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Cassiterite U-Pb dating and micro-XRF analysis constraint on the formation of Xinlu Sn-Zn deposit, South China

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This study presents rock slab and thin section reflection mapping, element mapping by Micro-XRF (µ-XRF) analysis, and in situ cassiterite U-Pb laser ablation-inductively coupled plasma mass spectrometer (LA-ICP-MS) ages of the Xinlu Sn-Zn deposit, Nanling Sn-W belt, South China. The new results provided constraints on the age and processes of Sn mineralization and thus further discussed the role of the Paleo-Pacific plate in the formation of Sn-W mineralization in the Nanling Sn-W belt. Based on mineralogy and µ-XRF analysis, four mineralization episodes are distinguished in the Baimianshan segment, Xinlu deposit, including 1) prograde skarn stage characterized by the mineral assemblages of garnet, pyroxene, marble, hornfels; 2) retrograde skarn-sulfide composed of massive pyrrhotite, sphalerite with parts of epidote, cassiterite; 3) quartz sulfide vein stage consisting of arsenopyrite, pyrite, quartz, and 4) barren calcite crosscutting the former episodes. The time of retrograde skarn-sulfide mineralization is dated by LA-ICP-MS cassiterite U-Pb at 168.7 \pm 1.3 Ma. These U-Pb ages, combined with previous geochronological studies, support one Sn mineralization event associated with the 159.5-168.7 Ma magmatism in the Guposhan-Huashan district. The Xinlu Sn-Zn mineralization and the Middle-Late Jurassic Sn-W mineralization in the Nanling belt, South China is genetically related to asthenospheric upwelling and crust-mantle interaction caused by the subduction of the Paleo-Pacific plate. Our results demonstrate that thin section reflection and µ-XRF element mapping are powerful tools for determining paragenesis in ore deposits and are highly effective to guide the selection of analytical positions for cassiterite U-Pb dating. This method may also be applicable to in-situ geochronology of other minerals.

KEYWORDS

micro-XRF, Xinlu deposit, Nanling Sn-W belt, South China, Paleo-Pacific, cassiterite U-Pb geochronology

Introduction

The Nanling Sn-W belt is a world-class Sn and W province within the South China Block (SCB) (Figures 1A,B; Mao et al., 2021b). Most tungsten and tin deposits are related to the late Jurassic felsic rocks (Xiao et al., 2019; Mao et al., 2021a). The spatial distribution of these rocks displays a southeastward younging trend (Figure 1C), which was interpreted to be related to the flat-slab subduction of the Paleo-Pacific Ocean plate (Figure 1D; Li and Li, 2007; Yang et al., 2020). In contrast, the Sn-W mineralization in this region does not display a similar age pattern. The ore-forming ages of most of large-scale W and Sn deposits in the Nanling Sn-W belt concentrated in 160–150 Ma. Therefore, whether the Paleo-Pacific plate

contributed to the formation of Sn-W mineralization is equivocal.

Cassiterite U-Pb dating has proved to be a robust method to precisely and directly date the tin mineralization of W-Sn deposits (Yuan et al., 2008; Zhang et al., 2015; Xiao et al., 2022). Most published cassiterite U-Pb dating required mineral separation and grain mounts. A major drawback of these two methods is that they may mix polygenetic populations of cassiterite. Hence, the rapid identification of cassiterite on the hand specimen or thin section before the cassiterite U-Pb dating is quite essential. Micro-XRF analyses are a powerful and non-destructive technique to determine the trace elements in the geological samples, and to identify the minerals and their contents (Flude et al., 2017; Wang et al., 2022).



Above all, the μ -XRF analyses have great potential to guide the selection of analytical spots for cassiterite U-Pb dating.

The Guposhan-Huashan Sn-W district is one of the most important districts in the western of the Nanling Sn-W belt (Figure 1C). The Sn-W deposits in this district (e.g., Xinlu Sn-Zn deposit, Shuiyanba Sn deposit, Keda Sn deposit, Huashan Sn deposit, and Shanhu W-Sn deposit) are located at the contact zone of the Guposhan-Huashan granite pluton and its vicinity (Figures 1C, 2A; Feng et al., 2019; Gu et al., 2007a, b; Kang et al., 2012; Li et al., 2016; Xiao et al., 2011; Zhang et al., 2020; Zhao et al., 2010). Considerable uncertainty and controversy remains regarding the exact timing of W-Sn mineralization. For example, some studies revealed that no direct genetic link between Sn-W mineralization in the Xinlu deposit and the Guposhan pluton based on whole-rock K-Ar ages and fluid inclusion Rb-Sr ages of the ores (Dong, 1991; Zhu et al., 2006; Gu et al., 2007a). Many researchers have demonstrated that W-Sn mineralization was related to the Guposhan pluton based on zircon U-Pb, molybdenite Re-Os, and muscovite/biotite Ar-Ar dating results in the Xinlu deposit (Kang et al., 2012; Li et al., 2016; Feng et al., 2019). Besides, molybdenite, with Re-Os ages representing the ore-forming ages, is not a major mineral in the Xinlu deposit, and most other tested minerals could not be directly dated for ore-forming ages (Feng et al., 2019). Thirdly, the skarn Sn-Zn mineralization with great economic values lacks investigation compared to the abundant studies on the quartz vein type Sn-W mineralization. All these lead to the ambiguous timing of the tin mineralization in the Guposhan-Huashan Sn-W district.

In this contribution, we applied microscopic observation and μ -XRF analysis to determine the paragenetic sequence and cassiterite occurrence in the Xinlu deposit. We performed laser ablation-inductively coupled plasma-mass spectrometry (LA-ICP-MS) U-Pb dating of cassiterite to directly date tin mineralization. From our new cassiterite U-Pb age, and previously published cassiterite U-Pb ages, we refined the spatial and temporal distribution of Sn-W mineralization and discussed the role of the Paleo-Pacific plate in the formation of Sn-W mineralization in the Nanling Sn-W belt.

Geological setting

Regional geology

The South China Block is composed of the Yangtze Block in the northwest and the Cathaysia Block in the southeast (Figure 1B). These two blocks were welded along the "Shi-Hang zone" during the Early Neoproterozoic, which was a tectonic-granitoid belt (Figure 1B; Gilder et al., 1996). The SCB subsequently underwent the Early Paleozoic NS shortening intraplate tectono-magmatism (Shu et al., 2021), the Triassic intraplate orogeny (Li et al., 2022), and the Late Mesozoic tectonic transformation (Shu, 2012). Large-scale Sn-W mineralization events occurred during the Mesozoic in the eastern SCB. Numerous studies have demonstrated that the subduction of the Paleo-Pacific plate was responsible for producing large-scale magmatism and Sn-W mineralization (Li and Li, 2007; Mao et al., 2021a, b; Zhou and Li, 2000).

The Nanling Sn-W belt locates in the central south of the SCB. In this belt, there are intensive folded and deformed basement, magmatic rocks, and rifting basins (Zhang et al., 2015). The basement is composed of a metamorphosed Neoproterozoic-Ordovician flysch-volcanic series, covered by unmetamorphosed Late Devonian to Early Triassic sedimentary rocks. Five episodic magmatism events (Neoproterozoic, Silurian, Permian, Triassic, Jurassic, and Cretaceous) occurred in this belt (Zhang et al., 2015). Previous studies indicate that W and Sn mineralization occurred in the belt in the Neoproterozoic and Mesozoic (Figure 1C; Hu and Zhou, 2012), of which the latter is economically more important.

This study focuses on the Guposhan-Huashan W-Sn district in the western of the Nanling Sn-W belt, which is located in the eastern of the Guangxi Zhuang Autonomous Region (GZAR). The study area mainly consists of semipelagic flysch sedimentary rocks and subsequently underwent the Early Paleozoic intraplate orogeny. Scattered Mo-W, Pb-Zn, and Au mineralization associated with this orogeny are represented by the Baishiding, Zhanggongling, and Longshui deposits in the eastern of the district (Figure 2A; Li et al., 2016). From Early Devonian to Jurassic, carbonate platforms with sporadic clastic rocks formed in this area under a shore-shallow sea-slope setting (Shu, 2012). The Jurassic clastic rocks, formed in a graben basin, unconformably overly the Paleozoic rocks (Figure 2A). Largescale magmatism and associated Sn-W mineralization occurred in response to the subduction of the Pacific plate in the Middle-Late Jurassic and Cretaceous (Figure 2A; Feng et al., 2019; Gu et al., 2007a; Gu et al., 2007b; Kang et al., 2012; Li et al., 2016; Xiao et al., 2011; Zhang et al., 2020; Zhao et al., 2010). The Quaternary system hosts alluvial tin, which is widely distributed in the south of the district. Folds and faults are widely developed in this district.

