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Light hydrocarbons composition and carbon isotopes unravel the origin of natural gas in Triassic marine strata of Moxi gas field, central Sichuan Basin

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The Triassic marine gas reservoir in the Moxi Gas Field is one of the large gas fields initially discovered in the Sichuan Basin. However, the origin of its natural gas has long been debated. To study the genetic types and origin of natural gas, the light hydrocarbons (LHs) composition, carbon isotopes, chemical compositions, and stable carbon and hydrogen isotopic compositions were analyzed in the Triassic strata of Moxi gas field in the Sichuan Basin. The natural gases from the second member of the Jialingjiang Formation $(T_2)^2$ and the first member of the Leikoupo Formation (T_2l^1) have dryness coefficients higher than 0.997 and contain low content non-alkane gases such as CO_2 and N_2 . The $\delta^{13}C_1$ values of natural gas in the Moxi (MX) gas field range from -35.3% to -32.9%, the $\delta^{13}C_2$ values range from –34.0‰ to –30.3‰, and the δD_1 values range from –144‰ to -120%. As for light hydrocarbons composition, the C₇ LHs are dominated by the methylcyclohexane (MCH), and the C_6-C_7 LHs have low normal alkane contents. The individual carbon isotopes of light hydrocarbons have the characteristics of δ^{13} C > -28‰. The origin identification indicates that the natural gas in this area is oil-type gas that results from the secondary cracking of crude oil. Our results show that the natural gas in $T_2 j^2$ of the MX gas field is sourced from the shale gas of the $\mathcal{E}_1 q$, the $T_2 l^1$ gas is a mixed-source gas from different thermal evolution stages of the $\mathcal{E}_1 q$ source rocks, shale gas and oil-cracking gas. The mixing effect is the main reason for the reversed phenomenon ($\delta^{13}C_1 > \delta^{13}C_2$).

KEYWORDS

Moxi gas field, light hydrocarbons, shale gas, genetic types, gas source rock

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Highlights

- 1. The $T_2 l^1$ gas is a mixed-source gas from the shale gas and oil-cracking gas of the $\varepsilon_1 q$ Fm.
- 2. The mixing effect gives $T_2 l^1 \operatorname{gas} \delta^{13} C_1 > \delta^{13} C_2$.
- 3. The dominance of MCH relative to C_7 LHs, $\delta^{13}C_{CH} > -24\%$ and $\delta^{13}C_{MCP} < \delta^{13}C_{CH}$ distinguish shale gas from oil–cracking gas.

1 Introduction

The MX gas field is located in the southern part of the central Sichuan low–flat tectonic area. The gas–bearing structure is at the top of the eastern segment of the Leshan–Longnyusi Caledonian Paleo–uplift. From bottom to top, there are three pay zones, namely, the gas reservoir in the T_2j^2 and T_2l^1 of the marine strata, and the gas reservoir in the Xujiahe Formation (T_2x) of the continental strata. This gas field spatially overlays the Anyue gas field, a super–large gas field in the Sinian–Cambrian System in central Sichuan. The gas–bearing range of the MX gas field is smaller than that of the Anyue Gas Field. The Leshan–Longnyusi Paleo–uplift developed in the Gaoshiti–Moxi area is the dominant direction for the long–term accumulation of oil and gas migration.

