



Scrutinizing GW-Based Methods Using the Hubbard Dimer

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Using the simple (symmetric) Hubbard dimer, we analyze some important features of the GW approximation. We show that the problem of the existence of multiple quasiparticle solutions in the (perturbative) one-shot GW method and its partially self-consistent version is solved by full self-consistency. We also analyze the neutral excitation spectrum using the Bethe-Salpeter equation (BSE) formalism within the standard GW approximation and find, in particular, that 1) some neutral excitation energies become complex when the electron-electron interaction U increases, which can be traced back to the approximate nature of the GW quasiparticle energies; 2) the BSE formalism yields accurate correlation energies over a wide range of U when the trace (or plasmon) formula is employed; 3) the trace formula is sensitive to the occurrence of complex excitation energies (especially singlet), while the expression obtained from the adiabatic-connection fluctuation-dissipation theorem (ACFDT) is more stable (yet less accurate); 4) the trace formula has the correct behavior for weak (*i.e.*, small U) interaction, unlike the ACFDT expression.

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1 INTRODUCTION

Many-body perturbation theory (MBPT) based on Green's functions is among the standard tools in condensed matter physics for the study of ground- and excited-state properties. (Aryasetiawan and Gunnarsson, 1998; Onida et al., 2002; Martin et al., 2016; Golze et al., 2019). In particular, the GW approximation (Hedin, 1965; Golze et al., 2019) has become the method of choice for bandstructure and photoemission calculations and, combined with the Bethe-Salpeter equation (BSE@ GW) formalism, (Salpeter and Bethe, 1951; Strinati, 1988; Albrecht et al., 1998; Rohlfing and Louie, 1998; Benedict et al., 1998; van der Horst et al., 1999a; Blase et al., 2018, 2020), for optical spectra calculations. Thanks to efficient implementations, (Duchemin and Blase, 2019, 2020, 2021; Bruneval et al., 2016; van Setten et al., 2013; Kaplan et al., 2015, 2016; Krause and Klopper, 2017; Caruso et al., 2012, 2013b,a; Caruso, 2013; Wilhelm et al., 2018), this toolkit is acquiring increasing popularity in the traditional quantum chemistry community, (Rohlfing and Louie, 1999; van der Horst et al., 1999b; Puschnig and Ambrosch-Draxl, 2002; Tiago et al., 2003; Boulanger et al., 2014; Jacquemin et al., 2015b; Bruneval et al., 2015; Jacquemin et al., 2015a; Hirose et al., 2015; Jacquemin et al., 2017a,b; Rangel et al., 2017; Krause and Klopper, 2017; Gui et al., 2018; Blase et al., 2018; Liu et al., 2020; Blase et al., 2020; Holzer and Klopper, 2018; Holzer et al., 2018; Loos et al., 2020), partially due to the similarity of the equation structure to that of the standard Hartree-Fock (HF) (Szabo and Ostlund, 1989) or Kohn-Sham (KS) (Hohenberg and Kohn, 1964; Kohn and Sham, 1965) mean-field methods. Several studies of the performance of various flavors of GW in atomic and molecular systems are now present in the literature, (Holm

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and von Barth, 1998; Stan et al., 2006; Stan et al., 2009; Blase and Attaccalite, 2011; Faber et al., 2011; Bruneval, 2012; Bruneval and Marques, 2013; Bruneval et al., 2015; Karlsson and van Leeuwen, 2016; Bruneval et al., 2016; Bruneval, 2016; Boulanger et al., 2014; Blase et al., 2016; Li et al., 2017; Hung et al., 2016, 2017; van Setten et al., 2015, 2018; Ou and Subotnik, 2016, 2018; Faber, 2014), providing a clearer picture of the pros and cons of this approach. There are, however, still some open issues, such as 1) how to overcome the problem of multiple quasiparticle solutions, (van Setten et al., 2015; Maggio et al., 2017; Loos et al., 2018; Véril et al., 2018; Duchemin and Blase, 2020; Loos et al., 2020), 2) what is the best way to calculate ground-state total energies, (Casida, 2005; Huix-Rotllant et al., 2011; Caruso et al., 2013b; Casida and Huix-Rotllant, 2016; Colonna et al., 2014; Olsen and Thygesen, 2014; Hellgren et al., 2015; Holzer et al., 2018; Li et al., 2019, 2020; Loos et al., 2020), and 3) what are the limits of the BSE in the simplification commonly used in the so-called Casida equations. (Strinati, 1988; Rohlfing and Louie, 2000; Sottile et al., 2003; Myöhänen et al., 2008; Ma et al., 2009a,b; Romaniello et al., 2009b; Sangalli et al., 2011; Huix-Rotllant et al., 2011; Sakkinen et al., 2012; Zhang et al., 2013; Rebolini and Toulouse, 2016; Olevano et al., 2019; Lettmann and Rohlfing, 2019; Loos and Blase, 2020; Authier and Loos, 2020; Monino and Loos, 2021). In the present work, we address precisely these questions by using a very simple and exactly solvable model, the symmetric Hubbard dimer. Small Hubbard clusters are widely used test systems for the GW approximation (e.g. Verdozzi et al., 1995; Schindlmayr et al., 1998; Pollehn et al., 1998; Puig von Friesen et al., 2010; Romaniello et al., 2009a, 2012). Despite its simplicity, the Hubbard dimer is able to capture lots of the underlying physics observed in more realistic systems, (Romaniello et al., 2009a, 2012; Carrascal et al., 2015, 2018), such as, for example, the nature of the band-gap opening in strongly correlated systems as bulk NiO. (Di Sabatino et al., 2016). Here, we will use it to better understand some features of the GW approximation and the BSE@GW approach. Of course, care must be taken when extrapolating conclusions to realistic systems.

The paper is organized as follows. **Section 2** provides the key equations employed in MBPT to calculate removal and addition energies (or charged excitations), neutral (or optical) excitation energies, and ground-state correlation energies. In **Sec. 3**, we present and discuss the results that we have obtained for the Hubbard dimer. We finally draw conclusions and perspectives in **Sec. 4**.

2 THEORETICAL FRAMEWORK

In the following we provide the key equations of MBPT (Martin et al., 2016) and, in particular, we discuss how one can calculate ground- and excited-state properties, namely removal and addition energies, spectral function, total energies, and neutral excitation energies. We use atomic units $\hbar = m = e = 1$ and work at zero temperature throughout the paper.

