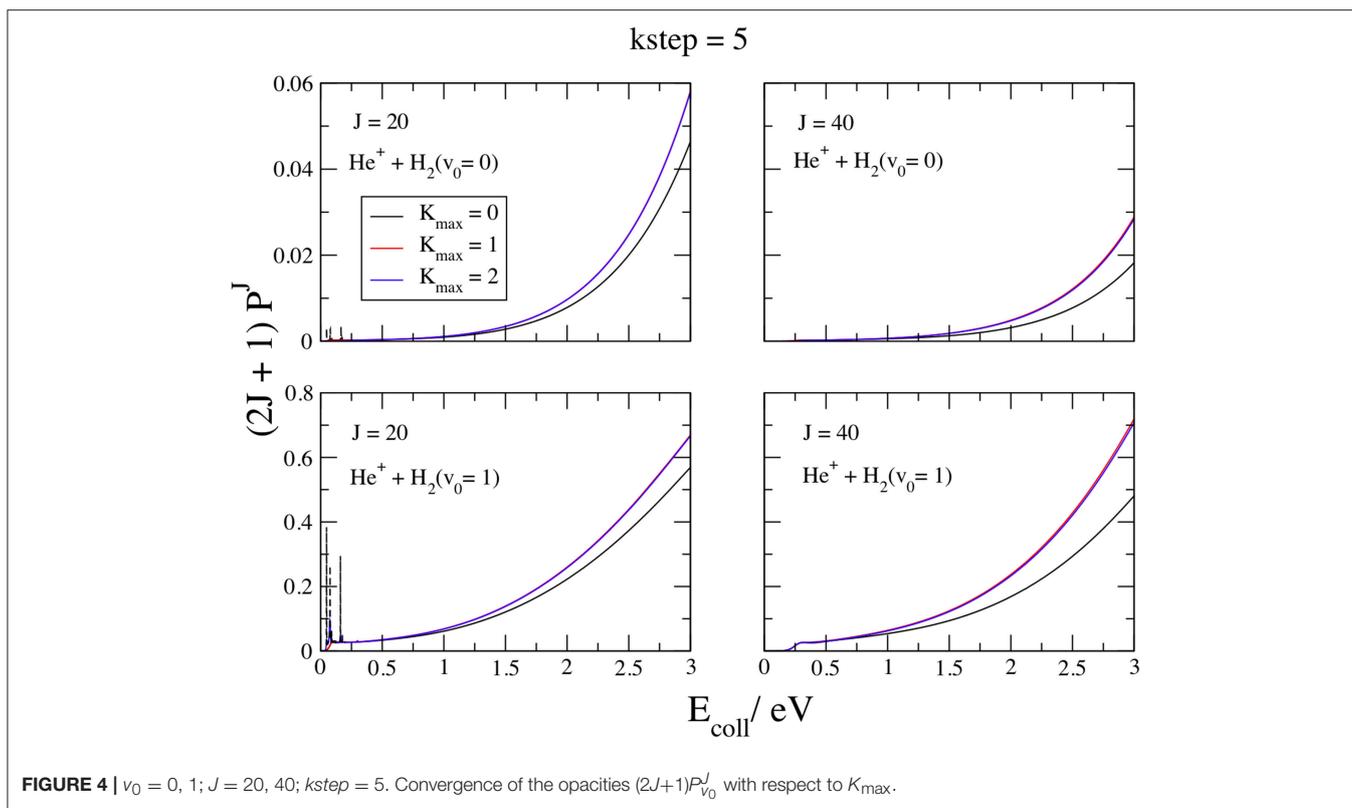
Initial Gaussian $g_0(R)$, Equation (4), α , R_0 , and E_0	0.2, 20, and 1.5 eV
R range and number of grid points	0.5–45 and 461
r range and number of grid points	0.5–15 and 153
Number of (associated) Legendre functions $ j_0\rangle$	50 ($j = \text{even}$)
R and r absorption start at	30 and 11
R and r absorption strength	0.0005 and 0.005
Flux analysis at r	10

