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One-degree resolution mascon solution over Antarctic derived from GRACE Level-2 data

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The mass loss of the Antarctic Ice Sheet (AIS) is an important contributor to global sea-level rise in response to the warming ocean and atmospheric temperatures as well as the changes in current systems and precipitation patterns. In this study, a regional mascon method is developed to squeeze more mass change signals, in which the pseudo-observations of the geopotential are generated from the unfiltered GRACE Level-2 data whereas the regularization matrix is constructed with the prior information derived from filtered GRACE Level-2 data. A series of mascon solutions with $1^{\circ} \times 1^{\circ}$ equal-area resolution over AIS is derived from the updated Tongji-Grace2018 model spanning April 2002 to December 2016. Compared to the filtering results from P4M6 decorrelation and 100 km Gaussian filtering, our mascon solutions can effectively suppress the strips, improve the spatial resolution over AIS, and get a stronger signal with an improvement of 116.86% in the Antarctic Peninsula Ice Sheet (APIS), and more coincide with the features of glaciers and ice streams, such as the most striking ice mass loss in Totten, Getz, Thwaites and Pine Island, and the ice mass gain in Kamb Ice Stream. During the period from 2002 to 2016, the mass change rates from our mascon solution are -103.6 ± 5.6 Gt/ yr, 63.0 \pm 4.3 Gt/yr, -143.3 \pm 4.9 Gt/yr and -23.29 \pm 1.2 Gt/yr in AIS, East AIS, West AIS, and APIS, respectively. The mass change signals at the basin scale are with even more distinguishing features, with the highest mass gain rates of 18.03 ± 1.88 Gt/yr and 14.55 ± 0.60 Gt/yr at Basin 7 and Basin 18, and the highest mass loss rates of -58.57 ± 2.48 Gt/yr and -44.12 ± 2.27 Gt/yr at Basin 21 and Basin 22. Relative to the cumulated surface mass balance from the regional atmospheric climate model, the correlation coefficients of our mascon solutions are 0.91, 0.94, and 0.96 in East AIS, West AIS, and APIS.

KEYWORDS

satellite gravity, time variable gravity, GRACE, Antarctica, inverse theory

1 Introduction

The contribution to global sea-level rise from the melting of polar ice sheets has been a focus of intensive study over the past several decades (Harig and Simons, 2012; Iz et al., 2021). As the largest ice sheet over the globe, the Antarctic Ice Sheet (AIS) is one of the major contributors to global mean sea level rise (Vaughan et al., 2013; Pattyn et al., 2018; Rignot et al., 2019). Therefore, the accurate estimate of mass changes over AIS is of utmost importance (Shepherd et al., 2018). Since the advent of satellite observations, three types of techniques have been used to estimate the mass balance over AIS: 1) Satellite altimetry method at intermediate resolution (1–10 km) by using the direct measurements of elevation changes combined with climatological/glaciological models for firn (snow) density and compaction (Shepherd et al., 2019); 2) input-output method (IOM) at a high resolution (100 m–1 km) by measuring the ice flow velocities with synthetic aperture radar data over outlet glaciers combined with glacier

Time span	GRACE/GRACE-FO model	Mass loss rates (Gt/yr)	GIA model	Reference	
2003-2011	CSR RL04 & RL05 GFZ RL04 & RL05	83 ± 36	empirical GIA model ^a	Barletta et al. (2013)	
2003-2013	CSR RL05	91 ± 26	W12a ^b	Schrama et al. (2014)	
2003-2014	CSR RL05	92 ± 10	IJ05_R2 ^c	Harig and Simons (2015)	
2003-2010	CSR RL05	165 ± 72	ICE5G ^d	Jacob et al. (2012)	
2003-2012	CSR RL05	107 ± 34	W12a ^b	Mu et al. (2017)	
2003-2013	CSR RL05	67 ± 44	IJ05_R2 ^c	Velicogna et al. (2014)	
2003-2013	CSR RL05	104 ± 5	IJ05_R2 ^c	Groh et al. (2019)	
2003-2012	CSR RL05	83 ± 49	IJ05_R2 ^c	Velicogna and Wahr (2013)	
	Contracto	147 ± 80	ICE5G ^d		
2002-2017	CSR RL06	163 ± 5	A model ^e	Zou et al. (2019)	
2002-2019		89 ± 43		Groh and Horwath. (2021)	
2002-2020	CSR RL06	90.9 ± 43.5	- IJ05_R2 ^c		
2002-2019	JPL RL06	126 ± 28	IJ05_R2 ^c	Loomis et al. (2019a)	
2002-2019	CSR RL06	107 ± 55			
	JPL RL06	104 ± 57	IJ05_R2 ^c	Velicogna et al. (2020)	
	GFZ RL06	89 ± 60	-		
2002-2015	ITSG-Grace2016	95 ± 50	W12a ^b	Forsberg et al. (2017)	

TABLE 1 Mass loss rates over AIS with GIA corrections.

^aRiva et al., 2009. ^bWhitehouse et al., 2012. ^cIvins et al., 2013. ^dPeltier, 2004.

^eGeruo et al., 2013.

thickness data to derive the ice discharge and surface mass balance (SMB) (van Wessem et al., 2018; Rignot et al., 2019); 3) satellite gravimetry and airborne gravimetry by measuring the gravity change due to ice mass variation, such as the satellite mission of gravity recovery and climate experiment (GRACE) (Velicogna and Wahr, 2006). Besides the GRACE mission, the other two techniques can hardly directly provide the AIS mass estimation with the monthly resolution, single satellite altimetry is restricted by temporal sampling issues and converting elevation change into mass change which requires assumptions on density due to firn density not known very well (Khan et al., 2015), and the penetration depth of radar altimetry may change during seasons of freezing-thaw (Nilsson et al., 2015). The IOM is limited by the annual temporal resolution of ice discharge (Rignot et al., 2019).

GRACE provides the direct measurement of ice mass change with a temporal resolution of 1 month at an error level of 2 cm in terms of equivalent water height (EWH) (Wahr et al., 2006; Tapley et al., 2019), which is less influenced by the temporal sampling issues and unknown surface properties compared to the single satellite altimetry mission. By using the GRACE gravity field solutions, many estimates related to the mass loss rates over AIS have been achieved (Table 1), though these results are very different from 67 ± 44 Gt/yr (Velicogna et al., 2014) to 165 ± 72 Gt/yr (Jacob et al., 2012), due to using different GRACE solutions, different Glacial Isostatic Adjustment (GIA) models and methodologies, as well as in different periods. Previous studies mainly focus on the estimation at large scale, i.g., the whole AIS, west AIS, east AIS, and APIS, since the estimation at the basin scale is limited by the GRACE spatial resolution. Moreover, the AIS mass changes before and after the epoch between 2007–2008 are quite different (Loomis et al., 2019a; Loomis et al., 2020), which are worthy to be further investigated.