2.1 The GW Approximation

Within MBPT a prominent role is played by the one-body Green's function G which has the following spectral representation in the frequency domain:

$$G(\mathbf{x}_1, \mathbf{x}_2; \omega) = \sum_{\nu} \frac{\psi_{\nu}(\mathbf{x}_1)\psi_{\nu}^{*}(\mathbf{x}_2)}{\omega - \epsilon_{\nu} + i\eta \operatorname{sgn}(\epsilon_{\nu} - \mu)},$$
(1)

where μ is the chemical potential, η is a positive infinitesimal, $\epsilon_{\nu} = E_{\nu}^{N+1} - E_{0}^{N}$ for $\epsilon_{\nu} > \mu$, and $\epsilon_{\nu} = E_{0}^{N} - E_{i}^{N-1}$ for $\epsilon_{\nu} < \mu$. Here, E_{ν}^{N} is the total energy of the ν th excited state of the *N*-electron system ($\nu = 0$ being the ground state). In the case of single-determinant many-body wave functions (such as HF or KS), the so-called Lehmann amplitudes ψ_{ν} (\mathbf{x}) reduce to one-body orbital energies.

The one-body Green's function is a powerful quantity that contains a wealth of information about the physical system. In particular, as readily seen from Eq. 1, it has poles at the charged excitation energies of the system, which are proper addition/ removal energies of the *N*-electron system. Thus, one can also access the (photoemission) fundamental gap

$$E_{\rm g} = I^N - A^N, \tag{2}$$

where $I^N = E_0^{N-1} - E_0^N$ is the ionization potential and $A^N = E_0^N - E_0^{N+1}$ is the electron affinity. Moreover, one can straightforwardly obtain the spectral function, which is closely related to photoemission spectra, as

$$A(\mathbf{x}_1, \mathbf{x}_2; \omega) = \frac{1}{\pi} \operatorname{sgn}(\mu - \omega) \operatorname{Im} G(\mathbf{x}_1, \mathbf{x}_2; \omega).$$
(3)

The ground-state total energy can also be extracted from *G* using the Galitskii-Migdal (GM) formula (Galitskii and Migdal, 1958)

$$E_0^{\rm GM} = -\frac{\mathrm{i}}{2} \int d\boldsymbol{x}_1 \lim_{2 \to 1^+} \left[\mathrm{i} \frac{\partial}{\partial t_1} + h(\boldsymbol{r}_1) \right] G(1,2), \tag{4}$$

where $1 \equiv (\mathbf{x}_1, t_1)$ is a space-spin plus time composite variable and $h(\mathbf{r}) = -\nabla/2 + v_{\text{ext}}(\mathbf{r})$ is the one-body Hamiltonian, $v_{\text{ext}}(\mathbf{r})$ being the local external potential.

The one-body Green's function can be obtained by solving a Dyson equation of the form $G = G_0 + G_0\Sigma G$, where G_0 is the non-interacting Green's function and the self-energy Σ is an effective potential which contains all the many-body effects of the system under study. In practice, Σ must be approximated and a well-known approximation is the so-called GW approximation in which the self-energy reads $\Sigma^{GW} = v_H + iGW$, where v_H is the classical Hartree potential, and $W = \varepsilon^{-1}v_c$ is the dynamically screened Coulomb interaction, with ε^{-1} the inverse dielectric function and v_c the bare Coulomb interaction. (Hedin, 1965).

The equations stemming from the *GW* approximation should, in principle, be solved self-consistently, since Σ is a functional of *G*. (Hedin, 1965). Self-consistency, however, is computationally demanding, and one often performs a single *GW* correction (for example using G_0 as starting point one builds *W* and Σ^{GW} as $\Sigma^{GW} = v_{\rm H} + iG_0W_0$, with $v_{\rm H} = -iv_cG_0$ and $W_0 = [1 + iv_cG_0G_0]^{-1}v_c$, from

which $G = \{1 - G_0 \Sigma^{GW} [G_0]\}^{-1} G_0$). This cost-saving and popular strategy is known as one-shot *GW*. The main drawback of the one-shot *GW* method is its dependence on the starting point (*i.e.*, the orbitals and energies of the HF or KS mean-field eigenstates) originating from its perturbative nature. To overcome this problem, one can introduce some level of self-consistency. Removal/addition energies are thus obtained by solving iteratively the so-called quasiparticle equation

$$\omega = \epsilon_i^{\rm HF} + \langle \phi_i^{\rm HF} | \Sigma_{\rm c}^{\rm GW} (\omega) | \phi_i^{\rm HF} \rangle.$$
(5)

Here, we choose to start from HF spatial orbitals $\phi_i^{\text{HF}}(\mathbf{r})$ and energies ϵ_i^{HF} , which are corrected by the (real part of the) correlation contribution of the *GW* self-energy $\Sigma_c^{GW} = \Sigma^{GW} - \Sigma_{\text{HF}}$, where $\Sigma_{\text{HF}} = v_H + iv_c G$ is the HF (hartree plus exchange) contribution to the self-energy. Σ_c^{GW} is evaluated with G_{HF} at the first iteration, where G_{HF} is the self-consistent solution of $G_{\text{HF}} = G_0 + G_0 \Sigma^{\text{HF}} G_{\text{HF}}$. At the *n*th iteration, Σ_c^{GW} is evaluated as $\Sigma_c^{GW}[G^{n-1}]$, where G^{n-1} has poles at the energies from the (n-1)-th iteration of **Eq. 5** and corresponding weights obtained from the *Z* factors given in **Eq. 6**. As a non-linear equation, **Eq. 5** has potentially many solutions $\epsilon_{i,v}^{GW}$. The so-called quasiparticle (QP) solution $\epsilon_{i,v=0}^{GW} \equiv \epsilon_i^{QP}$ has the largest renormalization factor (or spectral intensity)

$$Z_{i,\nu} = \left[1 - \langle \phi_i^{\rm HF} | \frac{\partial \Sigma_c^{\rm GW}(\omega)}{\partial \omega} | \phi_i^{\rm HF} \rangle \middle|_{\omega = \epsilon_{i,\nu}^{\rm GW}}\right]^{-1}, \tag{6}$$

while the satellite (sat) peaks $\epsilon_{i,\nu>0}^{GW} \equiv \epsilon_{i,\nu}^{sat}$ share the remaining of the spectral weight. Moreover, one can show that the following sum rule is fulfilled (von Barth and Holm, 1996)

$$\sum_{\nu} Z_{i,\nu} = 1,$$
 (7)

where the sum runs over all the solutions of the quasiparticle equation for a given mean-field eigenstate *i*. Throughout this article, *i*, *j*, *k*, and *l* denote general spatial orbitals, *a* and *b* refer to occupied orbitals, *r* and *s* to unoccupied orbitals, while *m* labels single excitations $a \rightarrow r$.