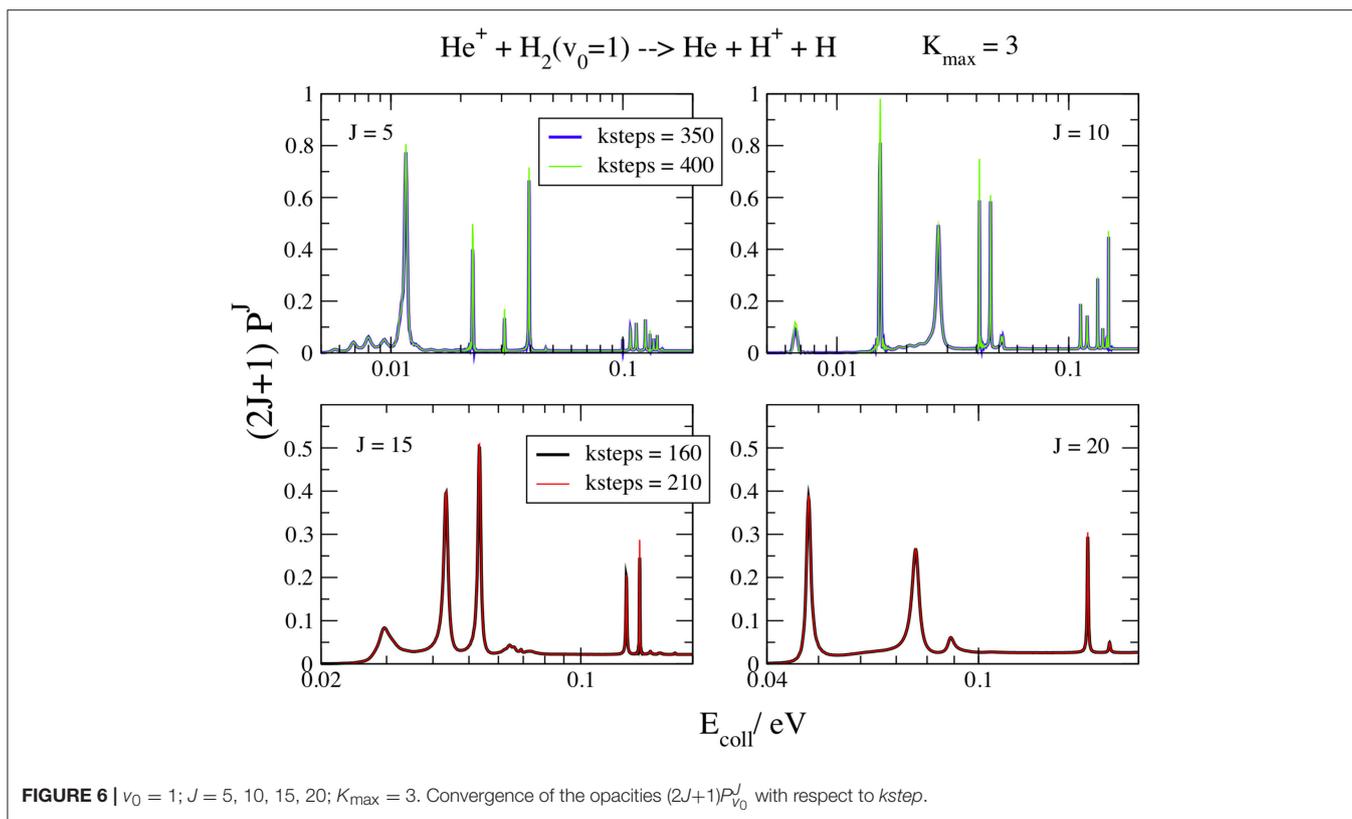
Values in atomic units, unless otherwise specified.