The GRACE gravity field solutions are expressed in terms of spherical harmonic coefficients (SHCs) or mascon solutions. Since significant north-south striping errors exist in the map of estimated mass change directly from GRACE SHCs solutions, several postprocessing techniques, such as Gaussian filters (Jekeli, 1981; Wahr et al., 1998), decorrelation filters (Swenson and Wahr, 2006), and DDK filters (Kusche et al., 2009), have been developed to mitigate these errors. However, these filters all cause the attenuation of real geophysical signals (Tapley et al., 2019) and leakage errors that reach 20 Gt/yr in AIS (Chen et al., 2015; Mu et al., 2017). To mitigate leakage error and increase the spatial resolution, Chen et al. (2021) developed the global regularized SHC solutions up to degree and order (d/o) 180 that is comparable with the three global mascon solutions, e.g., CSR (Center for Space Research, the University of Texas) mascon (Save et al., 2016), JPL (Jet Propulsive Laboratory) mascon (Watkins et al., 2015) and GSFC (Goddard Space Flight Center) mascon (Loomis et al., 2019b) solutions. JPL and GSFC mascon solutions are both developed from the GRACE Level-1B observations while CSR mascon solution is generated from the SHCs up to degree and order 120. The regularization matrix is constructed differently, from the GRACE information (CSR), the

prior geophysical signals (JPL), and the spatial/temporal correlated exponential function (GSFC). However, operating on the Level-1B observations is very complex with a huge computation burden, and the implementation and further improvement of them are almost impossible for others (Baur and Sneeuw, 2011; Ran, 2017). Besides, different global mascon solutions show obvious discrepancies at the regional scale for hydrology (Scanlon et al., 2016). For example, the JPL mascon solution overestimates the decreasing and increasing rates of the terrestrial water storage anomalies (TWSA) over the Tibet Plateau relative to other GRACE solutions (Jing et al., 2019). Moreover, when focusing on the signals at the region less than $3^{\circ} \times 3^{\circ}$, global mascon solutions should be used with caution (Zhang et al., 2019). Therefore, the SHCs are still used as the standard solutions for GRACE data while needing further processing to reduce noise in higher degree SHCs and restore regional signals due to the cut-off of SHCs. The regional mascon method utilizes the SHCs to generate the observations, also called pseudo-observations, the mascons to be solved are set in the place interested with the expected resolution, and only regional prior information is needed to construct the regularization matrix, therefore it is much easier to be implemented (Ran, 2017). The pseudo-observations can be either the geoid heights (Sasgen et al., 2010), filtered mass changes at the ground (Schrama and Wouters, 2011; Yi and Sun, 2014) or the radial gravity disturbances at mean satellite altitude (Forsberg and Reeh, 2007; Sørensen and Forsberg, 2010; Baur and Sneeuw, 2011). Among these pseudo-observations, the radial gravity disturbances at GRACE satellite altitude are more natural as the synthesized data better resemble original observations at satellite altitude. However, the observation information in the other two directions is usually abandoned except for Su et al. (2019) who considered the pseudo-observations of three directions in their mascon solutions. Moreover, previous studies on the regional mascon methods did not pay enough attention to the construction of the regularization matrix and used an identity matrix as a regularization matrix to stabilize the mascon solutions (Forsberg and Reeh, 2007; Sørensen and Forsberg, 2010; Baur and Sneeuw, 2011; Barletta et al., 2013; Yi and Sun, 2014). However, the regularization matrix plays an important role in leakage error correction and should reflect the power spectrum of the parameters. Using an identity matrix means all parameters with the same power spectrum and the signal leakage remains between adjacent basins (Barletta et al., 2013). Therefore, we introduce the regularization matrix with prior information from filtered GRACE SHCs into the mascon modeling method to reduce signal leakage and improve the spatial resolution of the mascon solutions.

Considering gravity disturbance contains full gravity information, in this paper, we propose a gravitational potential-based regional mascon method where the gravity disturbance at satellite altitude is used as the pseudo-observations to construct the observation equation. The proposed mascon method is used to derive the regional mascon solutions over AIS from April 2002 to December 2016 with a regularization matrix constructed with the prior information from filtered SHCs. The generated mascon solutions are first compared with the filtered counterparts to preliminary show the improvement in signal recovery and the sensitivity of mascon solutions to the prior information. Then, the spatiotemporal mass change signals of our mascon solutions over AIS are analyzed in different periods by compared with those of CSR, GSFC, and JPL mascon solutions. Subsequently, using the detrend mass change of the monthly time series of cumulative surface mass balance (SMB) from the regional atmospheric climate model (RACMO), the interannual mass change variations over AIS and its subregions are also analyzed. The rest of the paper is organized as follows. The used data and proposed mascon method are described in Section 2, 3 respectively. In Section 4, the mascon solutions are generated, and then the mass change rates and interannual signals over AIS are analyzed. Finally, conclusion are drawn in Section 5.

2 Data

2.1 GRACE data

The Tongji-Grace2018 model up to degree and order 90 (Chen et al., 2019) from April 2002 to December 2016 is used in our mascon solutions. The Earth's geo-center corrections are applied with the degree-1 coefficients from Technical Note-13 (Swenson et al., 2008; Sun et al., 2016). All C_{20} coefficients of the Tongji-Grace2018 model are replaced by those from Technical Note-14 as recommended by Loomis et al. (2020) and the C_{30} terms after August 2016 are also replaced by those from Technical Note-14 (Loomis et al., 2019a; Loomis et al., 2020). The regional IJ05_R2 model is applied to correct the long-term GIA signals (Ivins et al., 2013).