The MX gas field is a large scale gas field discovered relatively early in the Sichuan Basin. The $T_2 l^1$ gas reservoir was proven in 1988, and it took 26 years for the $T_2 j^2$ gas reservoir to be proven (Wang et al., 1998). However, the gas source has been less discussed and exists significant dispute over the source of the gas. Some scholars believe that the gas in $T_2 l^1$ sourced from the coal measures of the Longtan Formation (P2l) (Wang et al., 1998; Li et al., 2006), and the gas in $T_2 j^2$ is contributed by the Permian carbonate source rocks (Li et al., 2006). However, the burial depth of the Triassic gas reservoir is only more than 2,600 m. The inclusion detection show that the temperature of the fluids trapped in the reservoir is 111-165°C, which has just reached the stage of oil-cracking gas generation, corresponding to the condensate oil and gas generation (Jin et al., 2019). But the gas dryness coefficient is extremely high, with $\delta^{13}C_1$ being heavy while $\delta^{13}C_2$ is light, which is obviously inconsistent with the thermal evolution history of the Permian-Triassic source rocks such as the P_2l and the third member of Leikoupo Formation (T_2l^3) in the central region Sichuan Basin (CRSB) (Yu, 2022). In prior research regarding the gas-source identification in the CRSB, the emphasis has been placed on the utilization of natural gas compositions, carbon and hydrogen isotopes of gas components, and biomarker indexes of reservoir bitumen. Nevertheless, the indicator parameters of the LHs associated with natural gas are yet to be researched.

Although the content of LHs compounds associated with natural gas is low, their geochemical characteristics play an important role in determining the genetic types of natural gas, gas–source correlation, and secondary transformation (Thompson, 1983; Hu et al., 2010). This paper provides data on the LHs components and its individual carbon isotopes. Combining with the data of natural gas components and the carbon and hydrogen isotopes of each component, a systematic analysis of the natural gas characteristics of $T_2 l^1$ and $T_2 j^2$ in the Triassic System of the MX gas field was carried out. The origin of natural gas in the marine carbonate rock strata of the MX gas field

in the CRSB were discussed, providing a basis for the subsequent exploration of the marine natural gas-rich area in the MX region.

2 Geological setting

The Sichuan Basin is located on the northwestern margin of the Yangtze Platform. It is a superimposed basin that has experienced multiple tectonic movements. The multiple tectonic-sedimentary cycles are related to the opening and closing evolution of the Tethys Ocean. The differentiation period of the marine craton can be divided into three stages, the Sinian-Silurian is the development period of the marine craton fault-depression basin (Ma et al., 2018), the Carboniferous-Middle Triassic is the period of the marine craton depression basin, and the Late Triassic-Cenozoic is the foreland basin period. During the multi-stage tectonic evolution, multiple marine strata have developed, including multiple source rock strata. From the oldest to the youngest, the strata are the Lower Sinian Doushantuo Formation $(Z_1 ds)$, the Lower Cambrian Qiongzhusi Formation $(\mathcal{E}_1 q)$, the Upper Ordovician Wufeng Formation-Lower Silurian Longmaxi Formation $(O_1 w - S_1 l)$, the Middle Permian Qixia Formation–Maokou Formation (P_2q-P_2m) , the Upper Permian Longtan Formation (Wujiaping Formation) and Dalong Formation (P_3l-P_3d) (Figure 1). The vertical distribution of conventional marine carbonate gas reservoirs is mainly controlled by the source rock strata, forming multiple composite petroleum systems.

At the same time, during the multi-stage tectonic evolution, the tectonic characteristics vary in different blocks of the basin. According to their characteristics and the distribution of oil and gas, six tectonic units can be divided: the high-steep tectonic area in eastern Sichuan, the gentle tectonic area in central Sichuan, the low-steep tectonic area in southern and southwestern Sichuan, the western Sichuan sag tectonic area, and the low-flat tectonic area in northern Sichuan (Jiao et al., 2022). However, in the entire Sichuan Basin, the Ordovician to Carboniferous strata are not well-developed. An approximately east-swest-trending extensional fault-depression basin developed, with a cumulative sediment thickness of less than 200 m. In most areas, the Lower Cambrian and Upper Permian sediments are in direct contact. The Silurian is only distributed in the eastern Sichuan area and gradually pinches out in the CRSB (Figure 2B), with only remnants on the edge of the study area. The Dalong Formation (P_3d) is distributed in the Kaijiang-Liangping trough in the northeastern Sichuan area and contributes little to the CRSB. The thickness of the Sinian Doushantuo Formation $(Z_1 ds)$ is 10-30 m, and the thickest part is developed in the area of Quang An-Nanchong-Suining. The P_2q-P_2m source rocks are carbonate rocks with low TOC contents, belonging to secondary source rocks. The Gufeng member of the Maokou Formation has an extremely high TOC (0.5%-31.04%, average = 5.08%, Fu et al., 2021), but it is distributed in the northern part of the basin. Therefore, in the CRSB, the $\in_1 q$, $P_3 l$, and $T_2 l^3$ are the three sets of source rocks that mainly contribute to the natural gas in the Triassic marine strata.