than $v_0 = 0$ by \sim one order of magnitude. Of course, this finding reflects the HeH_2^+ (\tilde{A}^2A') early potential barrier of 1.34 eV and the $v_0 = 1$ density maximum at 1.73 a_0 (**Figure 2**), close to the NACME maximum. Comparing the curves inside each panel, we observe a very fast convergence in K_{max} , so that the Coupled-State approximation (McGuire and Kouri, 1974; Pack, 1974) at $K_{\text{max}} = 0$ gives already reasonable results. Just two K s converge the plots within the graphical accuracy, implying that the Coriolis couplings between K and $K \pm 1$ vanish quickly when the initial K_0 is equal to zero. To reach the high accuracy claimed above at all the partial waves required to give convergent ICSs in all the collision energy range considered, we employ $K_{\text{max}} = 2$ and 3 for $v_0 = 0$ and 1, respectively. In the $J = 20$ left panels, the dashed lines show the convergent results below 0.5 eV and the weight of the resonances with respect to the background. The resonance features are probably due to rotational metastable states of the C_{2v} ion-induced dipole minimum in the entrance channel. These states, embedded in the collisional continuum, are particular Feshbach resonances induced by the CI (Cederbaum and Friedman, 2003), with lifetimes of the order of the nanosecond. The $v_0 = 0$ opacities then increase in a monotonic way above 0.6 eV up to a maximum value of ~ 0.06 at 3 eV, that is one-two orders of magnitude larger than the strongest resonance. The $v_0 = 1$ opacities show the same behavior, but the value at 3 eV is just two times larger than the resonance maximum. However, to converge the resonance energy pattern very different values of the convergent parameters K_{max} and $kstep$ are required.

We show in **Figure 5** the opacities at $J = 10$ and 20, for $v_0 = 0$, $E_{\text{coll}} \leq 0.3$ eV, $K_{\text{max}} = 0, 3$, and 4, and $kstep = 120$. In the upper panels a blow up of 0.05 eV also shows the details of the resonance features. From this figure we can observe that the convergence in K_{max} is slower than for high energies. The curves at $K_{\text{max}} = 0$ exhibit less peaks and the resonances shift in energy by increasing this parameter. Three main groups of resonances could be identified, decreasing and then disappearing at high E_{coll} . This result suggests that the different peaks inside each group correspond to different bending energies of the collision complex (Aquilanti et al., 2005) while different groups correspond to different energies of the symmetric stretching motion. The $v_0 = 1$ curves have a similar behavior, with the K_{max} convergence slightly faster and the resonances at lower energies.

In **Figure 6** the resonance energy patterns at $J = 5, 10, 15$, and 20, $v_0 = 1$, and $K_{\text{max}} = 3$ are plotted for $E_{\text{coll}} \leq 0.11$ eV at different $kstep$ values. The K_{max} value gives a graphical convergence to



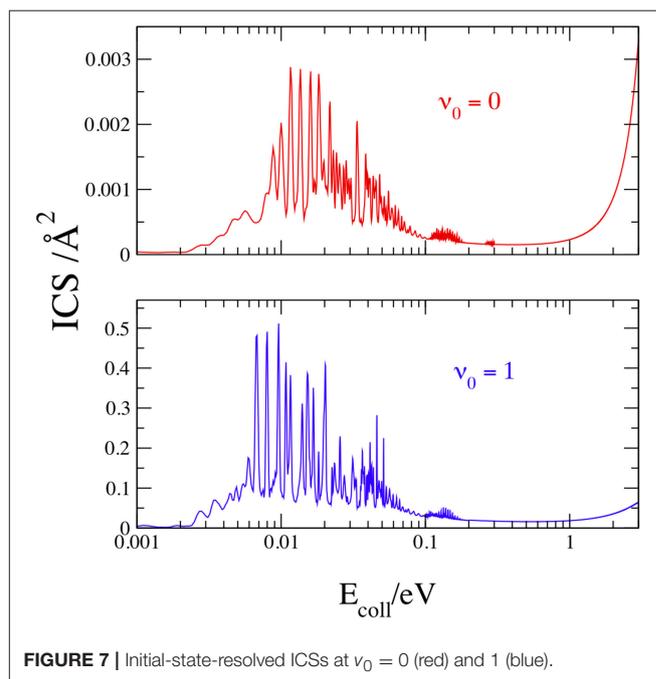


the curves and the logarithmic scale of the energy points out the resonance details. Here $kstep$ is much greater than for the background and changes markedly with J and E_{coll} . In fact, $kstep = 350$ is still not enough to converge all the $J = 5$ and 10 features, especially below 0.01 eV, but $kstep = 160$ perfectly converges the resonances at $J = 15$ and 20 . As expected, the slowest convergent resonances are the narrowest ones, with largest lifetimes (Aquilanti et al., 2004). Moreover, the lifetimes of the collision complexes decrease at high J and E_{coll} , because the centrifugal barrier increases and the resonances are less trapped inside the shallower well. Above $J = 25$ the well is so shallow that it does not support any resonance and the features disappear. Of course, increasing $kstep$ means increase the collision time that must be of the same order of magnitude of the lifetime of the resonance intermediate to give convergent results.