Ore deposit geology

The Xinlu Sn-Zn deposit is located at the south contact zone of the Guposhan granite pluton and its vicinity. It consists of the Shihuichong Sn-Fe, Xinlu Sn-Zn, Liuheao Sn, Muqiaomian Sn-Zn, Baimianshan Sn-Zn, and Shimen Sn-Zn segments from north to south (Figure 2B). The main sedimentary rocks in the deposit are composed of middle Devonian Xindu Formation (D_2x) and Yujiang Formation (D_2y) clastic rocks, Donggangling Formation (D_2d) carbonate rocks, upper Devonian Guilin (D_3g) and Rongxian (D_3r) formations, and



(A) Geological map of Guposhan-Huashan Sn-W district (modified after Geological Prospecting Institute of Guangxi (GPIG), 2021). (B) Geological map of Xinlu Sn-Zn deposit (modified after Feng et al., 2019). (C) Geological map of Baimianshan Zn-Sn segment (modified after GPIG, 2021).

Lower Carboniferous carbonate rocks. The NNW-, NW-, NE-, and EW-striking faults are widely developed in the deposit. The NNW-striking faults control the distribution of major ore segments in the deposit (Figure 2B). The dextral EW-striking faults crosscut the NNW-, and NW-striking faults. The NNWstriking faults plunge to the directions of 260°–290° at dip angles of 58°–85°. Most of the NE-striking faults plunge in the directions of 330°–340° at angles of 72°–85°. Fault-parallel folds are developed in the southeastern of the deposit. The Guposhan granite is distributed in the northern half of the deposit. Minor intermediate-felsic dykes are distributed along or crosscut the NE-, and NW-striking faults in the southern part of the deposit. Minor skarn and fault-controlled oxidized orebodies crop out on the surface in the deposit (Figure 2C).

The Baimianshan Sn-Zn segment is controlled by the NNWstriking normal faults (F_{23}) and is located in the middle of the Xinlu deposit (Figures 2B,C). The carbonate rocks of the Donggangling Formation and clastic rocks of the Xindu Formation are distributed on the east and west sides of the segments, respectively. The Devonian system developed in the west limb of the Kejiashan anticline and the stratum dipped SWW at 12° - 30° (Hu et al., 1992). The Quaternary sediments



unconformably overly the Devonian in the middle of the segment. The NS-striking F_{23} and F_4 normal faults vertical cross throughout the segment. Previous studies revealed that these NS-striking faults underwent nearly E–W compression following an E–W extension (Ou, 1989). The main NE-trending concealed orebody was formed between these two faults. Parts of oxidized orebodies were formed along the NE-, and small NS-striking faults in the eastern of the segment. Most of the NE-striking faults are normal faults except for the F_{55} fault. A few of NS-, and NE-striking dykes spread in the south of the segment. The alteration is not developed on the surface, and it includes silicification, potassic alteration, and sericitization. Some small skarn is found in the F_{23} fault. The orebodies are mainly buried underground with insignificant oxidized ore

Sampling and analytical methods

The hand specimens in this study were collected from drill cores ZK0201 and ZK0801 from two of the new holes of the Baimianshan segment in 2019 (Figures 3, 4). The two drill

cores have ferrous-sphalerite-cassiterite sulfide orebodies at the lithological contacts in the upper of the Xindu Formation (D_2x) with a small number of sulfide orebodies in the Donggangling Formation (D_2d) limestone. Two types of ores including XLM1901 and XLM1902 were sampled. Sample XLM1901 is massive ferrous sphalerite-cassiteritearsenopyrite sulfide that was formed in the Donggangling Formation limestone at 49.64 m of drill hole ZK0201. The Sn and Zn ore grades of XLM1901 are 0.31% and 8.65%, respectively. Sample XLM1902 is a layer-parallel skarn cassiterite sulfide orebody that was formed in the upper part of the Xindu Formation at 133 m of drill hole ZK0801 with 2.46% Sn and 2.04% Zn. The ore was formed in the contacts between pellet marlstone and argillaceous siltstone. We carried out observation 10 hand specimens, micro-observation, reflection mapping, and µ-XRF analysis on 2 representative polished slabs and 1 polished thin section, and cathodoluminescence (CL) imaging of targeted cassiterite grains, and LA-ICPto obtain detailed paragenesis MS dating and cassiterite U-Pb ages of the Baimianshan segment, Xinlu deposit.



FIGURE 4

Photographs of representative ore and rocks within the Baimianshan Zn-Sn segment. (A) Core from drill hole ZK0801. (B) Massive garnet and marble were crosscut by steep-dipping calcite veins. (C) Steep-dipping branching veins or veinlets and layer-parallel massive ores overly on the layer-parallel marble and hornfels. (D) Steep-dipping pyrite-quartz-calcite veins crosscut the marble. (E) The arsenopyrite replaced the massive pyrrhotite-chalcopyrite-sphalerite and both were crosscut by the pyrite-quartz-calcite veins. (F) Massive pyrrhotite-chalcopyrite-sphalerite ore.

Thin section reflection mapping and element mapping by $\mu\text{-XRF}$

Sample XLM1901 is a 6.5*6.5 cm² half drill core piece and XLM1902 is a cylindrical rock chip of 6.8 cm in diameter and 1.6 cm in height. The analytical areas included the whole surface of the polished slabs and thin sections for petrographic observations and element mapping. Polished sections and slabs were examined using transmitted and reflected optical microscopes (DM4500P, Leica, Germany). Reflection mapping of the whole polished slabs on the microscope was completed by manual splicing through Image Analysis System software (Figures 5A,B). The µ-XRF analyses on rock slabs and thin sections were performed on the benchtop M4 Tornado (Bruker, Germany) under vacuum conditions (20 mbar, 50 kV, 200 $\mu A)$ with a diameter of 20 μm X-ray beam with 40 µm pixel size. Element mapping scanning and sample navigation on the benchtop M4 were run by a motorized stage that moves the sample beneath the static X-ray beam. Data were collected for Fe, Zn, W, Ga, Ge, As, Pb, Bi, Sn, Cu, W, Si, Al, Ca, and S (Figures 6-8). The analytical procedures of element mapping were described in detail by Flude et al. (2017) and Wang et al. (2022). These analyses were conducted at the Laboratory of Dynamic Diagenesis and Metallogenesis, Institute of Geomechanics, Chinese Academy of Geological Science (LDDM-IGM, CAGS). All data acquisition and processing were carried out using the proprietary Bruker software supplied with the instrument.

Cassiterite U-Pb dating and cathodoluminescence imaging

Based on the microscopic observation and μ -XRF analysis, we mixed the two samples to separate cassiterite grains at the LDDM-IGM, CAGS. And then, the cassiterite was mounted on epoxy resin discs. Cassiterite grains were observed under the microscope, and CL and BSE images were obtained to characterize the texture of cassiterite and avoid cracks and mineral/fluid inclusions before U–Pb dating. Selected cassiterite domains without any obvious mineral inclusion were analyzed for U-Pb ages using the Agilent 7900 LA-ICP-MS at the Stake Key Laboratory of Ore Deposit Geochemistry, Institute of Geochemistry, Chinese Academy of Sciences. Cassiterite grains have a core-rim microstructure, and



FIGURE 5

Representative textures and mineral assemblages within ores and wall rocks of Baimianshan Zn-Sn segment. (A) The reflected mapping of polished slab of XLM1902. (B) The reflected mapping of thin section XLM1901T showing the pyrite vein replaced the massive pyrrhotite-sphalerite. (C–D) Cassiterite coexisted with arsenian pyrite and replaced by the pyrrhotite, sphalerite. (E) The pyrrhotite laths and lamellae, chalcopyrite blebs intergrew with sphalerite and pyrrhotite. (F) The euhedral arsenopyrite quartz veinlet crosscut the pyrrhotite and sphalerite. (G) The pyrite quartz replaced the sphalerite and pyrrhotite. (H) The pyrite-calcite vein cut the massive pyrrhotite and sphalerite which were replaced sericite.



FIGURE 6

Photograph and corresponding µ-XRF single-element maps (Fe, S, Si, Sn, Zn, W, Cu, Ga, Ge) of rock slab XLM1901.



thirty spots including 17 fully-dark cores and 13 gray oscillatory rims were analyzed. The reference materials used were AY-4 cassiterite for 206 Pb/ 238 U ratios (ID-TIMS U-Pb age of 158.2 ± 0.4 Ma; Yuan et al., 2008) and NIST SRM 612 for U, Th, and Pb concentrations of cassiterite.