The gentle tectonic area in CRSB ranges from Yanting in the north to Neijiang–Leshan in the south. The MX Structure, at the high–position of the eastern segment of the Leshan–Longnyusi Paleo–uplift, is the prime target for hydrocarbon accumulation.



In the Late Sinian, the paleo-uplift began to take shape, while the Cambrian-Ordovician strata were not well- developed. By the late Caledonian at the end of the Silurian, after regional uplift and subsidence, it was finally formed. During the Indosinian–Himalayan movement, the rapid rise of the Longmen Mountains and the subsequent squeezing transformed the tectonic



pattern and traps. Currently, the Weiyuan gas field (WY) is in the western part of the paleo–uplift's top, and the Anyue (AY) and MX gas fields are in the east. In the CRSB, large–scale strike–slip fault developed in the Sinian–Permian strata, experienced two–stage activities (Ma et al., 2018; Jiao et al., 2022), have laid the groundwork for 3D reservoir exploration (Figure 2C).

3 Samples and experiments

3.1 Samples

A total of 20 natural gas samples were collected. 10 samples were from the $T_2 l^1$ and five samples were from the $T_2 j^2$ from the Triassic marine carbonate gas reservoirs of the MX gas filed in the CRSB. For comparison, five samples from the $\mathcal{E}_1 l$ in the MX structure of Anyue gas filed were also collected. Prior to sampling, a double valve aluminum alloy cylinder (with a volume of 2L) capable of withstanding a gas pressure of 20 Mpa was utilized. The pressure gauge was removed at the wellhead, and the gas cylinder was connected to the wellhead *via* a hose, both gas valves were opened to flush the cylinder for 3–5 min. Subsequently, one of the

valves was closed to fill the cylinder with gas at a pressure equal to that of the wellhead. During sampling, wells with a pressure ranging from 3 to 10 Mpa and that had been operating normally for an extended period were selected. This was to ensure the success of the light hydrocarbon enrichment experiment.

3.2 Analytical methods

3.2.1 Natural gas analysis

The compositions of natural gas alkane and non–alkane gases, CH_4 and C_2H_6 carbon isotopes, CH_4 hydrogen isotope, and LHs (C_6-C_7) molecular composition were measured using industry–standard techniques at the PetroChina Research Institute of Petroleum Exploration and Development (Beijing). Gas components were detected by an Agilent 7890B gas chromatograph (GC) with five valves, seven columns, and three detectors (one FID and two TCDs). The FID quantified alkane gas proportions; the first TCD measured N₂, CO₂, and H₂S proportions, with helium as the carrier gas for both. The second TCD, using N₂ as the carrier gas, measured only the He proportion. Carbon isotope compositions were analyzed online with a MAT 253 gas isotope mass spectrometer.

Gaseous hydrocarbons, separated by a PQ capillary column, were transferred by the carrier gas (He) to a combustion furnace, oxidized to CO₂ by CuO at 850°C, and then the converted products were carried to the mass spectrometer for isotope measurement. δ^{13} C values were calibrated against the Vienna Peedee Belemnite (VPDB) standard, and δ D values against the Vienna Standard Mean Ocean Water (VSMOW) standard. Before daily sample analysis, standard samples were measured to ensure instrument stability. The analysis errors for components and δ^{13} C were less than 0.3% and 0.5‰ respectively, and that for δ D was less than 5‰.