Integral Cross Sections and Rate Constants

We plot in **Figure 7** the total ICS at $v_0 = 0$ and 1 , with $J \leq 150$ to obtain convergent results up to 3 eV, noting that the convergent requirements change markedly with E_{coll} and J . To minimize the computational effort, the total number of partial waves was therefore shared in different groups, and different values of K_{max} and $kstep$ were employed in each group, as shown in **Table 2**.

Notwithstanding the large $kstep$ used for the lowest partial waves, test calculations show that some narrow ICS peak slightly increases with more steps. With these input data, a convergence within 0.1% is reached only by the resonance peaks above



0.05 eV. The results in **Figure 7** confirm the general scenario of the opacities: sharp resonances below 0.3 eV and a monotonic increase at larger E_{coll} . The J sum now merges the resonances

TABLE 2 | Parameters of ICS calculations.

ν_0	J	$kstep$	K_{max}
0	0–24	160	4
	25–52	40	3
	53–82	20	3
	83–150	10	1
1	0–14	350	3
	15–25	160	3
	26–38	110	3
	39–102	20	2
	103–150	10	1

TABLE 3 | Present quantum ICSs/ \AA^2 vs. those semiclassical (Aguillon, 1998).

ν_0	E_{coll}/eV	present work	(Aguillon, 1998)
0	2	8.7×10^{-4}	$(\sim 7 \times 10^{-4})$
	3	3.3×10^{-3}	$(\sim 3 \times 10^{-3})$
1	2	3.5×10^{-2}	$(\sim 3.5 \times 10^{-2})$
	3	6.5×10^{-2}	$(\sim 6 \times 10^{-2})$

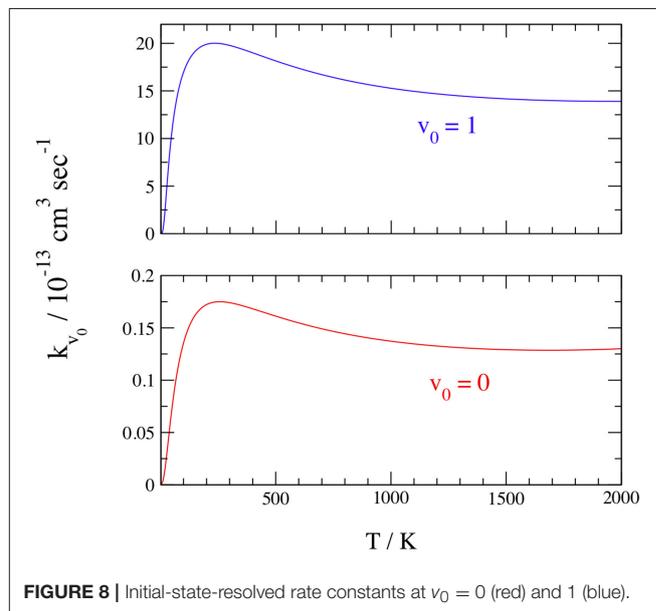
in two main groups, the first and stronger below 0.1 eV, and the latter and weaker from 0.1 to 0.3 eV. At $\nu_0 = 0$, the strongest resonance is equal to 0.003\AA^2 at 0.0116 eV and the maximum at 3 eV has nearly the same value. At $\nu_0 = 1$, the strongest resonance is larger than the value at 3 eV by \sim one order of magnitude, and is equal to 0.455\AA^2 at 0.0096 eV. In conclusion, the one-quantum vibrational excitation of $\text{H}_2(X^1\Sigma_g^+)$ increases the cross section by more than two orders of magnitude at ~ 0.01 – 0.02 eV and by ~ 20 times at 3 eV.

As said in the Introduction, only the semiclassical theoretical works by Aguillon (1994, 1998) have obtained ICSs at or below 3 eV. We present in **Table 3** a comparison between our quantum ICSs and those estimated from **Figure 4** of Aguillon (1998), showing the good agreement between the results that differ at the most by $\sim 24\%$ at $\nu_0 = 0$ and $E_{coll} = 2$ eV. Taking into account the different theoretical treatment and the small reactivity at these conditions, this implies that the Aguillon semiclassical treatment of the DCT reaction works remarkably well at this collision energies, where the ICSs do not present any quantum resonance.