Cassiterite grains were analyzed using LA-ICP-MS with a laser energy density of 5 J/cm², a spot size of 60 μ m, and a repetition rate of 6 Hz. NIST SRM 612 was analyzed twice and AY-4 was analyzed three times for every ten analyses of the unknown cassiterite grains. Each spot analysis incorporated a background signal acquisition of approximately 20 s followed by 35 s of sample data acquisition. Other conditions were listed by Tang et al. (2022). The uncertainty of the single population, ratio uncertainty of the AY-4 reference material, and decay constant uncertainties were propagated to the final results during the process of offline raw data reduction by ICPMSDataCal (Liu et al., 2008). Data errors for isotopic ratios in the cassiterite samples are 1σ . Isoplot 3.0 was used to calculate the U–Pb ages and draw Concordia diagrams (Ludwig, 2003).

Results

Mineralogy and paragenesis

According to the field observation, the Baimianshan segment is similar to most Zn skarn deposits which occur distal from the

causative pluton (Meinert et al., 2005). Most ores on drilling cores are present as massive with minor skarn minerals or hosted by veins and veinlets with skarn alteration halos (Figure 4). Based on microscopic observations, the mineralization process of the Baimianshan segment could be divided into prograde skarn (stage I), retrograde skarn-sulfide (stage II), quartz sulfide veins (stage III), and barren carbonate (stage IV) stages. The prograde stage is characterized by the mineral assemblages of garnet, pyroxene, marble, and hornfels (Figure 4B). The prograde stage is zoned, and the zoning pattern is thick hornfels with parts of pyroxene-bearing pale-brown garnet, banded skarn, marbleization-limestone from the deep to shallow (Figures 3, 4A). The retrograde skarn-sulfide stage is the main mineralization stage. This stage is associated with the formation of steep-dipping branching veins or veinlets and layer-parallel massive ores with dominated ore minerals such as massive pyrrhotite, sphalerite with parts of epidote, cassiterite, arsenian pyrite, and minor chalcopyrite. The euhedral coarse cassiterite (100–700 μ m) is followed by pyrrhotite, and then sphalerite base metal sulfides in this stage (Figures 5C,D). The pyrrhotite laths and lamellae, and chalcopyrite blebs intergrew with sphalerite (Figure 5E). The quartz sulfide veins stage is associated with the formation of middle- and steep-dipping veins with the mineral assemblage of arsenopyrite, pyrite, quartz, sericite, and small amounts of chlorite. This stage could be divided into two sub-stages. Anhedral pyrite, euhedral coarse



Photograph, single-element µ-XRF maps (Fe, S, Sn, Zn, W, Ga, Ge, As, Pb, Bi) and multi-element map of rock slab XLM1902.

arsenopyrite, and fine quartz, sericite of sub-stage I intensively replaced the cassiterite, pyrrhotite, and sphalerite of stage II (Figures 5G,H, 7). The euhedral arsenopyrite-quartz veinlets of sub-stage ii crosscut the veins of sub-stage I (Figure 5F). Steepdipping barren pyrite-calcite veins (stage IV) crosscut the former veins (Figures 4B,D, 5A,H).

Element mapping

Two polished slabs (XLM1901 and XLM1902) and one thin section (XLM1901T) were measured and elemental distribution maps were obtained. Each map was compared to a reflected light image that came from a section close to the one data was collected from (Figures 6–8).

Figure 6 shows the Fe, S, Si, Sn, Zn, W, Cu, Ga, and Ge distributions compared to the hand specimen of XLM1901. Copper, W, Ga, and Ge show an intensive positive correlation

with Zn. Copper could be proved by chalcopyrite blebs coexisting with sphalerite (Figure 5E). Tungsten mineral is not found in this study. Gallium and Ge are thought to be substituted with Zn in sphalerite. Tin distributed in the top-left corner corresponds to the position of cassiterite. The distribution of Fe with help of S could identify different iron-bearing minerals (pyrrhotite and pyrite) and could determine the paragenetic sequence that pyrite replaced pyrrhotite and sphalerite at the left-bottom of the sample.

Figure 7 shows the Fe, S, Zn, W, Cu, Sn, Al, and Si distributions compared to the hand specimen and reflected light image of XLM1901T. The observed depositional sequences are pyrrhotite + sphalerite + epidote followed by quartz + potassium feldspar + arsenopyrite and pyrite + quartz + calcite veinlet. Some disseminated Sn is distributed in the sphalerite aggregation. The distribution characteristics of other elements are similar to those of the XLM1901.



Figure 8 shows single-element distributions of 9 elements (Fe, S, Sn, Zn, W, Pb, Bi, Ga, and Ge), a multi-element map with Fe, Sn, and As, and phases based on multi-element Fe-Zn-As compared to the hand specimen and reflected light image of XLM1902. The distribution characteristics of most elements (Fe, S, Sn, Zn, W, Ga, and Ge) are similar to those of the XLM1901 and XLM1901T. This sample has obvious high counts of As, Pb, and Bi. The high content of As corresponds to the position of arsenopyrite on the reflected light image (Figure 5A), whereas no lead and bismuth minerals were found in this sample. Multi-element maps with Fe, Sn, and As, and phases based on multi-element Fe-Zn-As also clearly reveal the paragenetic sequences of stages II and III.

Cassiterite U-Pb ages

The cassiterite grains are dark brown to black, generally euhedral to subhedral with a size of $100-700 \,\mu$ m. Internal textures of cassiterite are solid dark cores with oscillatory rims. Some cassiterite grains have cracks and are porous with tiny metallic inclusions as revealed by the CL and BSE images (Figure 9A).

The analytical results of U–Pb isotopes of the cassiterite grains are listed in Table 1 and illustrated in Figure 9B. These grains have relatively low total Pb contents of <0.1 ppm, low Th contents of <0.01 ppm, and U contents varying from 0.31 to

2.85 ppm. The analyzed dark cores have higher U concentrations (1.01–2.85 ppm, 1.77 ppm average) than the oscillatory rims (0.31–1.63 ppm, 0.82 ppm average). The plots of 17 cores and 13 rims define a good Concordia line with a lower intercept at 168.7 \pm 1.3 Ma (1 σ ; MSWD = 0.6; n = 30) (Figure 9B).

Discussion

Application of thin section reflection and element mapping

Our results demonstrate the reflection mapping and μ -XRF element mapping on thin section and hand specimens together with detailed petrographic paragenesis in ore deposits can guide the selection of analytical spots for cassiterite U-Pb dating, which would enhance the study of ore formation processes.

Reflection and element mapping of representative samples could provide compositional and mineralogical information across a range of scales. This work is summarized as follows. First, the element mapping could easily distinguish different iron-bearing or sulfur-bearing minerals [such as pyrrhotite (Fe: 63.53%, S: 36.47%), pyrite (Fe: 46.55%, S: 53.45%), arsenopyrite (Fe: 34.3%, S:19.69%), and sphalerite (Fe: <10%, S: 32.9%)] owing to the differences in theoretical composition. The color intensity of Fe-mineral gradually decreases from pyrrhotite, pyrite, arsenopyrite, and sphalerite (Figures 6-8). At the bottom of Figure 6, the light green parts overlap and crosscut the bright green parts, suggesting that the pyrite was formed later than pyrrhotite. In Figure 8, it is obvious that the light green areas overlap and crosscut the bright green and black areas. This is consistent with the microscopic observation that arsenopyrite replaced pyrrhotite and sphalerite. The color intensity of S gradually decreases from pyrite, pyrrhotite, sphalerite, and arsenopyrite (Figures 6-8). The distribution of Fe and S elements can accurately determine the sequence of mineral formation in the Xinlu deposit. Second, element mapping could provide some information on mineralization that cannot be observed under a microscope. The multielements distribution shows that W is positively related to Zn (Figure 8). However, no tungsten mineral (such as sanmartinite) was recognized, implying that the oreforming fluid was rich in tungsten as supported by the quartz vein type tungsten mineralization after the skarn mineralization in the Xinlu deposit (Gu et al., 2007a; Li S. Z. et al., 2012). In the quartz sulfide veins stage, As, Pb, and Bi show a positive correlation and correspond to the position of arsenopyrite (Figures 6-8). Bismuth-lead sulfosalts may be a reasonable explanation. Bismuth-lead sulfosalt is common in the late stage of tin deposits and well documented in the Snsulfide-sulfosalts deposits of the Bolivia Sn province