In eigenvalue self-consistent GW (commonly abbreviated as ev*GW*), (Hybertsen and Louie, 1986; Shishkin and Kresse, 2007; Blase and Attaccalite, 2011; Faber et al., 2011; Rangel et al., 2016; Gui et al., 2018), one only updates the poles of *G*, while keeping fix the orbitals (or weights). *G* is then used to build Σ^{GW} and *W*. At the *n*th iteration, the removal/addition energies are obtained from the *GW* quasiparticle solutions computed from $G_{n-1}W$ (G_{n-1}) where the satellites are discarded at each iteration. Nonetheless, at the final iteration one can keep the satellite energies to get the full spectral function (**Eq. 3**). In fully self-consistent *GW* (sc*GW*), (Caruso et al., 2012, 2013b,a; Caruso, 2013; Koval et al., 2014), one updates the poles and weights of *G* retaining quasiparticle and satellite energies at each iteration.

It is instructive to mention that, for a conserving approximation, the sum of the intensities corresponding to removal energies equals the number of electrons, *i.e.*, $\sum_{\epsilon_{i,v}^{GW} < \mu} Z_{i,v} = N$. scGW is an example of conserving approximations, while, in general, the one-shot GW does not conserve the number of electrons.

2.2.1 Neutral Excitations

Linear response theory (Oddershede and Jorgensen, 1977; Casida, 1995; Petersilka et al., 1996) in MBPT is described by the Bethe-Salpeter equation. (Strinati, 1988). The standard BSE within the static GW approximation (referred to as BSE@ GW in this work, which means the use of GW quasiparticle energies to build the independent-particle excitation energies and of the GW self-energy to build the static exchangecorrelation kernel) can be recast, assuming a closed-shell reference state, as a non-Hermitian eigenvalue problem known as Casida equations:

$$\begin{pmatrix} \boldsymbol{A}^{\lambda} & \boldsymbol{B}^{\lambda} \\ -\boldsymbol{B}^{\lambda} & -\boldsymbol{A}^{\lambda} \end{pmatrix} \begin{pmatrix} \boldsymbol{X}_{m}^{\lambda} \\ \boldsymbol{Y}_{m}^{\lambda} \end{pmatrix} = \Omega_{m}^{\lambda} \begin{pmatrix} \boldsymbol{X}_{m}^{\lambda} \\ \boldsymbol{Y}_{m}^{\lambda} \end{pmatrix},$$
(8)

where Ω_m^{λ} is the *m*th excitation energy with eigenvector $(X_m^{\lambda} Y_m^{\lambda})^{\top}$ at interaction strength λ , $^{\top}$ is the matrix transpose, and we have assumed real-valued spatial orbitals. The non-interacting and physical systems correspond to $\lambda = 0$ and 1, respectively. The matrices A^{λ} and B^{λ} are of size $OV \times OV$, where O and V are the number of occupied and virtual orbitals, respectively, and O+V is the total number of spatial orbitals. Introducing the so-called Mulliken notation for the bare two-electron integrals

$$(ij|kl) = \iint d\mathbf{r}_1 d\mathbf{r}_2 \phi_i(\mathbf{r}_1) \phi_j(\mathbf{r}_1) v_c(\mathbf{r}_1 - \mathbf{r}_2) \phi_k(\mathbf{r}_2) \phi_l(\mathbf{r}_2), \quad (9)$$

and the corresponding (static) screened Coulomb potential matrix elements

$$W_{ij,kl}(\omega = 0) = \iint d\mathbf{r}_1 d\mathbf{r}_2 \phi_i(\mathbf{r}_1) \phi_j(\mathbf{r}_1) W(\mathbf{r}_1, \mathbf{r}_2; \omega = 0) \phi_k(\mathbf{r}_2) \phi_l(\mathbf{r}_2),$$
(10)

the BSE matrix elements read (Maggio and Kresse, 2016).

$$A_{ar,bs}^{\lambda,\sigma\sigma'} = \delta_{ab}\delta_{rs}\left(\epsilon_{r}^{\rm QP} - \epsilon_{a}^{\rm QP}\right) + \lambda \left[\alpha_{\sigma\sigma'}\left(ar|sb\right) - W_{ab,sr}\left(\omega = 0\right)\right],$$
(11a)

$$B_{ar,bs}^{\lambda,\sigma\sigma'} = \lambda \big[\alpha_{\sigma\sigma'} \left(ar | bs \right) - W_{as,br} \left(\omega = 0 \right) \big], \tag{11b}$$

where ϵ_i^{QP} are the *GW* quasiparticle energies, and $\alpha_{\uparrow\downarrow} = 2$ and $\alpha_{\uparrow\uparrow} = 0$ for singlet (*i.e.*, spin-conserved) and triplet (*i.e.*, spin-flip) excitations, respectively.

In the absence of instabilities (*i.e.*, when $A^{\lambda} - B^{\lambda}$ is positivedefinite), (Dreuw and Head-Gordon, 2005), **Eq. 8** is usually transformed into an Hermitian eigenvalue problem of half the dimension

$$(\boldsymbol{A}^{\lambda} - \boldsymbol{B}^{\lambda})^{1/2} (\boldsymbol{A}^{\lambda} + \boldsymbol{B}^{\lambda}) (\boldsymbol{A}^{\lambda} - \boldsymbol{B}^{\lambda})^{1/2} \boldsymbol{V}_{m}^{\lambda} = (\Omega_{m}^{\lambda})^{2} \boldsymbol{V}_{m}^{\lambda}, \qquad (12)$$

where the excitation amplitudes are

$$(\boldsymbol{X}^{\lambda} + \boldsymbol{Y}^{\lambda})_{m} = (\Omega_{m}^{\lambda})^{-1/2} (\boldsymbol{A}^{\lambda} - \boldsymbol{B}^{\lambda})^{+1/2} \boldsymbol{V}_{m}^{\lambda}, \qquad (13a)$$

$$(\boldsymbol{X}^{\lambda} - \boldsymbol{Y}^{\lambda})_{m} = (\Omega_{m}^{\lambda})^{+1/2} (\boldsymbol{A}^{\lambda} - \boldsymbol{B}^{\lambda})^{-1/2} \boldsymbol{V}_{m}^{\lambda}.$$
(13b)

Singlet $(\Omega_m^{\uparrow\downarrow} \equiv \Omega_m^{\lambda=1,\uparrow\downarrow})$ and triplet $(\Omega_m^{\uparrow\uparrow} \equiv \Omega_m^{\lambda=1,\uparrow\uparrow})$ excitation energies are obtained by diagonalizing **Eq. 8** at $\lambda = 1$.