3.2.2 Light hydrocarbon analysis

The analysis of LHs components was conducted using an Agilent 7890A gas chromatograph (GC) outfitted with a PONA capillary column. At the injection port, the PONA column was coiled repeatedly to create an enrichment loop. Prior to sample injection, this loop was submerged in a liquid–nitrogen cold trap. Ten minutes subsequent to injection, the cold trap was withdrawn, and upon closing the column oven, the analysis was initiated. The column oven temperature was maintained at 30°C for 15 min, then gradually increased to 70°C at a rate of 1.5°C/min, and further elevated to 280°C at a rate of 2.5°C/min. The carbon isotope ratios of LHs (C_6 – C_7) were determined *via* GC–C–IRMS, employing the identical temperature programming as that utilized for the determination of LHs components. The detection precisions for LHs components and δ^{13} C were ±0.3% and ±0.5‰, respectively.

4 Results

4.1 Chemical composition of natural gas

The molecular compositions and carbon isotopes of 20 gas samples from the Triassic gas reservoirs $(T_2 j^2 \text{ and } T_2 l^1)$ in MX gas field and the $\in_1 l$ gas samples in MX structure of AY gas field are presented in Table 1. The gas is extremely dry, with a dryness coefficient $(C_1/C_n) > 0.99$. The gas is predominantly composed of alkane gases, with CH4 being in an absolutely dominant position, ranging from 96.83% to 99.36% with an average of 96.96%. The C_2H_6 is extremely low (0.2%-0.26%), and the C_3H_8 is lower than that of C_2H_6 (0.01%–0.03%). In certain samples, the C_3H_8 content fell beneath the instrument's detection threshold, precluding the measurement of corresponding values. The content of non-alkane gases is low. The N_2 content ranges from 0.72% to 2.43%, with an average of 1.14%. The CO2 content is unevenly distributed, ranging from 0.07% to 2.18%, with an average of 0.29%. The H_2S content also varies considerably, ranging from 0% to 2.94%, with an average of 1.36%.

4.2 Stable carbon and hydrogen isotopic composition

Both $\delta^{13}C_1$ and $\delta^{13}C_2$ values show narrow distribution ranges: $\delta^{13}C_1$ ranges from -35.3‰ to -31.4‰ (average = -33.6‰), and $\delta^{13}C_2$ from -34.0‰ to -28.6‰ (average = -32.8‰). Owing to the relatively heavy $\delta^{13}C_1$ values, carbon isotope reversal ($\delta^{13}C_1 >$ $\delta^{13}C_2$) occurs in M5, M030–H13 and M91 samples. δD_1 values are generally heavy, varying narrowly from –144‰ to –120‰, with an average of –130‰.

4.3 Molecular composition of $C_6 - C_7$ light hydrocarbons

Table 2 shows the molecular composition and distribution of LHs associated with gas in the MX gas field and MX structure of AY gas field of the CRSB. In the samples, LHs components are characterized by high average contents of cycloalkanes (37.8%) and iso–alkanes (39.5%), while normal alkanes are relatively low, with a maximum of 35% of total C₆–C₇ LHs and an average of 22.6%. For C₇ LHs components, methylcyclohexane (MCH) has the highest average content (55.8%), followed by n–heptane (nC₇, 31.9%) and dimethylcyclopentanes (Σ DMCPs, 14.1%). The Mango coefficient K₁ ratio [(2–MH + 2.3–DMP)/(3–MH + 2.4–DMP)] of T₂*j*² gas ranges from 1.02 to 1.17, whereas T₂*l*¹ shows a wider range, from 0.96 to 1.77. In the C₁*l* Fm, low iso–alkane contents precluded K₁ calculation for most samples; the average K₁ of samples M201–1 and M201–2 is 1.16.