XLM02 1.01 0.00 0.04 0.01751 0.00867 0.4230 0.0207 0.0302 0.0207 0.00072 0.000 XLM04 1.04 0.00 0.05 0.06571 0.00072 0.2246 0.0179 0.0272 0.00 XLM04 2.25 0.00 0.01 0.0263 0.00019 0.2239 0.01127 0.0269 0.000 XLM11 1.99 0.00 0.02 0.06670 0.0011 0.2248 0.0188 0.0271 0.00 XLM14 1.22 0.00 0.00 0.06670 0.0011 0.2384 0.0188 0.0276 0.001 XLM14 1.29 0.00 0.01 0.06670 0.0011 0.2384 0.0144 0.0264 0.000 XLM17 1.63 0.00 0.01 0.06571 0.0011 0.2384 0.0144 0.0264 0.000 XLM18 1.29 0.00 0.01 0.0552 0.0014 0.0127 0.0128 0.0014 0.0172 0.2172 0.0128 0.021 0.02662 0.0055 0.0216	Analysis spots		U-Th-Pb contents			U-Pb isotopic ratios					
NLM01 1.8 0.00 0.01 0.05703 0.00533 0.2104 0.01273 0.02713 0.00032 XLM02 1.01 0.00 0.05 0.06571 0.00357 0.42390 0.00397 0.0399 0.02349 0.0397 0.0399 0.02239 0.01120 0.02639 0.060 XLM11 1.99 0.00 0.0357 0.00319 0.02349 0.03274 0.000 XLM14 1.82 0.00 0.0567 0.00317 0.2344 0.01120 0.02676 0.000 XLM17 1.63 0.00 0.04 0.06697 0.00317 0.2344 0.0110 0.00576 0.000 0.00576 0.000 0.02776 0.000 0.001 0.00535			²³⁸ U	²³² Th	Common Pb	²⁰⁷ Pb/ ²⁰⁶ Pb	1σ	²⁰⁷ Pb ^{/235U}	1σ	²⁰⁶ Pb/ ²³⁸ U	1σ
XLM02 1.01 0.00 0.04 0.01751 0.00867 0.4230 0.0207 0.0302 0.0207 0.00072 0.000 XLM04 1.04 0.00 0.05 0.06571 0.00072 0.2246 0.0179 0.0272 0.00 XLM04 2.25 0.00 0.01 0.0263 0.00019 0.2239 0.01127 0.0269 0.000 XLM11 1.99 0.00 0.02 0.06670 0.0011 0.2248 0.0188 0.0271 0.00 XLM14 1.22 0.00 0.00 0.06670 0.0011 0.2384 0.0188 0.0276 0.001 XLM14 1.29 0.00 0.01 0.06670 0.0011 0.2384 0.0144 0.0264 0.000 XLM17 1.63 0.00 0.01 0.06571 0.0011 0.2384 0.0144 0.0264 0.000 XLM18 1.29 0.00 0.01 0.0552 0.0014 0.0127 0.0128 0.0014 0.0172 0.2172 0.0128 0.021 0.02662 0.0055 0.0216	XLM fro	om the Baimia	anshan segm	ient, Xinlu S	n-Zn deposit, South Cl	nina					
XLM03 1.40 0.00 0.05 0.06371 0.00375 0.24246 0.0179 0.02792 0.00 XLM04 1.20 0.00 0.01 0.05935 0.00420 0.22082 0.0159 0.02740 0.00 XLM10 2.10 0.00 0.00 0.06203 0.0039 0.2339 0.01127 0.0299 0.00 XLM11 1.91 0.00 0.02 0.05741 0.0013 0.0748 0.01485 0.0229 0.00 XLM14 1.82 0.00 0.05 0.051423 0.00317 0.1918 0.0220 0.0271 0.00 XLM14 1.82 0.00 0.01 0.05570 0.0017 0.23344 0.0118 0.0264 0.00 XLM14 1.82 0.00 0.01 0.05583 0.00130 0.0434 0.0161 0.0266 0.00 XLM20 2.35 0.00 0.01 0.05582 0.0261 0.0194 0.0267 0.00 XLM20 1.93 0.00 0.02 0.05582 0.0014 0.01782 0.0298 0.001<	Core	XLM01	1.86	0.00	0.01	0.05703	0.00353	0.21094	0.01273	0.02713	0.00070
X1M04 1.02 0.00 0.01 0.05935 0.00420 0.22022 0.01500 0.02429 0.00 X1M10 2.50 0.00 0.04 0.05236 0.00251 0.18233 0.00890 0.02439 0.00 X1M11 1.91 0.00 0.02 0.05784 0.00113 0.02748 0.01127 0.02692 0.00 X1M11 1.99 0.00 0.05 0.04123 0.00119 0.01188 0.02270 0.00 X1M14 1.82 0.00 0.03 0.04517 0.00371 0.20845 0.01188 0.02664 0.00 X1M14 1.82 0.00 0.01 0.05570 0.00371 0.23844 0.0120 0.02664 0.00 X1M18 1.29 0.00 0.01 0.05580 0.00310 0.0014 0.02668 0.00 X1M14 1.51 0.00 0.01 0.05582 0.00249 0.0144 0.02687 0.00 0.01 0.05582 0.0025 0.20566 0.0122 0.02680 0.00 0.00 0.0269 0.01 0.01696 <td></td> <td>XLM02</td> <td>1.01</td> <td>0.00</td> <td>0.04</td> <td>0.11751</td> <td>0.00687</td> <td>0.48280</td> <td>0.02807</td> <td>0.03032</td> <td>0.00084</td>		XLM02	1.01	0.00	0.04	0.11751	0.00687	0.48280	0.02807	0.03032	0.00084
X1M09 2.85 0.00 0.04 0.05236 0.00251 0.18233 0.00890 0.0240 0.00 X1M11 1.91 0.00 0.00 0.05734 0.00130 0.27339 0.01127 0.02392 0.00 X1M11 1.91 0.00 0.05 0.05734 0.00131 0.1748 0.0123 0.0274 0.001 X1M12 1.92 0.00 0.05670 0.00317 0.20845 0.01188 0.02762 0.00 X1M17 1.43 0.00 0.04 0.06677 0.00317 0.20845 0.0114 0.0264 0.00 X1M17 1.43 0.00 0.01 0.05530 0.00310 0.01285 0.00256 0.000 X1M19 1.93 0.00 0.01 0.05530 0.00213 0.1985 0.02657 0.001 X1M24 1.51 0.00 0.01 0.05582 0.0041 0.0128 0.0267 0.000 X1M27 1.93 0.00 0.01 0.0139 0.01294 0.0269 0.000 X1M37 0.31 0		XLM03	1.40	0.00	0.05	0.06371	0.00387	0.24246	0.01479	0.02792	0.00079
XLM10 2.10 0.00 0.00 0.06203 0.0309 0.22339 0.01127 0.02639 0.00 XLM11 1.91 0.00 0.02 0.05744 0.00413 0.20748 0.01435 0.01220 0.02711 0.00 XLM113 1.27 0.00 0.00 0.05570 0.00137 0.02945 0.01120 0.02664 0.00 XLM14 1.82 0.00 0.00 0.06577 0.00317 0.2344 0.0120 0.02664 0.00 XLM14 1.82 0.00 0.01 0.05530 0.0011 0.00123 0.0101 0.02663 0.001 XLM19 1.95 0.00 0.01 0.05530 0.0014 0.0124 0.0123 0.02651 0.000 XLM2 1.51 0.00 0.01 0.05538 0.0226 0.20641 0.01228 0.02667 0.001 XLM2 1.51 0.00 0.01 0.05538 0.02626 0.20641 0.01228 0.02690 0.0279		XLM04	1.02	0.00	0.01	0.05935	0.00420	0.22082	0.01569	0.02742	0.00078
XI.M11 1.91 0.00 0.02 0.05784 0.00413 0.20748 0.01435 0.0292 0.00 XI.M13 1.27 0.00 0.05 0.05473 0.00319 0.1918 0.0120 0.02762 0.00 XI.M14 1.82 0.00 0.05 0.06371 0.23344 0.0120 0.02664 0.00 XI.M14 1.82 0.00 0.01 0.05570 0.0031 0.2344 0.0120 0.02664 0.00 XI.M18 1.29 0.00 0.01 0.05570 0.0011 0.02043 0.0161 0.02666 0.00 XI.M12 1.53 0.00 0.01 0.05570 0.0014 0.0216 0.0266 0.00 XI.M23 1.54 0.00 0.01 0.05570 0.0014 0.0103 0.0124 0.4002 0.0770 0.0267 0.000 XI.M25 1.58 0.00 0.02 0.0266 0.0286 0.0280 0.000 0.0104 0.0103 0.0124 0.4005 0.0123 0.0270 0.0270 0.0280 0.000 0.000		XLM09	2.85	0.00	0.04	0.05236	0.00251	0.18823	0.00890	0.02640	0.00068
XLM12 1.99 0.00 0.05 0.05423 0.00319 0.1918 0.01220 0.02711 0.00 XLM14 1.27 0.00 0.00 0.05670 0.00317 0.2045 0.01188 0.02762 0.00 XLM14 1.82 0.00 0.04 0.066677 0.00317 0.2344 0.0120 0.02266 0.00 XLM18 1.29 0.00 0.00 0.06184 0.00389 0.22389 0.01404 0.02664 0.00 XLM19 1.55 0.00 0.01 0.05573 0.00131 0.1918 0.0161 0.02664 0.00 XLM20 2.53 0.00 0.01 0.05572 0.00131 0.10163 0.01284 0.0164 0.02669 0.000 XLM27 1.98 0.00 0.01 0.01588 0.00289 0.20491 0.01044 0.02669 0.000 XLM27 1.98 0.00 0.01 0.01224 0.44026 0.07799 0.02870 0.000 XLM37 0.41 0.00 0.02 0.08688 0.00599 0.21787 <td></td> <td>XLM10</td> <td>2.10</td> <td>0.00</td> <td>0.00</td> <td>0.06203</td> <td>0.00309</td> <td>0.22339</td> <td>0.01127</td> <td>0.02639</td> <td>0.00068</td>		XLM10	2.10	0.00	0.00	0.06203	0.00309	0.22339	0.01127	0.02639	0.00068
XLM13 1.27 0.00 0.03 0.05670 0.00317 0.20845 0.01188 0.02722 0.00 XLM14 1.82 0.00 0.03 0.05670 0.00371 0.2384 0.010997 0.02676 0.00 XLM17 1.63 0.00 0.04 0.06513 0.00371 0.2384 0.0144 0.02664 0.00 XLM18 1.25 0.00 0.01 0.05530 0.0031 0.2043 0.0161 0.02667 0.00 XLM24 1.51 0.00 0.01 0.05588 0.0034 0.21861 0.00968 0.0267 0.00 XLM25 1.98 0.00 0.01 0.05662 0.00256 0.20861 0.00968 0.0269 0.00 XLM05 0.44 0.00 0.01 0.01294 0.44026 0.0779 0.0270 0.026 XLM06 0.47 0.00 0.03 0.07552 0.00702 0.26956 0.0212 0.02808 0.000 0.000 0.000		XLM11	1.91	0.00	0.02	0.05784	0.00413	0.20748	0.01435	0.02692	0.00075
XLM14 1.82 0.00 0.03 0.05473 0.00289 0.19620 0.00997 0.02676 0.00 XLM17 1.63 0.00 0.04 0.06507 0.00371 0.23344 0.0120 0.02644 0.00 XLM18 1.25 0.00 0.01 0.05330 0.02013 0.10404 0.02661 0.00 XLM22 2.53 0.00 0.01 0.05530 0.00289 0.21762 0.01228 0.0263 0.00 XLM24 1.51 0.00 0.01 0.05588 0.00289 0.20491 0.0104 0.02667 0.00 XLM27 1.93 0.00 0.02 0.05662 0.00256 0.20851 0.00086 0.02699 0.0120 XLM05 0.44 0.00 0.01 0.10130 0.01294 0.44026 0.07709 0.02570 0.000 XLM06 0.47 0.00 0.03 0.05772 0.02696 0.02123 0.0000 0.051 0.02773 0.02780 0.		XLM12	1.99	0.00	0.05	0.05423	0.00319	0.19918	0.01220	0.02711	0.00081