2.2.2 Correlation Energies

Our goal here is to compare the BSE correlation energy E_c^{BSE} obtained using two formulas, namely the trace (or plasmon) formula (Rowe, 1968; Ring and Schuck, 1980) and the expression obtained using the adiabatic-connection fluctuation-dissipation theorem (ACFDT) formalism. (Furche and Van Voorhis, 2005; Toulouse et al., 2009, 2010; Hellgren and von Barth, 2010; Angyan et al., 2011; Heßelmann and Görling, 2011; Colonna et al., 2014; Maggio and Kresse, 2016; Holzer et al., 2018; Loos et al., 2020). The two approaches have been recently compared at the randomphase approximation (RPA) level for the case of Be2, (Li et al., 2020), showing similar improved performances at the RPA@GW@PBE level with respect to the RPA@PBE level and an impressive accuracy by introducing BSE (BSE@GW@HF) correction in the trace formula. Here we would like to get more insights into the quality of these two approaches.

The ground-state correlation energy within the trace formula is calculated as

$$E_{c}^{\mathrm{Tr}@\mathrm{BSE}} = E_{c,\uparrow\downarrow}^{\mathrm{Tr}@\mathrm{BSE}} + E_{c,\uparrow\uparrow}^{\mathrm{Tr}@\mathrm{BSE}} + \frac{1}{2} \left[\sum_{m} \Omega_{m}^{\uparrow\downarrow} - \mathrm{Tr}(\mathbf{A}^{\uparrow\downarrow}) \right] + \frac{1}{2} \left[\sum_{m} \Omega_{m}^{\uparrow\uparrow} - \mathrm{Tr}(\mathbf{A}^{\uparrow\uparrow}) \right],$$
(14)

where $A^{\sigma\sigma'} \equiv A^{\lambda=1,\sigma\sigma'}$ is defined in **Eq.11a** and Tr denotes the matrix trace. We note that the trace formula is an approximate expression of the correlation energy since it relies on the so-called quasi-boson approximation and on the killing condition on the zeroth-order Slater determinant ground state (Li et al., 2020 for more details). Note that here both sums in **Eq. 14** run over all resonant (hence real- and complex-valued) excitation energies while they are usually restricted to the real-valued resonant BSE excitation energies. Thus, the Tr@BSE correlation energy is potentially a complex-valued function in the presence of singlet and/or triplet instabilities.

The ACFDT formalism, (Furche and Van Voorhis, 2005), instead, provides an in-principle exact expression for the correlation energy within time-dependent density-functional theory (TDDFT). (Runge and Gross, 1984; Petersilka et al., 1996; Ullrich, 2012). In practice, however, one always ends up with an approximate expression, which quality relies on the approximations to the exchange-correlation potential of the KS system and to the kernel of the TDDFT linear response equations. In this work, therefore, we use the ACFDT expression within the BSE formalism and we explore how well it performs and how it compares to the trace **Eq. 14**.

Within the ACFDT framework, only the singlet states do contribute for a closed-shell ground state, and the ground-state BSE correlation energy

$$E_{\rm c}^{\rm AC@BSE} = \frac{1}{2} \int_{0}^{1} d\lambda {\rm Tr} \left(\boldsymbol{K}^{\uparrow\downarrow} \boldsymbol{P}^{\lambda,\uparrow\downarrow} \right)$$
(15)

is obtained via integration along the adiabatic connection path from the non-interacting system at $\lambda = 0$ to the physical system $\lambda = 1$, where

$$\boldsymbol{K} = \begin{pmatrix} \tilde{\boldsymbol{A}}^{\lambda=1} & \boldsymbol{B}^{\lambda=1} \\ \boldsymbol{B}^{\lambda=1} & \tilde{\boldsymbol{A}}^{\lambda=1} \end{pmatrix}$$
(16)

is the interaction kernel, (Angyan et al., 2011; Holzer et al., 2018; Loos et al., 2020) $\tilde{A}_{ar,bs}^{\lambda,\sigma\sigma'} = \alpha_{\sigma\sigma'}\lambda(ar|sb)$, and

$$\boldsymbol{P}^{\lambda} = \begin{pmatrix} \boldsymbol{Y}^{\lambda} (\boldsymbol{Y}^{\lambda})^{\top} & \boldsymbol{Y}^{\lambda} (\boldsymbol{X}^{\lambda})^{\top} \\ \boldsymbol{X}^{\lambda} (\boldsymbol{Y}^{\lambda})^{\top} & \boldsymbol{X}^{\lambda} (\boldsymbol{X}^{\lambda})^{\top} \end{pmatrix} - \begin{pmatrix} 0 & 0 \\ 0 & 1 \end{pmatrix}$$
(17)

is the correlation part of the two-body density matrix at interaction strength λ . Here again, the AC@BSE correlation energy might become complex-valued in the presence of singlet instabilities.

Note that the trace and ACFDT formulas yield, for any set of eigenstates, the same correlation energy at the RPA level. (Angyan et al., 2011). Moreover, in contrast to density-functional theory where the electron density is fixed along the adiabatic path, (Langreth and Perdew, 1979; Gunnarsson and Lundqvist, 1976; Zhang and Burke, 2004), at the BSE@*GW* level, the density is not maintained as λ varies. Therefore, an additional contribution to **Eq. 15** originating from the variation of the Green's function along the adiabatic connection should, in principle, be added. However, as commonly done within RPA (Toulouse et al., 2009, 2010; Angyan et al., 2011; Colonna et al., 2014) and BSE, (Holzer et al., 2018; Loos et al., 2020), we neglect this additional contribution.