4.4 Isotopic characteristics of C₆–C₇ light hydrocarbons

As the gases are over-mature gas and the concentrations of LHs are low, reliable individual carbon isotope data were obtained for only six samples. Meanwhile, over-mature natural gas also suffers from poor separation between adjacent LHs compounds, which affects the detection results of individual carbon isotope values of LHs. The carbon isotope values of seven LHs compounds, namely, 2-MP, 3-MP, nC₆, MCP, CH, nC₇, and MCH, were selected for comparison, as shown in Table 2. Among the seven LHs compounds, $\delta^{13}C_{MCH}$ is the heaviest, with $\delta^{13}C$ values ranging from -23.5‰ to -21.3‰ and an average of -22.8‰. The carbon isotope values of normal alkanes, $\delta^{13}C_{nC6}$ and $\delta^{13}C_{nC7}$, are the lightest, with averages of -26.7‰ and -27.0‰ respectively. CH and MCP have the widest distributions of carbon isotope values, showing significant differences among different samples. The $\delta^{13}C_{CH}$ values range from -26.8% to -22.3%, with an average of -24.4%, and the $\delta^{13}C_{MCP}$ values range from -26.3% to -23.9%, with an average of -24.9%.

Due to the samples being over – mature gas with low LHs concentrations, reliable individual carbon isotope data were available for only six samples. Moreover, over – mature natural gas exhibits poor separation between adjacent LHs compounds, which impacts the carbon isotope detection of individual LHs compounds. As presented in Table 2, the carbon isotope values of seven LHs compounds (2–MP, 3–MP, nC₆, MCP, CH, nC₇, and MCH) were selected for comparison. Among them, ¹³C_{MCH} is the heaviest, with values from –23.5‰ to –21.3‰ (average = –22.8‰). In contrast, the carbon isotope values of normal alkanes ($\delta^{13}C_{nC6}$ and $\delta^{13}C_{nC7}$) are the lightest, averaging –26.7‰ and –26.9‰ respectively. CH and MCP show the widest carbon – isotope value distributions, varying significantly across samples. The $\delta^{13}C_{MCP}$ values from –26.8‰ to –22.3‰ (average = –24.4‰), and $\delta^{13}C_{MCP}$ values from –26.3‰ to –24.0‰ (average = –24.9‰).