Three global mascon solutions from the CSR (Save et al., 2016), GSFC (Loomis et al., 2019b) and JPL (Watkins et al., 2015) are also used to compare the results of our mascon solutions. The CSR mascon solutions can be downloaded from the website of http://www2.csr. utexas.edu/grace/RL06_mascons.html, the GSFC mascon solutions from the link https://earth.gsfc.nasa.gov/geo/data/grace-mascons and JPL mascon solutions from https://grace.jpl.nasa.gov/data/getdata/jpl_global_mascons. It is worthwhile to mention that the three GRACE mascon solutions all use the global ICE-6G model (Peltier, et al., 2018) to remove the GIA signals. To directly compare the results with the previous studies, we added back the GIA signals with the ICE-6G model and then deduct the GIA signals with the IJ05_R2 model according to the recommendation of Ivins et al. (2013).

2.2 RACMO data

The regional atmospheric climate model (RACMO) is developed by the Institute for Marine and Atmospheric Research Utrecht at Utrecht University. RACMO generates the SMB at 27-km resolution by using the physics package of the Integrated Forecast System of the European Centre for Medium-Range Weather Forecasts along with the High-Resolution Limited Area Model (Undén et al., 2002; van Wessem et al., 2014; Rignot et al., 2019). The new release version RACMO2.3p2 spanning from 1979 to 2016, which is updated from RACMO2.3p1, including improved topography, precipitation, and snow properties (van Wessem et al., 2018), is used to compute the interannual variation of surface mass balance over AIS. The time series of cumulative SMB anomalies relative to the average of 1979-2008 (Velicogna et al., 2014; Rignot et al., 2019) are then deducted with the average of the study period, the same as that in processing the GRACE time series. Focusing on the annual and interannual variability over AIS, the ice discharge and linear trend from cumulative SMB should be excluded. Therefore, we finally detrend the GRACE series and cumulative SMB series and apply a 7-month smooth to them.



3 Mascon modeling and its regularized solution

3.1 Gravitational potential-based mascon modeling

Taking the Earth's elastic deformation into account, the disturbing gravitational potential δT_i at GRACE satellite attitude can be computed with the unfiltered SHCs as follows,

$$\delta T_i = T\left(\lambda_i, \varphi_i, \rho\right) = \frac{GM}{\rho} \sum_{l=1}^{l_{\max}} \left(\frac{a}{\rho}\right)^l \frac{1}{\left(1+k_i'\right)} \sum_{m=0}^l P_{lm}\left(\sin\varphi_i\right) \left(\Delta C_{lm}\cos m\lambda_i + \Delta S_{lm}\sin m\lambda_i\right)$$
(1)

where, $(\lambda_i, \varphi_i, \rho)$ are longitude, latitude, and radial distance of a space point, *a* is the radial distance of the points at the ellipsoidal surface and it is the function of the latitude, i.e., $a = R(1 - f)/\sqrt{1 - \sin^2 \theta}$, *R* is the radius of the spherical Earth, *f* is the ellipsoid flattening (WGS84 value: f = 1/298.257223563) and θ is the colatitude (Li et al., 2017; Ditmar 2018). $\rho = a + 500$ km, *GM* indicates the product of gravitational constant *G* and the mass of the Earth *M*, P_{lm} denotes the fully normalized associated Legendre functions of degree *l* and order *m*, k'_l is the load Love number of degree *l*; ΔC_{lm} and ΔS_{lm} are the unfiltered SHCs after the mean for the study period is removed, l_{max} is the maximum degree.

Figure 1 shows the geometry of the space location $S(\lambda_i, \varphi_i, \rho)$ at the satellite attitude and ground point $P(\lambda_j, \varphi_j, a)$, where a point disturbing mass is located; $d_{i,j}$ and $\psi_{i,j}$ are the Euclidean distance and spherical distance between $S(\lambda_i, \varphi_i, \rho)$ and $P(\lambda_j, \varphi_j, a)$.

Supposing the disturbing potential δT_i at point *i* is caused by the *t* points of mass anomalies in the study area, according to Newton's law of gravitation we have the following expression,

$$\delta T_{i} = G \sum_{j=1}^{t} \frac{\delta m_{j}}{\left(a^{2} + \rho^{2} - 2a\rho\cos\psi_{i,j}\right)^{1/2}}$$
(2)

where δm_j is the point mass anomaly at the location *j*. Since the disturbing potential δT_i is derived from the SHCs with Eq. 1, it is called pseudo-observation, and the point mass unknown δm_i can be

solved from Eq. 2 with the pseudo-observations over the study area. If there are n (n > t) number of pseudo-observations, we reformulate Eq. 2 in the matrix form as follows,

$$y = Ax + e \tag{3}$$

where $\mathbf{y} = (\delta T_1 \cdots \delta T_n)^T$ denotes the *n*-vector of observations, \mathbf{A} is an $n \times t$ design matrix (its ij th element \mathbf{A} $(i, j) = G \frac{1}{(a^2 + \rho^2 - 2ar \cos \psi_{ij})^{1/2}}$), \mathbf{x} is a *t*-vector of mass anomalies to be estimated, $\mathbf{e} \sim (0, \sigma_0^2 \mathbf{P}^{-1})$ represents the random error vector, σ_0^2 and \mathbf{P} are the variance of unit weight and the weight matrix. The weight matrix \mathbf{P} can be directly derived from Eq. 1 *via* the law of error propagation as $\mathbf{P} = (\mathbf{B}\mathbf{D}\mathbf{B}^T)^{-1}$, where \mathbf{D} is the full covariance matrix of the SHCs and it can be obtained from the inverse of the normal equation when solving the SHCs. \mathbf{B} is the coefficient matrix when the right hand of Eq. 1 is written in a matrix-vector form. The *i* th row elements of \mathbf{B} are expressed as (Chen et al., 2020),

$$\begin{cases} \boldsymbol{B}(i, j_{\Delta C_{lm}}) = \frac{\mathrm{GM}}{\rho} \left(\frac{a}{\rho}\right)^{l} \frac{1}{1+k_{l}} P_{lm}(\sin\varphi_{i}) \cos m\lambda_{i} \\ \boldsymbol{B}(i, j_{\Delta S_{lm}}) = \frac{\mathrm{GM}}{\rho} \left(\frac{a}{\rho}\right)^{l} \frac{1}{1+k_{l}} P_{lm}(\sin\varphi_{i}) \sin m\lambda_{i} \end{cases}$$
(4)

where, $j_{\Delta C_{lm}}$ and $j_{\Delta S_{lm}}$ are the column numbers corresponding to ΔC_{lm} and ΔS_{lm} , respectively.