3 RESULTS

As discussed in **Sec. 1**, in this work, we consider the (symmetric) Hubbard dimer as test case, which is governed by the following Hamiltonian

$$\hat{H} = -t \sum_{\sigma=\uparrow,\downarrow} \left(\hat{c}_{1\sigma}^{\dagger} \hat{c}_{2\sigma} + \hat{c}_{2\sigma}^{\dagger} \hat{c}_{1\sigma} \right) + U \left(\hat{n}_{1\uparrow} \hat{n}_{1\downarrow} + \hat{n}_{2\uparrow} \hat{n}_{2\downarrow} \right).$$
(18)

Here $\hat{n}_{1\sigma} = \hat{c}^{\dagger}_{1\sigma}\hat{c}_{1\sigma}$ ($\hat{n}_{2\sigma} = \hat{c}^{\dagger}_{2\sigma}\hat{c}_{2\sigma}$) is the spin density operator on site 1 (site 2), $\hat{c}^{\dagger}_{1\sigma}$ and $\hat{c}_{1\sigma}$ ($\hat{c}^{\dagger}_{2\sigma}$ and $\hat{c}_{2\sigma}$) are the creation and annihilation operators for an electron at site 1 (site 2) with spin σ , U is the on-site (spin-independent) interaction, and -t is the hopping kinetic energy. The physics of the Hubbard model arises from the competition between the hopping term, which prefers to delocalize electrons, and the on-site interaction, which favors localization. The ratio U/t is a measure for the relative contribution of both terms and is the intrinsic, dimensionless coupling constant of the Hubbard model, which we use in the following. In this work we consider the dimer at one-half filling.

3.1 Quasiparticle Energies in the *GW* Approximation

We test different flavors of self-consistency in *GW* calculations: one-shot *GW*, ev*GW*, partial self-consistency through the alignment of the chemical potential (psc*GW*), where we shift G_0 or $G_{\rm HF}$ in such a way that the resulting *G* has the same chemical potential than the shifted G_0 or shifted $G_{\rm HF}$, (Schindlmayr, 1997), and sc*GW*. In the one-shot formalism, we also test two different starting points: the truly noninteracting Green's function G_0 (U = 0) and the HF Green's function G_{HF}. These two schemes are respectively labeled as $G_0 W_0$ and $G_{\rm HF} W_{\rm HF}$ in the following.

The $G_0 W_0$ self-energy (in the site basis) and removal/ addition energies are already given in Ref. (Romaniello et al., 2012) for the Hubbard dimer at one-half filling. For completeness we report them in Supplementary Appendix S1, together with the renormalization factors, which are discussed in Sec. 3.1.1.

Starting from $G_{\rm HF}$, which reads

$$G_{\mathrm{HF},II}(\omega) = \frac{1}{2} \left[\frac{(-1)^{(I-J)}}{\omega - (t+U/2) + \mathrm{i}\eta} + \frac{1}{\omega + (t-U/2) - \mathrm{i}\eta} \right], \quad (19)$$

where I and J run over the sites, the (correlation part of the) $G_{\rm HF}W_{\rm HF}$ self-energy is $\Sigma_{c,II}^{GW}(\omega) = \Sigma_{II}^{GW}(\omega) - \delta_{IJ}U/2$ with

$$\Sigma_{c,IJ}^{GW}(\omega) = \frac{U^2 t}{2h} \left[\frac{1}{\omega - (t+h+U/2) + i\eta} + \frac{(-1)^{I-J}}{\omega + (t+h-U/2) - i\eta} \right],$$
(20)

where $h = \sqrt{4t^2 + 4Ut}$. Here we used the following expression for the polarizability P = -iGG with elements

$$P_{IJ}(\omega) = \frac{(-1)^{I-J}}{4} \left[\frac{1}{\omega - 2t + i\eta} - \frac{1}{\omega + 2t - i\eta} \right]$$
(21)

to build the screened interaction $W = v_c + v_c PW$, whose only nonzero matrix elements read

$$W_{II,JJ}(\omega) = U\delta_{IJ} + (-1)^{I-J} \frac{U^2 t}{h} \left[\frac{1}{\omega - h + i\eta} - \frac{1}{\omega + h - i\eta} \right]$$
(22)

due to the local nature of the electron-electron interaction. The quantities defined in Eqs 19-22 can then be transformed to the bonding (bn) and antibonding (an) basis (which is used to recast the BSE as Eq. 8) thanks to the following expressions:

$$|\mathrm{bn}\rangle = \frac{|1\rangle + |2\rangle}{\sqrt{2}}, |\mathrm{an}\rangle = \frac{|1\rangle - |2\rangle}{\sqrt{2}}.$$
 (23)

Therefore, the one-shot removal/addition energies read

$$\epsilon_{1,\pm} = +\frac{h}{2} + \frac{U}{2} \pm \frac{\sqrt{(h+2t)^2 + 4tU^2/h}}{2},$$
 (24a)

$$\epsilon_{2,\pm} = -\frac{h}{2} + \frac{U}{2} \pm \frac{\sqrt{(h+2t)^2 + 4tU^2/h}}{2},$$
 (24b)

with the quasiparticle solutions being $\epsilon_{bn}^{QP} = \epsilon_{1,-}$ and $\epsilon_{an}^{QP} = \epsilon_{2,+}$, which correspond to the bonding and antibonding energies, respectively. As readily seen in Eqs 24a, 24b, in addition to the quasiparticle, there is a unique satellite per eigenstate given by $\epsilon_{bn}^{sat} = \epsilon_{1,+}$ and $\epsilon_{an}^{sat} = \epsilon_{2,-}$. Moreover, the closed-form expression of the renormalization factors (Eq. 6) reads

$$Z_{\rm bn/an}^{\rm QP} = \frac{t \left[h^2 + 2ht + 2U^2 + h\sqrt{(h+2t)^2 + 4tU^2/h} \right]}{h^3 + 4h^2t + 4ht^2 + 4tU^2 - h^2\sqrt{(h+2t)^2 + 4tU^2/h}}$$
(25)

and $Z_{bn/an}^{sat} = 1 - Z_{bn/an}^{QP}$. The ev*GW* and sc*GW* calculations were performed numerically using the meromorphic representation of G, following Ref. (Puig von Friesen et al., 2010) with some slight modifications (Supplementary Appendix S2 for more details). At each iteration, the solution of the Dyson equations for G and W (Sec. 2.1) produces extra poles. In order to keep the number of poles under control in scGW, the poles with intensities smaller than a user-defined threshold (set from 10^{-4} to 10^{-6} depending on the ratio U/t) are discarded and the corresponding spectral weight is redistributed among the remaining poles.