Finally, we report in **Table 4** and **Figure 8** the initial-state-resolved rate constants $k_{\nu_0}(T)$, at $\nu_0 = 0$ and 1, as functions of the temperature T up to 2,000 K. Their accuracy is $\sim 2\%$, considering the slow convergence of the resonances, and the rates are stable with respect to changes of the parameters in **Tables 1, 2**. Both rates increase quickly by a factor of ~ 30 from 20 to 200 K, with a maximum at ~ 250 K equal to $1.75 \cdot 10^{-14}$ and $1.97 \cdot 10^{-12} \text{ cm}^3 \text{ sec}^{-1}$ for $\nu_0 = 0$ and 1, respectively. This behavior is associated with the sharp ICS resonances above ~ 0.006 eV and is similar to that of $\text{H}+\text{HeH}^+$ (De Fazio, 2014). Note that a rate-constant maximum was also found in the adiabatic ground PES dynamics of $\text{H}+\text{HeH}^+$ (Esposito et al., 2015) using the PES of

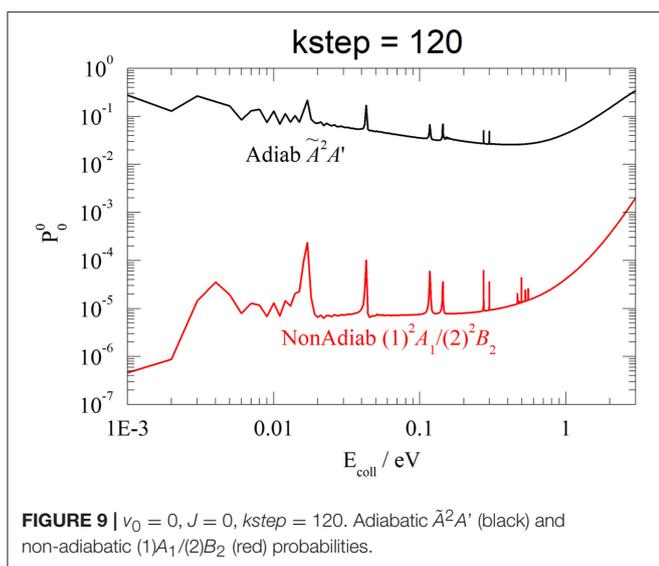
TABLE 4 | Rate constants k_0 and $k_1/\text{cm}^3 \text{ s}^{-1}$.

T/K	$k_0/10^{-14}$	$k_1/10^{-12}$
50	0.73	1.12
100	1.36	1.71
200	1.72	1.99
300	1.74	1.98
500	1.61	1.81
1,000	1.37	1.53
2,000	1.30	1.39

**FIGURE 8** | Initial-state-resolved rate constants at $\nu_0 = 0$ (red) and 1 (blue).

Ramachandran et al. (2009), but it was at $\sim 10,000$ K and due to a complete different mechanism. Then the rates slowly decrease with T and become nearly constant above 1,500 K, where they differ by two orders of magnitude. On the overall, this behavior is due to the ICS resonances shown in **Figure 7** overlapped to a nearly Langevin decrease of the associated cross sections up to ~ 0.5 eV.

Because we have considered just the ground rotational state of $\text{H}_2(X^1\Sigma_g^+)$, $j_0 = 0$, the present rate constant at $\nu_0 = 0$ and 300 K underestimates by a factor of ~ 6 the thermal experimental $k(330) = (1.1 \pm 0.1) \times 10^{-13} \text{ cm}^3 \text{ s}^{-1}$ (Johnsen et al., 1980), which is however 39 years old. The agreement increases if we consider H_2 in the excited rotational states that are open at 300 K, because test calculations suggest that the reactivity can be enhanced by $\sim 33\%$ when $j_0 = 1$. Taking into account the $\text{H}_2(X^1\Sigma_g^+)$ Fermi-Dirac nuclear spin statistics, the room-temperature populations of $j_0 = 0$ and 1 are ~ 0.13 and 0.66 , respectively, and this implies a rotational enhancement of the rate constant by a factor of ~ 7 . Owing to the room-temperature branching ratio of *para*- and *ortho*- H_2 and the rotational effects on the reactivity, we can roughly estimate that *ortho*- H_2 is ~ 4 times more reactive than the *para* species.