3.2 Regularized solution to the ill-conditioned mascon modeling

Since Eq. 3 is ill-conditioned, it is usually solved with Tikhonov regularization and the solution is derived by minimizing the following cost function (Tikhonov, 1963),

$$\Phi(\mathbf{x}) = (\mathbf{y} - \mathbf{A}\mathbf{x})^{\mathrm{T}} \mathbf{P}(\mathbf{y} - \mathbf{A}\mathbf{x}) + \alpha \mathbf{x}^{\mathrm{T}} \mathbf{R}\mathbf{x}$$
(5)

where α indicates the regularization parameter ($\alpha > 0$) and R is the symmetric regularization matrix. On the right hand of Eq. 5, the first term represents how well the mascons fit the temporal geopotential anomalies derived from the unfiltered SHCs, the second term denotes the total signal variance of the mascons. An identity regularization matrix was adopted by Baur and Sneeuw (2011) and Chen et al. (2016); Chen et al. (2020), which implies that all mascon parameters have the same signal variance. However, it is not reasonable for the mascons over AIS. Since the regularization matrix plays an important role in correcting the signal leakage and improving the spatial resolution, it should reflect, as far as possible, the signal strengths of the mascons to be solved. Therefore, the regularization matrix should be constructed with the signal variance of each mascon as,

$$\boldsymbol{R} = \begin{bmatrix} \frac{\sigma^2}{\sigma_1^2} & \cdots & 0\\ 0 & \ddots & 0\\ 0 & \cdots & \frac{\sigma^2}{\sigma_t^2} \end{bmatrix}$$
(6)

where *t* is the number of mascons, σ_i^2 ($1 \le i \le t$) represents the variance of mass change signal, the scale factor σ^2 can be chosen as the mean of σ_i^2 . To determine the reasonable signal variances over AIS, we first filter the Tongji-Grace2018 model using Gaussian smoothing (Wahr

	CSR	GSFC	JPL	This study
Coverage	global	global	global	AIS and its coastal ocean
Input data	GRACE SHCs	Level-1B data	Level-1B data	GRACE SHCs
Mascon size	$1^{\circ} \times 1^{\circ}$	1° × 1°	3° × 3°	$1^{\circ} \times 1^{\circ}$
Regularization Matrix	Filtered GRACE SHCs	Spatial/temporal correlated function	Geophysical models	Filtered GRACE SHCs
Regularized method	Tikhonov regularization using the L-ribbon approach	Iterative solution with an empirical damping parameter	Bayesian estimation and sequential Kalman filter	Tikhonov regularization by minimizing traced MSE
Error assessment	no	Data-driven method	no	MSE

TABLE 2 Summary of the mascon solutions.

et al., 1998) and P4M6 decorrelation (Swenson and Wahr, 2006), then estimate the signal variances of all mascons over AIS with the filtered model, same as that in Chen et al. (2021) and Save et al. (2016). However, considering the signal leakage from coastal ice sheets to the buffer zone of AIS (coastal ocean within 600 km, about twice GRACE natural resolution), as suggested by Save et al. (2016) the σ_i (also called RMS) over the buffer zone are all set to 4 cm for the mascons of 1° × 1° grids. If we can estimate better signal variances by using the available signals in the buffer zone, such as by comparing different ocean mass simulations from different ocean models (Uebbing et al., 2019; Dobslaw et al., 2020), the constructed regularization matrix will be more effective in correcting leakage signals.

Once **R** and α are given, the regularized solution x_{α} can be uniquely determined by minimizing the cost function (5) as

$$\boldsymbol{x}_{\alpha} = \left(\boldsymbol{A}^{T}\boldsymbol{P}\boldsymbol{A} + \alpha\boldsymbol{R}\right)^{-1}\boldsymbol{A}^{T}\boldsymbol{P}\boldsymbol{y} = \boldsymbol{Q}_{\alpha}\boldsymbol{A}^{T}\boldsymbol{P}\boldsymbol{y}$$
(7)

with $Q_{\alpha} = (N + \alpha R)^{-1}$ and $N = A^T P A$. Since a regularized solution is biased, the bias vector b_{α} is estimated with,

$$\boldsymbol{b}_{\alpha} = -\alpha \boldsymbol{Q}_{\alpha} \boldsymbol{R} \, \bar{\boldsymbol{x}} \tag{8}$$

in which \bar{x} is the vector of true parameters. For biased estimate x_{α} , its precision is evaluated by using the mean squared error (MSE) as,

$$\boldsymbol{M}(\boldsymbol{x}_{\alpha}) = \sigma_0^2 \boldsymbol{Q}_{\alpha} \boldsymbol{N} \boldsymbol{Q}_{\alpha} + \boldsymbol{b}_{\alpha} \boldsymbol{b}_{\alpha}^{T}$$
(9)

where, $M(x_{\alpha})$ consists of the covariance and bias of the regularized solution, e.g., the first and second terms of Eq. 9. Since the variance of unit weight σ_0^2 is unknown, we estimate it with the following equation,

$$\hat{\sigma}_0^2 = \frac{\hat{\boldsymbol{e}}^T \boldsymbol{P} \hat{\boldsymbol{e}} - \alpha^2 \bar{\boldsymbol{x}}^T \left(\boldsymbol{Q}_{\alpha} - \alpha \boldsymbol{Q}_{\alpha}^2 \right) \bar{\boldsymbol{x}}}{(n-t) + \alpha^2 tr \left(\boldsymbol{Q}_{\alpha}^2 \right)}$$
(10)

where \hat{e} is the residual vector, *n* and *t* are the numbers of the measurements and parameters to be estimated, respectively. Since the true value \bar{x} remains unknown, it is replaced with its estimate in Eq. 10 when calculating the variance of unit weight. The regularization parameter α plays a key role in balancing the contribution of observation errors and biases to MSE. The computation of α is commonly based on the generalized cross-validation (Golub et al., 1979), minimizing traced MSE (Hoerl and Kennard, 1970; Xu, 1998) and the L-curve method (Hansen and O'Leary 1993; Save et al., 2012), or estimated as a variance component (Koch and Kusche, 2002). In this work, the regularization parameter is updated based on the criterion of minimizing traced MSE. It ensures that the sum of the square of bias and error variance of the regularization solution is



minimized. And the true parameters in Eqs.9, 10 are replaced with their estimates (Xu, 1992; Shen et al., 2012). Since the regularization matrix in Eq. 9 is derived from filtered GRACE SHCs, it is worth to be mentioned that the regularization parameter α determined by minimizing traced MSE is based on filtered GRACE SHCs. Because the pseudo-observation vector y is generated from the unfiltered GRACE SHCs with Eq. 1, the regularized solution of Eq. 7 is looking forward to squeezing more signals.