In Figure 1, we present the spectral function of G (Eq. 3) for different values of the ratio U/t (U/t = 1, 5, 10, and 15) and using $G_{\rm HF}$ as starting point. We consider three GW variants: $G_{\rm HF}W_{\rm HF}$ ev*GW*, and sc*GW*. For $U/t \leq 3$, all the schemes considered here provide a faithful description of the quasiparticle energies. For larger U/t, GW (regardless of the level of self-consistency) tends to underestimate the fundamental gap E_{g} (Eq. 2), as shown in the upper left panel of Figure 2. $G_{\rm HF}W_{\rm HF}$ and evGW give a very similar estimate of $E_{\rm g}$, whereas the quasiparticle intensity $Z_{\rm bn/an}^{\rm QP}$ defined in Eq. 25 is quite different and overestimated by both methods, at least in the range of U/t considered in Figure 2 (center left panel).

The main effects of full self-consistency are the reduction of E_{σ} (see upper left panel of Figure 2), and the creation of extra satellites with decreasing intensity (see upper panel of Figure 1). For small U/t, the fundamental gap is similar to the one predicted by other methods while for increasing U/t the agreement worsen and E_{g} is grossly underestimated. The quasiparticle intensity is very similar to the one predicted by $G_{\rm HF}W_{\rm HF}$. Concerning the position of the satellites, we observe that the one-shot $G_{\rm HF}W_{\rm HF}$ scheme gives the most promising results. Numerical values of quasiparticle and first satellite energies as well as their respective intensities in the spectral functions presented in Figure 1 are gathered in Table 1.

We notice that a similar analysis for H_2 in a minimal basis has been presented in Ref. (Hellgren et al., 2015) with analogous conclusions.

For the sake of completeness, we also report in the bottom left panel of Figure 2 the total energy calculated using the Galitskii-Migdal formula (Eq. 4). Since the Galitskii-Migdal total energy is not stationary with respect to changes in G, one gets meaningful energies only at self-consistency. However, for the Hubbard dimer, we do not observe a significant impact of selfconsistency, as one can see from Figure 1 by comparing the total energy at the $G_{\rm HF}W_{\rm HF}$, evGW, and scGW levels. For each of these schemes which correspond to a different level of selfconsistency, the Galitskii-Migdal formula provides accurate total energies only for relatively small $U/t (\leq 3)$.

If we consider $G_{\rm HF}$ as starting point and we define the chemical potential as $\mu = (\epsilon_{an}^{QP} + \epsilon_{bn}^{QP})/2$, then the alignment of the chemical potential has no effect on the spectrum, this means that $G_{\rm HF}W_{\rm HF}$ and pscGW are equivalent.

3.1.1 G₀: A Bad Starting Point

In the following we will illustrate how the starting point can influence the resulting quasiparticle energies. The Green's







function obtained from the one-shot $G_0 W_0$ does not satisfy particle-hole symmetry, the fundamental gap is underestimated (top right panel of Figure 2) yet more accurate than $G_{\rm HF}W_{\rm HF}$ (top left panel of Figure 2), the quasiparticle intensity relative to the bonding component is close to the exact result up to $U/t \approx 16$ (center right panel of Figure 2), while overestimated for the antibonding components. Moreover, we note that the intensities of the two poles of the bonding component crosses at U/t = 24. This means that if we sort the quasiparticle and the satellite according to their intensity at a given U/t, the nature of the two poles is interchanged when one increases U/t, which results in a discontinuity in the QP

energy. Meanwhile, the total number of particle is not conserved (N < 2). For G_0W_0 we found a small deviation from N = 2 for small U/t (e.g. N = 1.98828 at U = 1), which becomes larger by increasing the interaction (e.g. N = 1.55485 for U/t = 10). Instead, starting from $G_{\rm HF}$ the particle number is always conserved. We checked that for the self-consistent calculations the total particle number is conserved, as it should.

Considering G_0 as starting point in evGW, we encounter the problem described in Ref. (Véril et al., 2018), namely the discontinuity of various key properties (such as the fundamental gap in the top right panel of **Figure 2**) with respect to the interaction strength U/t. This issue is solved, for

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the Hubbard dimer, by considering a better starting point or using the fully self-consistent scheme scGW. Note, however, that improving the starting point does not always cure the discontinuity problem as this issue stems from the quasiparticle approximation itself. Full self-consistency, instead, avoids systematically discontinuities since no distinction is made between quasiparticle and satellites. Unfortunately, full self-consistency is much more involved from a computational point of view and, moreover, it does not give an overall improvement of the various properties of interest, at least for the Hubbard dimer, for which $G_{\rm HF}W_{\rm HF}$ is to be preferred. For more realistic (molecular) systems, it was shown in Ref. (Berger et al., 2020). that the computationally cheaper selfconsistent COHSEX scheme solves the problem of multiple quasiparticle solutions.

3.2 Bethe-Salpeter Equation

For the Hubbard dimer the matrices A^{λ} and B^{λ} in Eq. (8) are just single matrix elements and they simply read, for both spin manifolds,

$$A^{\lambda,\uparrow\downarrow} = \Delta \epsilon^{GW} + \lambda \frac{U}{2}, \ B^{\lambda,\uparrow\downarrow} = \lambda \frac{U}{2} \left(\frac{4tU}{h^2} + 1 \right),$$
(26a)

$$A^{\lambda\uparrow\uparrow} = \Delta \epsilon^{GW} - \lambda \frac{U}{2}, \ B^{\lambda\uparrow\uparrow} = \lambda \frac{U}{2} \left(\frac{4tU}{h^2} - 1\right), \tag{26b}$$

while $\tilde{A}^{\lambda,\uparrow\downarrow} = \lambda U$. We employ the screened Coulomb potential given in **Eq. 22** at $\omega = 0$ for the kernel, and the *GW* quasiparticle energies from **Eqs 24a** and **24b** to build the *GW* approximation of the fundamental gap $\Delta \epsilon^{GW} = \epsilon_{an}^{QP} - \epsilon_{bn}^{QP}$. For comparison purposes, we also use the *exact* quasiparticle energies [see Eq. (C3) of Ref. (Romaniello et al., 2012).], which consists in replacing $\Delta \epsilon^{GW}$ by the *exact* fundamental gap $E_g = \sqrt{16t^2 + U^2} - 2t$. In such a case, one is able to specifically test how accurate the BSE formalism is at catching the excitonic effect via the introduction of the screened Coulomb potential.