XLM17 1.63 0.00 0.04 0.06507 0.00371 0.23344 0.01210 0.02664 0.00 XLM18 1.29 0.00 0.01 0.05530 0.00310 0.20343 0.0161 0.02664 0.00 XLM12 1.51 0.00 0.01 0.05530 0.00310 0.2043 0.0161 0.02663 0.00 XLM24 1.51 0.00 0.01 0.05588 0.00256 0.2041 0.01228 0.02663 0.00 XLM25 1.98 0.00 0.01 0.1130 0.01294 0.44026 0.07709 0.02870 0.00 XLM05 0.44 0.00 0.01 0.10130 0.01294 0.44026 0.07709 0.02870 0.00 XLM05 0.44 0.00 0.01 0.10130 0.01293 0.00960 0.01212 0.02880 0.000 XLM05 0.44 0.00 0.02 0.02850 0.000 0.02860 0.000 0.02864 0.000 0.02864 </td <td></td> <td>XLM13</td> <td>1.27</td> <td>0.00</td> <td>0.00</td> <td>0.05670</td> <td>0.00317</td> <td>0.20845</td> <td>0.01188</td> <td>0.02762</td> <td>0.00095</td>		XLM13	1.27	0.00	0.00	0.05670	0.00317	0.20845	0.01188	0.02762	0.00095
XLM18 1.29 0.00 0.00 0.0184 0.00389 0.2289 0.01404 0.02648 0.00 XLM19 1.95 0.00 0.01 0.05530 0.00301 0.2043 0.0161 0.02666 0.00 XLM24 1.51 0.00 0.01 0.05982 0.00344 0.19805 0.02857 0.00 XLM25 1.98 0.00 0.01 0.05982 0.00344 0.1228 0.02687 0.00 XLM27 1.93 0.00 0.01 0.1130 0.01294 0.44026 0.07012 0.02870 0.00 XLM07 0.41 0.00 0.01 0.10130 0.01294 0.44026 0.07012 0.02870 0.00 XLM06 0.47 0.00 0.02 0.08698 0.0088 0.0966 0.0223 0.0269 0.008 XLM16 0.55 0.00 0.02 0.08572 0.00590 0.1787 0.01968 0.02668 0.000 XLM12 1.63 0.00 0.02 0.05514 0.01970 0.0274 0.000 X		XLM14	1.82	0.00	0.03	0.05473	0.00289	0.19620	0.00997	0.02676	0.00068
XLM19 1.95 0.00 0.01 0.05530 0.00311 0.2043 0.0161 0.02666 0.00 XLM20 2.53 0.00 0.02 0.05379 0.00213 0.19805 0.00855 0.02651 0.00 XLM25 1.98 0.00 0.02 0.05588 0.00256 0.2061 0.0128 0.02667 0.00 XLM27 1.93 0.00 0.01 0.0130 0.01294 0.44026 0.07709 0.02870 0.00 XLM07 0.44 0.00 0.01 0.0133 0.0752 0.00726 0.2956 0.02122 0.02887 0.00 XLM07 0.41 0.00 0.02 0.08572 0.00796 0.5381 0.0366 0.00 XLM22 0.51 0.00 0.02 0.05570 0.0030 0.20488 0.03066 0.00 XLM22 0.51 0.00 0.05217 0.00488 0.0448 0.04683 0.00488 0.03051 0.22383 0.01560 0.02724<		XLM17	1.63	0.00	0.04	0.06507	0.00371	0.23344	0.01210	0.02664	0.00052
XLM20 2.53 0.00 0.02 0.05379 0.00213 0.19805 0.02851 0.00 0.00 XLM24 1.51 0.00 0.01 0.05982 0.00444 0.21782 0.01234 0.02663 0.00 XLM25 1.98 0.00 0.02 0.05662 0.00286 0.2081 0.01644 0.02667 0.00 XLM27 1.93 0.00 0.01 0.10130 0.01294 0.44026 0.07709 0.02670 0.00 XLM06 0.47 0.00 0.03 0.07352 0.00702 0.26956 0.02723 0.02709 0.00 XLM07 0.31 0.00 0.02 0.08898 0.00885 0.30960 0.02733 0.0700 0.37888 0.03066 0.000 XLM16 0.55 0.00 0.02 0.05570 0.0030 0.24986 0.000 0.000 0.02744 0.000 XLM16 0.55 0.00 0.02 0.05570 0.00330 0.24088 0.00727		XLM18	1.29	0.00	0.00	0.06184	0.00389	0.22389	0.01404	0.02648	0.00046
XLM20 2.53 0.00 0.02 0.05379 0.00213 0.19805 0.02851 0.00 0.00 XLM24 1.51 0.00 0.01 0.05982 0.00444 0.21782 0.01234 0.02663 0.00 XLM25 1.98 0.00 0.02 0.05662 0.00286 0.2081 0.01644 0.02667 0.00 XLM27 1.93 0.00 0.01 0.10130 0.01294 0.44026 0.07709 0.02670 0.00 XLM06 0.47 0.00 0.03 0.07352 0.00702 0.26956 0.02723 0.02709 0.00 XLM07 0.31 0.00 0.02 0.08898 0.00885 0.30960 0.02733 0.0700 0.37888 0.03066 0.000 XLM16 0.55 0.00 0.02 0.05570 0.0030 0.24986 0.000 0.000 0.02744 0.000 XLM16 0.55 0.00 0.02 0.05570 0.00330 0.24088 0.00727		XLM19									0.00046
XLM24 1.51 0.00 0.01 0.05982 0.00444 0.21782 0.0128 0.02663 0.00 XLM25 1.98 0.00 0.02 0.05662 0.00256 0.20491 0.0134 0.02667 0.00 XLM27 1.93 0.00 0.01 0.10130 0.01294 0.44026 0.07709 0.02870 0.00 XLM05 0.44 0.00 0.01 0.10130 0.01294 0.44026 0.07709 0.02870 0.00 XLM06 0.47 0.00 0.02 0.08698 0.00885 0.3960 0.02723 0.02709 0.00 XLM07 0.31 0.00 0.02 0.05872 0.0059 0.21787 0.0168 0.02696 0.00 XLM22 0.51 0.00 0.02 0.12528 0.00910 0.49936 0.0468 0.0306 0.00 XLM22 0.51 0.00 0.05217 0.00330 0.24008 0.01195 0.02678 0.000 XLM23											0.00042
XLM25 1.98 0.00 0.03 0.05588 0.00299 0.20491 0.01034 0.02687 0.000 Nim XLM27 1.93 0.00 0.01 0.1030 0.0124 0.44026 0.07799 0.02870 0.00 XLM05 0.44 0.00 0.01 0.0133 0.07352 0.00702 0.26956 0.02123 0.02709 0.02870 0.00 XLM06 0.47 0.00 0.02 0.08698 0.0085 0.30966 0.02123 0.02709 0.02870 0.00 XLM06 0.47 0.00 0.02 0.08698 0.0085 0.30966 0.02122 0.02068 0.00 XLM07 1.35 0.00 0.02 0.12523 0.00790 0.51881 0.03066 0.00 XLM15 0.55 0.00 0.02 0.05570 0.00330 0.01608 0.02724 0.00 XLM23 0.51 0.00 0.05514 0.00477 0.20333 0.01608 0.02724 0.00											0.00042
XLM27 1.93 0.00 0.02 0.05662 0.00256 0.2081 0.00968 0.02669 0.00 kim XLM05 0.44 0.00 0.01 0.0130 0.01294 0.44026 0.0709 0.02870 0.00 XLM06 0.47 0.00 0.03 0.07352 0.00702 0.26956 0.02122 0.02808 0.00 XLM07 0.31 0.00 0.02 0.08698 0.00885 0.30960 0.02723 0.02709 0.00 XLM16 0.55 0.00 0.00 0.12933 0.00790 0.53831 0.03588 0.03066 0.00 XLM16 0.55 0.00 0.02 0.15570 0.00330 0.24088 0.01195 0.02658 0.00 XLM22 0.51 0.00 0.00 0.05514 0.0044 0.19877 0.01966 0.02723 0.000 XLM28 0.58 0.00 0.02 0.06435 0.00248 0.19697 0.09277 0.02723 0.000<											0.00033
XLM06 0.47 0.00 0.03 0.07352 0.00702 0.26956 0.02122 0.02808 0.00 XLM07 0.31 0.00 0.02 0.08698 0.00855 0.3960 0.02723 0.02709 0.00 XLM08 0.73 0.00 0.00 0.12933 0.00790 0.53831 0.03588 0.03066 0.00 XLM15 0.55 0.00 0.02 0.12528 0.00910 0.49936 0.03468 0.03066 0.00 XLM21 1.63 0.00 0.02 0.05570 0.00330 0.20408 0.01195 0.02588 0.00 XLM22 0.51 0.00 0.00 0.05514 0.00477 0.20333 0.01608 0.02724 0.00 XLM23 0.69 0.00 0.00 0.05514 0.00477 0.20333 0.01608 0.02723 0.00 XLM24 1.93 0.00 0.02 0.06435 0.00521 0.23583 0.01727 0.02293 0.00 XLM30 1.01 0.00 0.02 0.08867 0.00476 0.302055<											0.00033
XLM06 0.47 0.00 0.03 0.07352 0.00702 0.26956 0.02122 0.02808 0.00 XLM07 0.31 0.00 0.02 0.08698 0.00855 0.3960 0.02723 0.02709 0.00 XLM08 0.73 0.00 0.00 0.12933 0.00790 0.53831 0.03588 0.03066 0.00 XLM15 0.55 0.00 0.02 0.12528 0.00910 0.49936 0.03468 0.03066 0.00 XLM21 1.63 0.00 0.02 0.05570 0.00330 0.20408 0.01195 0.02588 0.00 XLM22 0.51 0.00 0.00 0.05514 0.00477 0.20333 0.01608 0.02724 0.00 XLM23 0.69 0.00 0.00 0.05514 0.00477 0.20333 0.01608 0.02723 0.00 XLM24 1.93 0.00 0.02 0.06435 0.00521 0.23583 0.01727 0.02293 0.00 XLM30 1.01 0.00 0.02 0.08867 0.00476 0.302055<	Rim	XLM05	0.44	0.00	0.01	0.10130	0.01294	0.44026	0.07709	0.02870	0.00143
XLM07 0.31 0.00 0.02 0.08698 0.00885 0.30960 0.02723 0.02709 0.00 XLM08 0.73 0.00 0.00 0.12933 0.00790 0.53831 0.03588 0.03066 0.00 XLM15 0.55 0.00 0.05 0.05872 0.00509 0.21787 0.01968 0.02696 0.00 XLM21 1.63 0.00 0.02 0.12528 0.00910 0.49936 0.03468 0.03066 0.00 XLM22 0.51 0.00 0.02 0.05570 0.00330 0.20408 0.01195 0.026588 0.00 XLM23 0.69 0.00 0.00 0.05504 0.00477 0.20303 0.01608 0.02724 0.00 XLM26 1.93 0.00 0.02 0.06435 0.00521 0.23583 0.01727 0.02723 0.00 XLM30 1.23 0.00 0.02 0.068087 0.00476 0.30205 0.01766 0.02723 0.00 XLM30 1.23 0.00 0.3 0.523922 0.0193 5.07543											0.00080
XLM08 0.73 0.00 0.00 0.12933 0.00790 0.53831 0.03588 0.03006 0.00 XLM15 0.55 0.00 0.05 0.05872 0.00509 0.21787 0.01968 0.02966 0.00 XLM16 0.55 0.00 0.02 0.12528 0.0010 0.49936 0.03468 0.03066 0.00 XLM21 1.63 0.00 0.02 0.05570 0.00330 0.20408 0.01195 0.02588 0.00 XLM22 0.51 0.00 0.00 0.05217 0.00494 0.19877 0.01966 0.02784 0.00 XLM26 1.93 0.00 0.04 0.05438 0.00477 0.20303 0.01608 0.02724 0.00 XLM28 0.58 0.00 0.02 0.06435 0.0021 0.23583 0.01727 0.02639 0.00 XLM30 1.23 0.00 0.02 0.08087 0.00476 0.30205 0.01766 0.02723 0.00 XLM30 1.23 0.00 0.3 0.523922 0.01993 5.07543 <td rowspan="11"></td> <td></td> <td></td> <td></td> <td></td> <td></td> <td></td> <td></td> <td></td> <td></td> <td>0.00103</td>											0.00103
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TABLE 1 LA-ICP-MS U-Pb dating results of cassiterite from the Baimianshan segment, Xinlu Sn-Zn deposit, South China.