As shown in Table 2, the global mascon solutions from GSFC and JPL are directly derived from the GRACE Level-1B data, while the CSR mascon and the regional mascon solution of this paper are derived from the GRACE SHCs, i.e., the Level-2 data. However, since the full covariance matrix of SHCs is adopted in our regional mascon method, its solution is looking forward to closing to the global mascon solution in the regional area. The advantage of our proposed method is with a lower computation burden and is easy to

TABLE 3	Description	of the	basins	over	AIS.
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Basin number	Region	Area (km²)	Length of grounding line (km)		
1	WAIS	783,290	4,521		
2	EAIS	933,754	1,343		
3	EAIS	1,615,608	1,303		
4	EAIS	329,331	2,548		
5	EAIS	238,176	1,334		
6	EAIS	693,328	3,199		
7	EAIS	501,239	3,604		
8	EAIS	159,742	1,545		
9	EAIS	166,335	1,401		
10	EAIS	943,263	154		
11	EAIS	273,145	957		
12	EAIS	773,999	3,065		
13	EAIS	1,126,542	3,302		
14	EAIS	726,359	3,688		
15	EAIS	133,755	3,533		
16	EAIS	271,666	16,278		
17	EAIS	2,100,069	3,852		
18	WAIS	411,835	1,350		
19	WAIS	481,061	1826		
20	WAIS	255,065	3,937		
21	WAIS	228,632	1,237		
22	WAIS	220,317	769		
23	WAIS	129,689	2,423		
24, 25, 26, 27	APIS	420,832	11,118		

be implemented due to using the Level-2 data. As to the regularization matrix, the same as CSR, we use filtered GRACE SHCs to estimate the prior signal variances in AIS and take a 4-cm RMS in the buffer zone, nevertheless, JPL computed the prior signal variances with the geophysical models and GSFC estimated the prior signal variances using the spatial/temporal correlated exponential function. For the regularized method, JPL used the sequential Kalman filter based on Bayesian estimation, GSFC employed the iterative solution with an empirical damping parameter, and we and CSR all used the Tikhonov regularization, but the regularization parameter was determined differently, i.e., CSR used the L-ribbon approach and we employ the minimum traced MSE criterion. Moreover, a regional optimal high-resolution solution could be combined with regional methods for inferring an empirical GIA model (Riva et al., 2009; Sasgen et al., 2017; Willen et al., 2018). As to the error assessment, GSFC used a data-driven method to evaluate the error, we use Eqs 8,9 to estimate the bias and MSE of the mascon solutions, while JPL and CSR have not presented their error assessment methods.

4 Results and discussion

4.1 Experimental design

4.1.1 Mascon set-up

The outlines of AIS and its 27 drainage basins defined by Zwally et al. (2012) are shown in Figure 2, where West Antarctic Ice Sheet (WAIS; Basins 1 and 18–23), East Antarctic Ice Sheet (EAIS; Basins 2–17) and Antarctic Peninsula Ice Sheet (APIS; Basins 24–27) are the three commonly interested regions. The description of the basins including area, number of mascon and length of grounding line is shown in Table 3. Considering more dense GRACE observations in polar regions, 2024 mascons of $1^{\circ} \times 1^{\circ}$ equal-area are parameterized covering both AIS and its 600 km buffer zone, and the mascons over AIS are also shown in Figure 2. To solve the mascons, we generate 3,616 pseudo-observations distributed with 0.75° × 0.75° equal-area at the satellite altitude covering AIS and its buffer zone. To recover the leaked signals while reducing the computational burden as possible, the buffer zone is set to twice GRACE natural resolution (i.e., 600 km).





Regularization parameter with the prior information of RMS derived from the mass change signals with Gaussian 100 km + P4M6 (left) and Gaussian 200 km + P4M6 (right) filtering.

4.1.2 Construction of regularization matrix and determination of regularization parameter

To derive the signal variance (or RMS) of each mascon, we filter the Tongji-Grace2018 model by using the P4M6 decorrelation (Swenson and Wahr, 2006) and Gaussian smoothing (Wahr et al., 1998) with a radius of 100 km and 200 km, respectively. With the filtered Tongji-Grace2018 model from April 2002 to December 2016, we compute the RMS of each mascon that is then used to construct the regularization matrix. According to Save et al. (2016), the time variability of the regularization matrix is needed only for the larger river basins, hence a constant regularization matrix is used in our regional solution over AIS with the dominant signal of ice mass loss. The spatial distributions shown in Figure 3 are the RMS values derived from 100 km to 200 km Gaussian smoothing radius over AIS and 4 cm RMS over the buffer zone (Save et al., 2016), where the RMS values over the entire AIS present strong spatial heterogeneity and exceed 60 cm in the Amundsen Sea Embayment of the West Antarctic, much larger than that in the other regions. Additionally, the RMS values at Basins 18, 13, and parts of coastal EAIS (Basin 4–9) are moderate while that of the rest is relatively smaller. Compared to 100 km Gaussian smoothing, the signals derived from 200 km Gaussian smoothing have smaller RMS and more signal leakage.

After the regularization matrix is constructed, the regularization parameter is determined by minimizing traced MSE and then the regularized solution is computed with Eq. 7. The regularization parameters of our solutions from April 2002 to December 2016 are



Region	Improvement ratio (%)				
5	100 km (%)	200 km (%)			
Basin 5–8 (DML)	21.65	38.07			
Basin 18 (Kamb)	14.07	70.41			
Basin 21and 22 (ASE)	35.29	53.81			
Basin 24–27 (APIS)	103.25	116.86			

TABLE 4 Improvements for the mass change signals.

shown in Figure 4. Since no seasonal variation exists in the regularization parameters, no unmodelled seasonal signals present in the residuals. Moreover, the regularization parameters in Figure 4 range from 10^{-8} to 10^{-6} and become larger before 2003 and after 2014 due to the worse quality of SHCs solutions.