In closing this section, we have also done some test calculations of the reaction probabilities in the adiabatic approximation, on the \tilde{A}^2A' PES of HeH_2^+ (see **Figure 1**). Without reaching the accuracy and the stability of the results in section Reaction Probabilities, we present an example at $v_0 = 0, J = 0$, and $kstep = 120$, contrasting in **Figure 9** adiabatic \tilde{A}^2A' and non-adiabatic CI $(1)^2A_1/(2)^2B_2$ results. Although both probabilities present a resonance structure at low E_{coll} and increase above ~ 1 eV, the reactivity is dramatically different, with the \tilde{A}^2A' probability hugely larger than the CI one, from 6 orders of magnitude at 0.001 eV to 2 orders at 3 eV. This finding shows that the \tilde{A}^2A' adiabatic PES drives the WP into the exothermic product channel, following a C_s pathway that avoids the C_{2v} barrier of the CI. On the contrary, the non-adiabatic CI WP remains essentially on the initial and repulsive $(1)^2A_1$ diabatic PES V_{11} , owing to the weak non-adiabatic interaction. As expected, only at energies larger than 3 eV the two probabilities seem to merge and the adiabatic approximation is probably less worse. We roughly estimate that the adiabatic \tilde{A}^2A' rate constants at 330 K is $\sim 1,000$ times larger than the experimental value (Johnsen et al., 1980).

CONCLUSIONS

In this article we have presented quantum non-adiabatic DCT results of the $\text{He}^+ + \text{H}_2$ collision, employing an electronic diabatic representation, previously computed by Aguado et al. (1993), a WP time dependent formalism, and a flux analysis. Specifically, we have taken into account the non-adiabatic CI coupling between the first two excited diabatic PESs of HeH_2^+ . Dynamical calculations are performed for the ground and first excited vibrational states of H_2 , for investigating vibrational effects on the DCT dynamics, and collision energies up to 3 eV are considered. Reaction probabilities and ICSs present strong and narrow resonance features up to 0.5 eV, due to quasi-bound molecular states embedded in the continuum and trapped in the ion-induced dipole minimum of the reactant

channel, near the CI. These features are very hard to converge and affect all the DCT dynamics. These intense resonance features have a determinant role in the radiative charge-transfer formation (Mrugala et al., 2013) of stable HeH_2^+ , probably present in the interstellar medium (Tennyson and Miller, 1987). At higher collision energies the computational load reduces drastically owing to a near conservation of the helicity quantum number. Our results are in reasonable agreement with previous experimental and theoretical studies of this reaction, confirming the strong vibrational enhancement previously founded. In fact, rate constants increase of about two orders of magnitude just adding one vibrational quantum to the H_2 reactants.

At the best of our knowledge, these are the first rigorous quantum DCT calculations presented for a chemical system, made possible by the joint implementation of time-dependent WP, time-to-energy Fourier transform, and flux methods. On the other hand, a more rigorous approach, as time independent close coupling calculations, cannot be attempted at the present state-of-the-art of the reaction dynamics theories, because of the difficulty of these methods to treat the three-bodies breakup (see e.g., (Pack et al., 1998), and references therein). The results achieved could have relevant consequences in astrophysics and in particular in the early Universe evolution models (Bovino et al., 2011). Although the reaction is probably negligible when the molecular hydrogen is relaxed in its ground vibrational state, the strong increase with the vibrational energy suggests that it could have a role during the early stages of the adiabatic expansion at high redshift, destroying the molecular hydrogen formed and slowing down further cooling of the primordial gas. Of course, to determinate its role, evolution Universe models with at least a vibrational resolution of the chemical network (Coppola et al., 2011, 2012) are required. Moreover, this reaction could be also relevant in other different astrophysical environments, as for example in the study of the Sun atmosphere where the temperature is very high and hydrogen and helium are very abundant species.

AUTHOR CONTRIBUTIONS

All authors listed have made a substantial, direct and intellectual contribution to the work, and approved it for publication.

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