We notice that, within the so-called Tamm-Dancoff approximation (TDA) where one neglects the coupling matrix B^{λ} between the resonant and anti-resonant parts of the BSE Hamiltonian (Eq. 8), BSE yields RPA with exchange (RPAx) excitation energies for the Hubbard dimer. This is the case also for approximations to the BSE kernel which are beyond GW, such as the T-matrix approximation. (Romaniello et al., 2012; Zhang et al., 2017; Li et al., 2021), and it is again related to the local nature of the electron-electron interaction. Hence, to test the effect of approximations on correlation for this model system we must go beyond the TDA.

3.2.1 Neutral Excitations

In **Figure 3**, we report the real part of the singlet and triplet excitation energies obtained from the solution of **Eq. 8** for $\lambda = 1$. For comparison, we report also the exact excitation energies obtained as differences of the excited- and ground-state total energies of the Hubbard dimer obtained by diagonalizing the Hamiltonian (18) in the Slater determinant basis $\{|1\uparrow, 1\downarrow, |1\uparrow, 2\downarrow, |1\downarrow, 2\uparrow, |2\uparrow, 2\downarrow\}$ built from the sites [Ref. (Romaniello et al., 2009a) for the exact total energies]. For the

Ut		ean ean	۵			- <u></u> .,	ϵ_{an}			Z an an	<u>م</u> -			\mathbf{Z}_{an}^{sat}	at n	
	Exact	G _{HF} W _{HF}	evGW	scGW	Exact	G _{HF} W _{HF}	evGW	scGW	Exact	G _{HF} W _{HF}	evGW	scGW	Exact	G _{HF} W _{HF}	evGW	scGW
	1.0615	1.0721	1.0702	1.0651	3.0615	3.9006	4.1175	4.0793	0.9851	0.9855	0.9864	0.9861	0.0149	0.0145	0.0135	0.0132
	2.2016	1.6739	1.6302	1.4334	4.2016	6.5728	8.8364	7.6389	0.8123	0.9183	0.9398	0.9239	0.1876	0.0817	0.0602	0.0593
	4.3852	2.4893	2.4001	1.7787	6.3852	9.1225	14.7136	10.8296	0.6857	0.8717	0.9182	0.8777	0.3143	0.1282	0.0818	0.0823
	6.7621	3.2887	3.1813	2.0542	8.7621	11.2887	20.5769	13.3847	0.6288	0.8430	0.9082	0.8472	0.3712	0.1570	0.0918	0.0934



singlet manifold, this yields, for the single excitation $\Omega_1^{\uparrow\downarrow}$ and double excitation $\Omega_2^{\uparrow\downarrow}$, the following expressions:

$$\Omega_1^{\uparrow\downarrow} = \frac{1}{2} \left(U + \sqrt{16t^2 + U^2} \right), \Omega_2^{\uparrow\downarrow} = \sqrt{16t^2 + U^2}, \qquad (27)$$

while the unique triplet transition energy is

$$\Omega_1^{\uparrow\uparrow} = \frac{1}{2} \Big(-U + \sqrt{16t^2 + U^2} \, \Big). \tag{28}$$

Of course, one cannot access the double excitation within the static approximation of BSE, (Strinati, 1988; Romaniello et al., 2009b; Loos and Blase, 2020), so only the lowest singlet and triplet excitations, $\Omega_1^{\uparrow\downarrow}$ and $\Omega_1^{\uparrow\uparrow}$, are studied below.

Using one-shot $G_{\rm HF}W_{\rm HF}$ quasiparticle energies (BSE@ $G_{\rm HF}W_{\rm HF}$) produces complex excitation energies (see right panel of Figure 3). We find the same scenario also with other flavors of GW (not reported in the figure), such as scGW. The occurrence of complex poles and singlet/triplet instabilities at the BSE level are well documented (Holzer et al., 2018; Blase et al., 2020; Loos et al., 2020) and is not specific to the Hubbard dimer. For example, one finds complex poles also for H₂ along its dissociation path, (Li and Olevano, 2021), but also for larger diatomic molecules. (Loos et al., 2020). For U/t > 12.4794, the singlet energy becomes pure imaginary, the same is observed for the triplet energy for 7.3524 < U/t < 12.4794. These two points corresponds to discontinuities in the first derivative of the excitation energies with respect to U/t (Figure 3). The BSE excitation energies are good approximations to their exact analogs only for $U/t \leq 2$ for the singlet and $U/t \leq 6$ for the triplet. Using exact quasiparticle energies instead produces real excitation energies, with the singlet energy in very good agreement with the exact result; the triplet energy, instead, largely overestimates the exact value. This seems to suggest that complex poles are caused by the approximate nature of the GW quasiparticle energies, although, of course, the quality of the kernel also plays a role. Indeed, setting W = 0 but using GWQP energies, BSE yields real-valued excitation energies. It would be interesting to further investigate this issue by using the exact kernel together with GW QP energies. This is left for future work.

3.2.2 Correlation Energy

For the Hubbard dimer, we have $E^{\text{HF}} = -2t + U/2$, and the correlation energy given in Eq. 15 can be calculated analytically. After a lengthy but simple derivation, one gets

$$\begin{split} E_{c}^{AC \oplus \text{BSE}} &= -\frac{U}{2} + \frac{t^{2} - 2U^{2}}{2U(2t + 3U)} \bigg\{ \Delta \epsilon^{GW} - \\ &\frac{1}{2(t + U)} \sqrt{[-U^{2} + 2(t + U)\Delta \epsilon^{GW}] [U(2t + 3U) + 2(t + U)\Delta \epsilon^{GW}]} \bigg\} \\ &- \frac{t + 2U}{2\sqrt{U(2t + 3U)}} \bigg(\frac{3t + 4U}{2t + 3U} + \frac{t}{U} \bigg) \Delta \epsilon^{GW} \\ &\text{atan} \left\{ - \frac{U\sqrt{U(2t + 3U)}}{2\Delta \epsilon^{GW} (t + U) + \sqrt{[-U^{2} + 2(t + U)\Delta \epsilon^{GW}] [U(2t + 3U) + 2(t + U)\Delta \epsilon^{GW}]} \right\} \end{split}$$

Results are reported in Figure 4 and are compared with the exact correlation energy (Romaniello et al., 2009a)

$$E_{\rm c} = -\frac{\sqrt{16t^2 + U^2}}{2} + 2t. \tag{29}$$

The AC@BSE correlation energy does not possess the correct asymptotic behavior for small U, as Taylor expanding Eq. 29 for small U, we obtain

$$E_{c}^{AC@BSE} = -\frac{U^{2}}{32t} - \frac{5U^{3}}{96t^{2}} + \frac{323U^{4}}{6144t^{3}} + \mathcal{O}(U^{4}),$$
(30)

while the exact correlation energy behaves as

$$E_{\rm c} = -\frac{U^2}{16t} + \frac{U^4}{1024t^3} + \mathcal{O}(U^6).$$
(31)

Moreover, we found that the radius of convergence of the small-U/t expansion of $E_c^{AC@BSE}$ is very small due to a square-root branch point for $U/t \approx -2/3$.