4.1.3 Mascon estimates with the regularization matrices from different prior information

To compare the filtered estimates by P4M6 and Gaussian smoothing with 100 km and 200 km radius with the correspondent mascon estimates with the regularization matrices constructed with the filtered estimates, we present in Figures 5A–D for the time series of mass change signals of the four estimates in Dronning Maud Land (DML), Kamb, Amundsen

Sea Embayment (ASE) and APIS, where the two mascon estimates agree well and show more significant variations than the correspondent filtered estimates. After the time series of the four estimates are fitted with constant, trend, annual, semiannual and S_2 terms, the fitted signals of the mascon estimates are obviously stronger than those of the correspondent filtered estimates. The improvement ratio of the signal RMS values is defined as,

improvement ratio = $\frac{RMS(mascon) - RMS(filter)}{RMS(filter)}$ *100%

And the improvement ratios are presented in Table 4, in which the improvement ratios in ASE and APIS are 35.29% and 103.25% for the 100 km Gaussian filter, 53.81% and 116.86% for the 200 km Gaussian filter, respectively. In DML and Kamb, the improvement ratios are 21.65% and 14.07% for the 100 km Gaussian filter, and 38.07% and 70.41% for the 200 km Gaussian filter, respectively. These improvements confirm the capability of the presented mascon method in recovering the leakage signal over coastal AIS.

The mass change signals in June 2004 derived with P4M6 and Gaussian smoothing with 100 km and 200 km radius are shown in Figures 6A, B, and the correspondent mascon estimates are presented in Figures 6C, D, respectively. In Figure 6A, non-geophysical strips are still observable due to the 100 km weak filtering, while in Figure 6B 200 km moderate filtering leads to signal attenuation and leakage.



Mass change signals in June 2004. (A) and (B): P4M6 with Gaussian 100 and 200 km; (C) and (D): the correspondent mascon solutions; (E) and (F): CSR and JPL mascon solutions.

Moreover, the mascon estimates in Figures 6C, D have higher spatial resolution and less signal leakage than their counterparts in Figures 6A, B. In detail, the mass loss signals over Basin 18 are clear in two mascon estimates and the mass gain over WAIS is also concentrated on the coast of WAIS in June 2004, which means that the spatial resolution is improved via the presented mascon method. Additionally, the mass change signals of our regularized mascon solutions in Figures 6C, D agree well with that of CSR and JPL mascon solutions in Figures 6E, F, much better than their filtering counterparts in Figures 6A, B. Therefore, our mascon solutions have obvious advantages in signal recovery compared to their filtering counterparts. The slight difference between Figures 6C, D is attributed to different prior information used in the regularization matrix, which means that prior information does affect the spatial resolution of mascon estimates. Therefore, the regularization matrix should be constructed carefully especially when a high-resolution solution is needed. Comparing Figures 6C, D, we find that the mascon estimates with the prior information from 100 km Gaussian filtering have a bit higher spatial resolution and better spatial patterns relative to the mascon estimates in Figures 6E, F than that from 200 km Gaussian filtering. Hence in the following sections, we derive our mascon solutions using the regularization matrix derived from 100 km Gaussian filtering.



4.1.4 Error assessment of regularized mascon estimates

With the proposed regional mascon method, there are 157 months of mascon estimates are derived from April 2002 to December 2016, in



which 20 months are missing, since the Tongji-Grace2018 model is not available during these months. In each estimate, 2024 mascons are estimated. Since the Tikhonov regularized estimates are biased, the biases must be taken into account when assessing the precisions of the estimates. Therefore, we use the MSE, which consists of covariance and squared biases, to assess the precisions of our mascon estimates. The average values of the MSE roots and absolute biases of 2024 mascon estimates are presented in Figure 7 for 157 months in terms of Gt. The mean absolute biases of the 2024 parameters range from 0.01 to 0.06 Gt with an average of 0.02 Gt, which are much smaller than the correspondent mean MSE roots that range from 0.10 to 0.17 Gt with an average of 0.13 Gt. Hence, the biases are well controlled in our regularized mascon solutions. The MSE roots of 157 months are shown in Figure 8 in terms of EWH for 2024 mascon estimates, where there are 1813 mascons (89.6%) with the mean MSE roots less than 2 cm and 13 mascons (0.64%) over 10 cm. And the large MSE roots appear on the coastline of the ASE in West Antarctic with a maximum value of 19.04 cm, where the signal variations are also very large with a maximum of 336.21 cm.

4.2 Experimental results

4.2.1 Characteristics of mass change signals over AIS

The summed time series of mass change signals over AIS from April 2002 to December 2016 of our regularized mascon solutions are presented in Figure 9, together with that of CSR (Save et al., 2016), GSFC (Loomis et al., 2019b), JPL (Watkins et al., 2015) and mascon with identity regularization matrix, where the AIS was close to a state of balance before 2007, but with a significant mass loss after 2008, which is consistent with the finding in Shepherd et al. (2018) and Loomis et al. (2019a); Loomis et al. (2020).

According to Loomis et al. (2019a); Loomis et al. (2020), five series in Figure 9 are divided into two periods (period-1 from April 2002 to June 2007 and period-2 from July 2007 to December 2016) and then fitted with constant, trend (rate), annual, semiannual and S₂ terms. The results of mass change rate, annual, and semiannual amplitudes are presented in Table 5, in which the uncertainties are at a 95% confidence level. The mass change rate of mascon with identity regularization matrix for the period from 2002 to 2016 is only -68.3 ± 5.1 Gt/yr due to large land-ocean leakage. For the same period, the mass change rate of our mascon solution is -103.6 ± 5.6 Gt/yr, consistent with the results of three global



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	Timespan	GRACE model									
				SSFC JPI ascon		JPL mascon		Mascon with identity matrix		This study	
	04/2002-06/2007	-17.6 ± 14.6		1.7 ±	.7 ± 17.2 -2.8		± 18.0 6.6 ±		14.4	-17.4 ± 17.7	
Rate [Gt/yr]	07/2007-12/2016	-145.1 ± 8.1		-148.6	± 10.6 -142.4		± 9.1	-98.2 ± 6.9		-144.7 ± 7.3	
	04/2002-12/2016	-111.7 ± 5.9		-103.3 ± 7.4 -105.2		± 6.6	-68.3 ± 5		-103.6 ± 5.6		
	04/2002-06/2007	152.4 ± 30.4		166.5	± 35.8	109.8	± 37.5	141.1	± 30.0	149.0 ± 36.6	
Annual amplitude [Gt]	06/2007-12/2016	137.9 ± 31.3		139.1 ± 41.0 109.4		± 35.1	123.8	± 26.2	134.6 ± 28.0		
	04/2002-12/2016	144.4 ± 34.1		151.8 ± 42.7		110.4 ±38.1		131.7 ± 29.1		141 ± 32.1	
	04/2002-06/2007	60.2 ± 30.0		77.5 ± 35.3		65.5 =	± 36.9	52.5 ±	29.8	52.9 ± 36.5	
Semi-annual amplitude [Gt]	06/2007-12/2016	28.6 ± 31.3		42.4 ± 40.8 46.0		46.0	± 34.6	26.1 ±	26.3	26.8 ± 28.8	
	04/2002-12/2016	36.1 ± 33.9		51.9 ± 42.5		51.0	± 37.7	34.8 ±	29.1	35.0 ± 32.1	