(Continued on following page)

Analysis spots	U-Th-Pb contents			U-Pb isotopic ratios					
	²³⁸ U	²³² Th	Common Pb	²⁰⁷ Pb/ ²⁰⁶ Pb	1σ	²⁰⁷ Pb ^{/235U}	1σ	²⁰⁶ Pb/ ²³⁸ U	1σ
AY-4std	3.14	0.00	0.0	0.11874	0.01239	0.46164	0.04956	0.02719	0.00104
AY-4std	0.08	0.00	0.0	0.10436	0.02845	0.30422	0.07150	0.02689	0.00149
AY-4std	15	0.00	0.0	0.08293	0.00559	0.31168	0.02182	0.02677	0.00071
AY-4std	66	0.00	1.7	0.12044	0.01668	0.46544	0.05747	0.02637	0.00092
AY-4std	0.97	0.00	0.0	0.08996	0.00466	0.32390	0.01752	0.02615	0.00040
AY-4std	213	0.00	3.9	0.12359	0.01338	0.45961	0.04720	0.02611	0.00072
AY-4std	49	0.00	0.5	0.06693	0.00801	0.23368	0.02575	0.02497	0.00077
AY-4std	14	0.00	1.0	0.07360	0.01043	0.21904	0.02460	0.02300	0.00071
NIST reference glass S	RM 612								
SRM 612	98	37		0.92112	0.12499	32.90046	3.71038	0.24484	0.00826
SRM 612	101	39		0.92721	0.11829	33.63997	3.56754	0.24918	0.00790
SRM 612	100	38		0.94022	0.07948	34.05289	2.52080	0.25511	0.00470
SRM 612	100	38		0.94842	0.08621	34.95647	2.77951	0.25968	0.00513
SRM 612	99	37		0.84393	0.04970	27.21569	1.65231	0.22700	0.00365
SRM 612	100	38		0.83753	0.04991	26.80127	1.64293	0.22487	0.00360
SRM 612	37	37		0.93942	0.07635	29.77904	2.28620	0.22463	0.00840
SRM 612	38	39		0.94003	0.07406	29.79229	2.22115	0.22455	0.00813
SRM 612	37	37		0.91044	0.02008	28.89959	0.72585	0.22739	0.00214
SRM 612	38	38		0.90875	0.02065	28.97643	0.74343	0.22848	0.00216
SRM 612	38	39		0.92143	0.01898	28.86141	0.71193	0.22596	0.00157
SRM 612	36	36		0.92808	0.01973	29.00782	0.72734	0.22568	0.00184