In the case of the trace formula **Eq. 14**, the singlet and triplet contributions behave as

$$E_{c,\uparrow\downarrow}^{\mathrm{Tr}\oplus\mathrm{BSE}} = -\frac{U^2}{32t} - \frac{7U^3}{128t^2} + \frac{99U^4}{2048t^3} + \mathcal{O}(U^5), \qquad (32a)$$

$$E_{c,\uparrow\uparrow}^{\mathrm{Tr@BSE}} = -\frac{U^2}{32t} + \frac{7U^3}{128t^2} - \frac{157U^4}{2048t^3} + \mathcal{O}(U^5), \qquad (32b)$$



derivative of the energy (corresponding to the appearance of complex poles) are indicated by open circles.

which guarantees the correct asymptotic behavior for the total Tr@BSE correlation energy

$$E_{\rm c}^{\rm Tr@BSE} = -\frac{U^2}{16t} - \frac{29U^4}{1024t^3} + \mathcal{O}(U^5), \tag{33}$$

and cancels the cubic term (as it should).

The trace formula is strongly affected by the appearance of the imaginary excitation energies: as shown in Figure 4 where we plot the real and complex components of the BSE@ $G_{HF}W_{HF}$ correlation energy as functions of U/t at various levels of theory, irregularities (*i.e.*, discontinuities in the first derivative of the energy) appear at the values of U/t for which the triplet and singlet energies become purely imaginary. The ACFDT expression, instead, is more stable over the range of U/t considered here with only a small cusp on the energy surface at the singlet instability point after which the real part of $E_c^{AC@BSE}$ behaves linearly with respect to U/t. Overall, however, the correlation energy obtained by the trace formula is almost on top of its exact counterpart over a wide range of U/t, with a rather small contribution from the triplet component, *i.e.*, $|E_{c,\uparrow\uparrow}^{\text{Tr@BSE}}| \ll |E_{c,\uparrow\uparrow}^{\text{Tr@BSE}}|$. For comparison purposes, the RPA correlation energy, which is obtained from the trace or ACDFT formula using $BSE@G_{HF}W_{HF}$ with W = 0 in the BSE kernel, is also reported in Figure 4. Both formulas yield the same correlation energies as expected, and they show no irregularities thanks to the fact that BSE excitation energies are real-valued at the RPA level. Also correlation energies obtained using BSE@exact (also shown in Figure 4) do not show irregularities for the same reason. Moreover, they show a visible upshift with respect to the corresponding AC@BSE@G_{HF}W_{HF} and Tr@BSE@ $G_{\rm HF}W_{\rm HF}$ results, which worsens the agreement with the exact correlation energy. Finally, we observe that both expressions for the correlation energy (at BSE@GW level) produce better results than the Galitskii-Migdal Eq. 4, as one can see from Figure 4, in particular at large U/t.

4 CONCLUSION

In this work we have used the symmetric Hubbard dimer to better understand some features of the *GW* approximation and of BSE@ *GW*. In particular, we have found that the unphysical discontinuities that may occur in quasiparticle energies computed using one-shot or partially self-consistent *GW* schemes disappear using full self-consistency. However, full self-consistency does not give an overall improvement in term of accuracy and, at least for the Hubbard dimer, $G_{\rm HF}W_{\rm HF}$ is to be preferred.

We have also analyzed the performance of the BSE@*GW* approach for neutral excitations and correlation energies. We have found that, at any level of self-consistency, the excitation energies become complex for some critical values of *U/t*. This seems related to the approximate nature of the *GW* quasiparticle energies, since using exact quasiparticle energies (hence the exact fundamental gap) solves this issue. The BSE excitation energies are good approximations to the exact analogs only for a small range of *U/t* (or *U/t* \leq 2 for the lowest singlet-singlet transition and *U/t* \leq 6 for the singlet-triplet transition), while the strong-correlation regime remains a challenge.

The correlation energy obtained from these excitation energies using the trace (or plasmon) formula has been found to be in very good agreement with the exact results over the whole range of U/tfor which these energies are real. The occurrence of complex singlet and triplet excitation energies shows up as irregularities in the correlation energy. The ACFDT formula, instead, is less sensitive to this. However, we have found that the AC@BSE correlation energy is less accurate than the one obtained using the trace formula. Both, however, perform better than the standard Galitskii-Migdal formula. Finally, we have studied the small-U expansion of the correlation energy obtained with the trace and ACFDT formulas and we found that the former, contrary to the latter, has the correct behavior when one includes both the singlet and triplet energy contributions. Our findings point out to a possible fundamental problem of the AC@BSE formalism.

Although our study is restricted to the half-filled Hubbard dimer, some of our findings are transferable to realistic (molecular) systems. In particular: 1) a fully self-consistent solution of the *GW* equation cures the problem of multiple QP solutions, avoiding in the process the appearance of discontinuities in key physical quantities such as total or excitation energies, ionization potentials, and electron affinities; 2) a "bad" starting point (G_0 in the case of the Hubbard dimer) may result in the appearence of multiple QP solutions; 3) potential energy surfaces computed with the trace formula and within the ACFDT formalism may exhibit irregularities due to the appearence of complex BSE excitation energies; 4) for the Hubbard dimer at half-filling, the trace formula has the correct asymptotic behavior (thanks to the inclusion of singlet and triplet excitation energies) for weak interaction, contrary to its ACFDT counterpart. It would be interesting to check if it is also the case in realistic systems.

DATA AVAILABILITY STATEMENT

The original contributions presented in the study are included in the article/**Supplementary Material**, further inquiries can be directed to the corresponding author.

AUTHOR CONTRIBUTIONS

All authors listed have made a substantial, direct, and intellectual contribution to the work and approved it for publication.

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SUPPLEMENTARY MATERIAL

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