TABLE 5 Rate, annual amplitude and semi-amplitude of mass change signals over AIS.

mascon solutions and -121.6 ± 14.7 Gt/yr from altimetry within the uncertainty (Shepherd et al., 2019). For the first period from April 2002 to June 2007, the mass change rate is very small and with large uncertainty, which is -17.4 ± 17.7 Gt/yr, while for the second period, the rate is -144.7 ± 7.3 Gt/yr, hence the ice mass loss became significant after 2007. As for the annual and semiannual amplitudes, the results from the five series are quite consistent within the uncertainty. From period-1 to period-2, the annual and semiannual amplitudes from our mascon solution diminished from 149.0 \pm 36.6 Gt to 134.6 \pm 28.0 Gt, and 52.94 \pm 36.5 Gt to 26.8 \pm 28.8 Gt, respectively.

The spatial patterns of mass change rates over AIS are shown in Figure 10 for period-1, period-2, and the whole study period.

4.2.2 Mass change signals in EAIS, WAIS, APIS, and 27 basins

The time series of mass change signals over EAIS, WAIS and APIS are presented in Figure 11. We can observe from Figure 11 that EAIS experienced a slow rise before 2009, followed by an intensive rise during 2009-2012 and then a slow decline after 2013. As explained in Section 4, the intensified accumulation was caused by the sharp increasing snowfall between 2009 and 2012 (Boening et al., 2012; Shepherd et al., 2019). WAIS experienced a slow mass loss before 2008, an intense loss during 2009-2014, and a slow loss again after 2014. APIS experienced a steady mass loss from the middle of 2006 to the beginning of 2016, which is consistent with that in Figure 10 where the pattern over APIS in period-2 is significant while that in period-1 is not obvious. Overall, the three regions have quite different mass change characteristics. The estimated mass change rates over EAIS, WAIS and APIS from April 2002 to December 2016 are shown in Table 6, in which our solutions are consistent with CSR, GSFC, and JPL solutions within the uncertainties.

The mass change rates and annual amplitude at the basin scale are also estimated from the $1^{\circ} \times 1^{\circ}$ equal-area mascon estimates by

the least-square fitting, and the results are shown in Figure 12, where APIS is the summed area of four small basins, that is Basins 24, 25, 26, and 27. The large mass gains are in Basins 7 and 18, with rates of 18.03 \pm 1.88 Gt/yr and 14.55 \pm 0.60 Gt/yr, while the high mass losses are in Basins 21 and 22, with rates of -58.57 ± 2.48 Gt/ yr and -44.12 ± 2.27 Gt/yr. The rates in most basins of EAIS show good consistency between ours and those of CSR, GSFC and JPL. However, for the basins in WAIS (Basins 1, 20, 21, 22, and 23), the discrepancies in mass change rates between the four mascon solutions are relatively large. Especially for the rates in Basins 18, 20, and 21, the GSFC solution does not agree with the other three solutions within the error bars of 95% confidence level. The mass change rate from the JPL solution in basin 1 also does not agree with the other three solutions within the error bars. However, our regional solution agrees well with the CSR solution except for Basin 23, and in Basin 23 our solution agrees with the JPL solution. The estimated annual amplitudes agree well between the four mascon solutions within the error bars, except for Basins 3 and 18, in which, however, our solution still agrees with the CSR solution. The probable reason for the better agreement is that we use the same method as CSR to construct the regularization matrix.

4.3 Discussion

4.3.1 Spatial pattern comparison of mass change rate over AIS

As shown in Figure 10, compared to the results derived from P4M6 + Gaussian 200 km and the mascon solution with identity regularization matrix (the first and second rows of Figure 10), the mass change rates of our mascon solutions are more clear with stronger mass loss and gain signals. The maximum mass gain rate over the coast of Basin 7 in period-2 is over 16 cm/yr, while that of



the filtered one and the mascon solution with identity regularization matrix is only 4 cm/yr and 8 cm/yr. Moreover, the mass loss rate in Totten glacier (Basin13) for period-2 is very significant, with an extreme value of over -16 cm/yr, while that of the filtered one and the mascon solution with identity

regularization matrix is only -6 cm/yr and -7 cm/yr. Therefore, our mascon solutions can squeeze more signals than the correspondent filtering ones and the mascon solution with an identity regularization matrix. In general, the three global mascon solutions have more similar spatial patterns to our



TABLE 6 Mass change rate from April 2002 to December 2016.

Regions	CSR mascon	GSFC mascon	JPL mascon	This study
WAIS	-145.4 ± 4.9	-150.0 ± 5.5	-148.2 ± 5.5	-143.3 ± 4.9
EAIS	60.0 ± 4.2	67.3 ± 5.1	69.6 ± 4.5	63.0 ± 4.3
APIS	-26.32 ± 1.3	-20.54 ± 1.3	-26.58 ± 1.4	-23.29 ± 1.2

mascon solutions than the filtering ones. The mass change signals of our mascon solutions are also more coincident with the shape of glaciers and ice streams than that of the filtering ones, such as the most striking ice mass loss in Getz (Basin 20), Thwaites (Basin 21) and Pine Island (Basin 22), and the ice mass gain in Kamb Ice Stream (Basin 18), which are consistent with the findings in Rignot et al. (2011); Shepherd et al. (2019). Therefore, our mascon estimates with the higher spatial resolution are more reliable than the correspondent filtered estimates and the mascon solution with identity regularization matrix. Compared to GSFC and JPL mascon solutions, the mass gain rates in Basin 18 from our mascon and CSR mascon solutions are more coincident with that of the Kamb Ice Stream. And the mass loss rate in Totten glacier of Basin 13 is as well. As shown in the first two columns of Figure 10, the spatial patterns in the two periods are obviously different, the mass loss area in period-2 is much larger than that in period-1, especially in the coastline of the ASE of West Antarctic, which is consistent with the finding in Harig and Simons (2015). The mass loss rates in Totten and Moscow (basin 13), as well as in basin 15, are also enhanced. Interestingly, some regions have opposite mass change rates in the two periods. For instance, the western Dronning Maud Land (basins 5-6) experienced ice mass loss in period-1 and