TABLE 1 (Continued) LA-ICP-MS U-Pb dating results of cassiterite from the Baimianshan segment, Xinlu Sn-Zn deposit, South China.

(Gemmrich et al., 2021; Routhier, 1980), the Dachang Sn-Zn district (Huang et al., 1986; Fan et al., 2004), and the Bawang Fe-Zn-Sn deposit of Wuxu Sb-Zn district, South China (Xiao et al., 2022). Based on the two above-mentioned works, the generated elemental maps can provide information on the mineralization process and more meaningful results than a single micro-observation.

Reflection and element mapping can also greatly improve the efficiency of positioning for cassiterite U-Pb dating when the sample contains minor cassiterite. Commonly, it is hard to identify the position of cassiterite in the coexistence of finegrained cassiterite and sphalerite sulfide. In our study, the element mapping by μ -XRF all display the distribution of tin (Figures 6–8), which was then confirmed to be cassiterite by quick microscopic observation (Figures 5C,D). Therefore, we confirm that the cassiterite was only formed in the retrograde skarn-sulfide stage. Moreover, the element mapping also avoided potential mixed ages of cassiterite. This method is also a powerful tool for *in situ* geochronology of polyphase minerals and provides a more accurate process of mineralization.

Age relationship between Sn-Zn (-W) mineralization and magmatism of Guposhan-Huashan district during Jurassic

Published high-precision geochronological investigations, including zircon U-Pb and biotite ⁴⁰Ar-³⁹Ar dating on the causative intrusions and multiple dating methods (Sm-Nd, Re-Os, ⁴⁰Ar-³⁹Ar, Rb-Sr) on the skarn and hydrothermal minerals, have revealed the magma emplacement and mineralization in the Guposhan-Huashan district. However, the relationship between the Sn-Zn (-W) mineralization and magmatism remains confusing based on these studies. An overview of the timing of magmatic-hydrothermal ore formation in the Guposhan-Huashan district is described as follows.

The Huashan and Guposhan complex plutons are the main intrusions in this district. The Huashan complex pluton composes of Huashan biotite granite in the core, the Tong'an quartz monzonite in the west, and the Niumiao diorite in the southeast. Available zircon U-Pb ages show these three intrusions

were formed at 160–163 Ma (Zhao et al., 2010). The $\varepsilon_{\text{Hf}}(t)$ (-2.8 to +7.4), $\varepsilon_{\text{Nd}}(t)$ (-3.18 to +0.91), and initial $({}^{87}\text{Sr}/{}^{86}\text{Sr})$ ratios (0.70472-0.70714) revealed these three intrusions were generated by mantle-crust interaction (Cai et al., 2018 and references therein; Zhao et al., 2010; Zhu et al., 2005, 2006). Zhu et al. (2006) reported the younger felsic dykes with an age of 148.0 \pm 4.0 Ma intruded into the Huashan granite along the fractures. The Guposhan complex pluton includes Lisong hornblende biotite granite in the core, Guposhan coarse biotite granite, and Fengmuchong fine felsic dykes in the south (Figure 2A). Published zircon U-Pb ages show that the Lisong and Guposhan granites were formed at 161.0-163.0 Ma and 159.5-165.0 Ma, respectively (Zhu et al., 2006; Gu et al., 2007b; Zhao et al., 2010; Shu et al., 2011; Cai et al., 2020; Cao et al., 2020). Li et al. (2016) dated the causative Shimen quartz monzodiorite in the Xinlu deposit and Fengmuchong biotite granite in the Shuiyanba deposit by igneous biotites ⁴⁰Ar-³⁹Ar dating at 168.7 \pm 1.9 Ma and 165.0 \pm 1.1 Ma, respectively. The $\varepsilon_{\text{Hf}}(t)$ (-3.6 to -2.3), $\varepsilon_{\text{Nd}}(t)$ (-5.5 to +1.9), and initial (⁸⁷Sr/⁸⁶Sr) ratios (0.70473-0.70742) of the Lisong and Guposhan granites and dark mafic enclaves imply the Guposhan main pluton also resulted from mantle-crust interaction. The later felsic dykes were dated by zircon U-Pb at 146.8 \pm 2.3 Ma to 154.2 \pm 3.1 Ma (Zhu et al., 2006; Wang et al., 2014). Initial Sr isotopic data ((⁸⁷Sr/ ⁸⁶Sr)_i = 0.7173) and $\varepsilon_{\rm Nd}(t)$ (-5) of later felsic dykes are obvious different from those of the main pluton in this district (Gu et al., 2007b). The high initial Sr isotopic and low $\varepsilon_{Nd}(t)$ values suggest the later dykes were generated from crust components. Thus, based on geochronology, two episodes of magmatism (159.5-168.7 Ma and 146.8-154.2 Ma) occurred in the Guposhan-Huashan district during Jurassic, of which the former was predominant.

Our robust in-situ LA-ICP-MS cassiterite U-Pb data, integrated with previous geochronological studies, support one stage magmatic-hydrothermal mineralization associated with 159.5-168.7 Ma magmatism in this district. Our new age directly yielded from cassiterite in this study demonstrates that the Sn-Zn mineralization in the Baimianshan segment occurred at 168.7 \pm 1.3 Ma, coinciding with the age of quartz monzodiorite in the Xinlu deposit. Cassiterite U-Pb age is a little older than muscovite ⁴⁰Ar- 39 Ar age of skarn-type ores (160.0 ± 2.0 Ma) and molybdenite Re-Os age of quartz-vein type wolframite ores (154.4 \pm 3.5 Ma) from the Liuheao segment (Feng et al., 2019). Due to the lack of detailed microscopic observation, the muscovite 40Ar-39Ar isotopic data are doubtful to represent the skarn stage. Li et al. (2016) obtained hydrothermal muscovite 40 Ar- 39 Ar age of 160.6 ± 1.5 Ma for the Liuheao altered granite. Therefore, we could speculate the muscovite ⁴⁰Ar-³⁹Ar isotopic data obtained by Feng et al. (2019) may represent the age of the quartz-vein stage rather than the skarn stage. This conclusion could be supported by the euhedral sericite-quartz of stage III replacing the massive skarn-type pyrrhotite and sphalerite of stage II (Figures 5A,H). From a temporal perspective, the skarn Sn-Zn mineralization is associated with quartz monzodiorite.

Spatially, the skarn Sn-Zn mineralization is related to the Guposhan coarse biotite granite.