mass gain in period-2, which is consistent with Shepherd et al. (2019) who derived the results from satellite altimetry. In the whole Dronning Maud Land (basins 5–8), a broad pattern of modest ice sheet mass gain spans much of the coastline and stretches inland for several kilometers in the whole study period, which is associated with sharp increases in snowfall that happened from 2009 to 2012 (Boening et al., 2012; Shepherd et al., 2019).

4.3.2 Annual and interannual signals comparison with RACMO

To compare the annual and interannual signals of the four mascon solutions with the cumulated SMB from RACMO, the detrended and smoothed (with a 7-month moving average filter) time series of the mass change signals in EAIS, WAIS and APIS are shown in Figure 13 together with that from RACOM. The estimated RMSE (root mean squared error) values from the differences between the four mascon solutions relative to the cumulated SMB in Figure 13 are illustrated in Table 7 together with correlation coefficients between the time series of cumulated SMB and the four mascon solutions. Except for the GSFC solution in EAIS, all the correlation coefficients and RMSE values in Table 7 are close to each other and the correlation coefficients are at least



TABLE 7 Correlation coefficients and RMSE of detrended and smoothed mass change from four mascon solutions and RACMO.

Region	CSI	R mascon	GSF	C mascon	JPL mascon		This study		
	Cor	RMSE [Gt]	Cor	RMSE [Gt]	Cor	RMSE [Gt]	Cor	RMSE [Gt]	
EAIS	0.91	42.92	0.87	62.15	0.94	44.79	0.91	42.18	
WAIS	0.94	15.68	0.92	18.92	0.94	16.66	0.94	16.20	
APIS	0.96	10.31	0.96	12.08	0.96	9.86	0.96	11.49	

equal to 0.91. The four mascon solutions have the same correlation coefficients in WAIS and APIS, however, the JPL solution has the highest correlation coefficient in EAIS, and our solution has the same correlation coefficient as the CSR solution. As for the RMSE values, our solution has the smallest value in EAIS, the CSR solution has the smallest value in WAIS and JPL solution is the smallest in APIS.

After detrended and smoothed, more obvious annual variations can be observed in Figure 13 than that in Figure 11, and the discrepancies between the four mascon solutions are obviously smaller than that relative to cumulated SMB. By the way, due to the enhanced precipitation during the extreme 2015–2016 EL Niño (Bodart and Bingham, 2019), a significant mass gain in APIS and WAIS can be observed by the mascon solutions.



5 Conclusion

This contribution proposed the gravitational potential-based regional mascon method, in which the pseudo-observations of gravitational potential are generated from unfiltered GRACE level-2 data, and the regularization matrix is constructed with the prior information derived from the GRACE level-2 data with two-step filtering of 100 km radius. With the regional mascon method, the $1^{\circ} \times 1^{\circ}$ equal-area mascon solutions over AIS are derived from the Tongji-Grace2018 model from April 2002 to December 2016. The results demonstrate that our regional mascon solutions can achieve a better spatial resolution, and effectively reduce the land-ocean signal leakage and the signal leakage between the adjacent mascons.

The mass change signals of our mascon solutions have been significantly enhanced relative to the filtering counterparts with the improvement ratios of 21.65%, 14.07%, 35.29%, and 103.25% in DML, Kamb, ASE and APIS, which confirmed the capability of the presented regional mascon method in recovering the leakage signal over coastal AIS. The mass change rates over AIS of our mascon solution are -103.6 ± 5.6 Gt/yr from 2002 to 2016, -17.4 ± 17.7 Gt/yr during 2002–2007, and -144.7 ± 7.3 Gt/yr from 2008 to 2016, the ice mass loss is significantly intensified after 2007. In EAIS, WAIS, and APIS, the mass change rates from 2002 to 2016 are quite different, with the rates of 63.0 ± 4.3 Gt/yr, -143.3 ± 4.9 Gt/yr and -23.29 ± 1.2 Gt/yr respectively. The mass change signals at the basin scale with even more distinguishing features, significant mass gain occurred in Basins 7 and 18, with the rates of 18.03 ± 1.88 Gt/yr and 14.55 ± 0.60 Gt/yr, while large mass loss presented in Basins 21 and 22, with the rates of -58.57 ± 2.48 Gt/yr and -44.12 ± 2.27 Gt/yr. Relative to the cumulated SMB from RACMO, the correlation coefficients of four mascon solutions are at least equal to 0.91 except for the GSFC solution in EAIS.

Since the pseudo-observations are generated from GRACE level-2 data, the proposed regional mascon method has a lower computation burden and can be more easily optimized than the present global mascon methods. Moreover, the regional mascon method can achieve

much better results than the filtered counterparts and its solutions are as good as the global mascon solutions, although the regularization matrix is simply constructed with the prior information from the filtered GRACE level-2 data in AIS and with the constant variance in the buffer zone. If the additional information from high-resolution global ocean and sea ice data synthesis, i.e., ECCO2 (Estimating the Circulation and Climate of the Ocean) or sea level change from satellite altimetry can be used to estimate the signal variances in the buffer zone, and the information from InSAR (Interferometric synthetic aperture radar) or ICESat (Ice, Cloud, and land Elevation Satellite) is used to refine the signal variances in AIS, the spatial resolution and the accuracy of the mascon solution are looking forwards to be further increased.

Data availability statement

The original contributions presented in the study are included in the article/supplementary material, further inquiries can be directed to the corresponding author.

Author contributions

WW performed the data processing, analyzed the experimental results, and drafted the manuscript. YS performed the methodology research, designed the study, conducted the analysis of the results and revised the manuscript. QC and TC checked the performance of this method and revised the manuscript. All authors read and approved the final manuscript.

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Conflict of interest

The authors declare that the research was conducted in the absence of any commercial or financial relationships that could be construed as a potential conflict of interest.

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