In the following, we also compare the obtained cassiterite U-Pb age to published geochronological data of mineralization in the Guposhan-Huashan district to further constrain the Sn-Zn-W formation age. Previous work on the mineralization age of the Shuiyanba Sn deposit used muscovite 40 Ar- 39 Ar dating (162.0 ± 1.9 Ma and 162.5 ± 1.2 Ma; Kang et al., 2012; Li et al., 2016) of quartz + muscovite + wolframite veins and Rb-Sr dating of quartz fluid inclusions (136.1 \pm 2.0 Ma; Gu et al., 2007a) from the wolframite + scheelite + arsenopyrite + cassiterite quartz veins in the Lantoushan segment. It is difficult to confirm whether the Rb-Sr isotopes of different hand specimens were at isotopic equilibrium or the ore-forming fluids have homogeneous initial 87Sr/86Sr ratios. Furthermore, quartz from the W-Sn quartz veins has a different initial 87Sr/86Sr ratio (0.72015-0.72735) from the Guposhan-Huashan pluton (0.7056-0.7066) (Gu et al., 2007b). To sum up, the muscovite ⁴⁰Ar-³⁹Ar dating results are more reliable and the quartz vein type W-Sn mineralization of the Shuiyanba Sn deposit probably occurred at 162 Ma.

Published Sm-Nd isotopic data of skarn minerals (pyroxene, hornblende, garnet) from the skarn-type ores in the Chuanlingjiao segment of the Keda Sn deposit suggest the skarn minerals occurred at 160.4 ± 4.8 Ma (Cao et al., 2020). Li et al. (2016) obtained muscovite 40 Ar- 39 Ar ages of 164.5 ± 2.6 Ma for W-Sn quartz veins hosted in the Devonian marble at Maohedong segment and 161.2 ± 1.3 Ma for wolframite bearing quartz veins hosted in the Devonian sandstone at Sanchachong segment. According to the above, both the skarn Sn and quartz vein W-Sn mineralization occurred during 160-164 Ma.

As aforementioned, we conclude the Sn-W mineralization in the Guposhan-Huashan district was related to the intermediate felsic intrusions rather than the later felsic dykes. The skarn mineral Sm-Nd isotopic data represent the formation age of the prograde skarn stage (Stage I). Our new cassiterite U-Pb age represents the formation age of retrograde massive skarn-sulfides (stage II). All the muscovite ⁴⁰Ar-³⁹Ar ages represent the formation ages of quartz sulfide veins (stage III). No geochronological data of barren carbonate stage (stage IV) have been obtained in the Guposhan-Huashan district to date.

Relationship between the subduction of Paleo-Pacific plate and Jurassic tintungsten mineralization in the Nanling belt of South China

To understand the timing of magmatic-hydrothermal ore formation on the scale of the Nanling belt and SCB, we incorporated our new geochronological data into the existing



geochronological framework of the Nanling Sn-W belt and South China W-Sn provinces during the Jurassic. On the probability density distribution plots of the Jurassic intrusions, two main peaks (~142 and 155 Ma) and two small peaks (~179 and 189 Ma) can be recognized (Figure 10A). The former two peaks with dominated negative $\varepsilon_{\rm Hf}(t)$ values were associated with W-Sn mineralization, whereas the latter two peaks with dominated positive $\varepsilon_{Hf}(t)$ values were associated with Mo-Co mineralization (Figure 10B; Li et al., 2018; Yang et al., 2021). The four age clusters indicate multiphase tectonic events in the SCB. On the probability density distribution plots of the Jurassic W-Sn mineralization, a single peak (~157 Ma) is observed. The W-Sn formation ages overlapped with the second peak of intrusions (Figure 10A). The spatial distribution of W-Sn mineralization and associated intrusions display a slight southeastward younging trend in the Nanling Belt. However, in the entire South China, this trend is not obvious (Figure 1C). Recent research shows that W-Sn mineralization causative intrusions are A-type granites (Li et al., 2018; Xiao et al., 2019; Mao et al., 2021b). It is commonly accepted that the Jurassic mafic-ultramafic intrusions, A-type granites, and bimodal basalt-rhyolite volcanic rocks are formed in the extensional setting, but the geodynamic setting is still puzzling (Zhou and Li, 2000; Zhou et al., 2006; Li and Li, 2007; He et al., 2010; Mao et al., 2021b; Cao et al., 2021). The generally accepted dynamic domain includes the Paleo-Tethyan domain to the southwest and the Paleo-Pacific domain to the southeast. Therefore, two models have been proposed for the Jurassic evolution, including the Paleo-Tethyan post-collisional intra-continental extension model (He et al., 2010; Gan et al., 2017) and Izanagi or Paleo-Pacific plate subduction-related model (Zhou et al.,

2006; Li and Li, 2007; Zaw et al., 2007; Mao et al., 2021b; Cao et al., 2021; Gan et al., 2021). Although the latter view is more popular, it is inconsistent with the fact that the orientation of the distribution of Middle-Late Jurassic rocks is opposite to the subduction of the Paleo-Pacific plate (Gan et al., 2017). The difference in geochemical signatures between the Early-Middle Jurassic mafic and shoshonitic rocks and Cretaceous arc-related rocks in the SCB also indicate that the Paleo-Pacific slabderived components didn't modify the lithospheric mantle (Gan et al., 2021 and references therein). This conclusion is also confirmed by previous whole rock geochemistry and Sr-Nd-Hf isotope and zircon Hf-O isotope data from the early Jurassic gabbro-syenite-alkali-feldspar granite in South China (Yang et al., 2021). As described above, many facts are still in debate in the Paleo-Pacific subduction model, such as the variations of slab dip angles (Zhou et al., 2006; Li et al., 2018); flat-slab subduction and slab foundering (Li and Li, 2007); polyphase slab subduction and retreat (Li et al., 2020); slab window in the back-arc region (Mao et al., 2021b); and mantle upwelling model (Cao et al., 2021). According to the available data, the Paleo-Pacific subduction model is currently the most reasonable one. First, the geochemical signatures and higher zircon $\epsilon Hf_{(t)}$ - $\delta^{18}O$ values of the Middle-Late Jurassic igneous rocks all indicate an extensional setting corresponding to asthenospheric upwelling and crust-mantle interaction (Cao et al., 2021; Gan et al., 2021 and references therein). Previous studies have shown that tin granites generally occurred in an extensional environment corresponding to mantle upwelling (Romer and Kroner, 2016; Xiao et al., 2019; Guo et al., 2021). The upwelling of the basaltic intrusions might provide additional heat and trigger a partial melting of the lower crust to form W-Sn causative granite. Second, the Jurassic compressional deformation in the SCB is consistent with the coeval advancing motion of the Izanagi trench (Li X. F. et al., 2012; Cao et al., 2021). The Jurassic arc-related granitoids have been also observed in the coastal regions and the South China Sea basin (Xu et al., 2017). Cao et al. (2021) also pointed out that Jurassic magmatism was accompanied by crustal deformation in the SCB. Last, numerical models suggest that flat-slab exists (Dai et al., 2020).

Conclusion

- (1) Thin section reflection and μ -XRF element mapping represent a powerful tool for determining paragenesis in ore deposits and are highly effective to guide the selection of analytical points for cassiterite U-Pb dating. This method may also be applicable for *in-situ* dating of other minerals.
- (2) Four mineralization episodes are identified in the Baimianshan segment, Xinlu deposit. They are prograde skarn (stage I), retrograde skarn-sulfide (stage II), quartz sulfide veins (stage III), and barren carbonate (stage IV) stages. The retrograde skarn-sulfide stage is the main mineralization stage and consists of massive pyrrhotite, and sphalerite with parts of epidote, cassiterite, and minor chalcopyrite and arsenian pyrite.
- (3) The time of retrograde skarn-sulfide mineralization is dated by LA-ICP-MS cassiterite U-Pb at 168.7 ± 1.3 Ma. Our robust *insitu* LA-ICP-MS cassiterite U-Pb data, coupled with previous geochronological studies, support one stage magmatichydrothermal mineralization associated with 159.5–168.7 Ma magmatism in the Guposhan-Huashan district.
- (4) The Xinlu Sn-Zn mineralization and the Middle-Late Jurassic Sn-W mineralization in the Nanling belt, South China corresponded to asthenospheric upwelling and crust-mantle interaction caused by the subduction of the Paleo-Pacific plate.

Data availability statement

The original contributions presented in the study are included in the article/supplementary material, further inquiries can be directed to the corresponding author.

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Author contributions

C-HX designed the study, collected the samples for analysis, and compiled the results and wrote the manuscript. P-PY and X-HL collected and processed previously published data. X-HL also finished the figures and table. C-HX and W-LW performed the μ -XRF analysis and process of the data. Y-WT performed cassiterite U-Pb dating and process of the data. X-CL, P-PY, YZ provided feedback to the manuscript.

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Conflict of interest

The authors declare that the research was conducted in the absence of any commercial or financial relationships that could be construed as a potential conflict of interest.

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