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TABLE 1	

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servoirs in the MX and AY gas fields.	mical Coposition (vol%) Dryness Coefficient δ^{13} C $\delta D (\%,VSMOW)$ ($\%,VPDB$)	C ₃₊ N ₂ CO ₂ H ₂ S CH ₄ C ₂ H ₆ CH ₄	0.01 0.52 0.00 0.00 0.979 -33.9 -34.1 -140	0.01 0.45 0.00 0.00 0.980 -34.6 -34.1 -139	0.01 0.72 2.18 0.06 0.978 -34.2 -33.6 -122	0.01 2.43 0.31 0.00 0.978 -33.2 -31.9 -	0.03 2.22 0.40 0.05 0.970 -33.1 -33.8 -	0.01 1.02 0.13 1.43 0.9980 -32.9 -33.9 -135	0.01 1.05 0.12 1.30 0.976 -35.3132	0.01 1.00 0.14 1.21 0.979 -33.5 -138	0.01 1.03 0.13 1.45 0.9980 -33.5 -32.6 -138	0.01 1.18 0.17 2.56 0.978 -34.9 -32.9 -144	0.02 1.13 0.15 2.16 0.978 -34.9 -33.5 -144	0.00 1.03 0.12 2.03 0.9980 -32.9 -33.1 -123	0.00 1.23 0.16 2.94 0.9980 -33.8 -33.2 -124	0.00 1.05 0.14 2.12 0.9980 -33.6 -124	0.00 1.19 0.15 2.17 0.975 -33.5 -120	0.01 0.80 2.83 0.32 0.9985 -33.1 -33.0 -133	0.01 0.73 2.53 0.62 0.9985 -33.4 -134	0.01 0.78 2.07 0.50 0.9984 -33.0 -32.8 -137	0.01 0.69 1.76 0.60 0.985 -32.4 -32.3 -133	0.00 0.62 -33.5 -33.3 -122
	Dryness C	S	.6.0	0.9	16 0.9	0.0	15 0.9	3 0.9	0.0	0.9	.5 0.9	.0	6.0	13 0.9	14 0.9	2 0.9	.0.0	.2	0.9	0.9	0.0	.6.0 0.
as fields.	<u>%</u>	² H ₂	0.0	0.0	3 0.0	0.0	0.0	1.4	. 1.3	۱.2 I.2	1.4	7 2.5	5 2.1	2.0	5 2.9	t 2.1	5 2.1	\$ 0.3	3.0.6	7 0.5	0.6	3 0.7
and AY g	on (vol?	S	00.00	00.00	2.18	0.31	0.40	0.13	0.12	0.14	0.13	0.17	0.15	0.12	0.16	0.14	0.15	2.83	2.53	2.07	1.76	2.08
in the MX	opositi	N_2	0.52	0.45	0.72	2.43	2.22	1.02	1.05	1.00	1.03	1.18	1.13	1.03	1.23	1.05	1.19	0.80	0.73	0.78	0.69	0.62
eservoirs	mical C	C ₃₊	0.01	0.01	0.01	0.01	0.03	0.01	0.01	0.01	0.01	0.01	0.02	0.00	0.00	0.00	0.00	0.01	0.01	0.01	0.01	0.00
mbrian re	Che	C ₂	0.20	0.19	0.20	0.20	0.26	0.18	0.23	0.20	0.19	0.20	0.19	0.19	0.19	0.20	0.24	0.13	0.13	0.14	0.14	0.14
sic and Ca		5	99.27	99.35	96.83	97.06	97.04	97.23	97.29	97.45	97.20	95.89	96.34	96.62	95.46	96.48	96.23	95.91	95.98	96.50	96.80	96.45
of gas from Triass	Depth/m		3,190–3,192	3,147-3,160	3,128-3,180	3,128-3,180	3,079-3,132	3,722	3,670	3,589	3,612	3,717	3,270	2,676	2,723	3,200	3,431	4,575	4,640	4,750	4,660	4,574-4,608
able isotopic compositions	Gas pool									MX gas field										AY gas field (MX structure)		
lar and st	E		$T_{2}j^{2}$	$T_{2}j^{2}$	$T_{2j^{2}}$	T_{2j^2}	$T_{2}j^{2}$	$T_2 l^1$	$T_2 l^1$	T_2l^1	T_2l^1	T_2l^1	$T_2 l^1$	$T_2 l^1$	$T_2 l^1$	T_2l^1	T_2l^1	$\epsilon_1 l$	$\epsilon_1 l$	$\epsilon_1 l$	$\epsilon_{_{\rm I}}l$	$\epsilon_{_{1}l}$
TABLE 1 Molecu	Well		M005-H9	M157	MS1	MS1-2	M5	M030-H13	M030-H14	M030-H16	M030-H17	M004-H9	M004-H10	M91	M89	M017-H3	M032-H2	M201-1	M12	M009-X1	M8	M201-2

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Well	E	Gas pool	Depth/m	C ₆₋₇ a area i	llkanes/GC (%)		C ₇ /GC	area ((%	(2–MH+2.3 –DMP)/nC ₇	(3–MH+2.4 –DMP)/nC ₇	, К	δ ¹³ C (‰	,VPDB)					
				lso	Normal	Cyc	MCH	nC_{7}	ZDMCPs				2–MP	3–MP	nC ₆	MCP	СН	nC_7	MCH
M005-H9	$T_{2}j^{2}$		3,190-3,192	21.53	9.59	68.88	28.01	61.99	10.00	1.03	0.93	1.11						1	
M157	$T_{2}j^{2}$		3,147-3,160	40.05	10.29	49.66	38.93	38.40	22.67	1.37	1.33	1.02	I	I		I	I	1	I
MS1	T_{2j}		3,128-3,180	31.32	20.97	47.71	47.98	43.77	8.25	0.69	0.58	1.17	-27.8	-28.1	-28.4	-26.8	-24.2	-29.1	-23.5
MS1-2	T_{2j}^{2}		3,128-3,180	33.38	22.29	44.32	51.13	40.70	8.16	0.65	0.59	1.11							
M5	T_{2j} ²		3,079-3,132	38.76	9.26	51.98	35.74	38.68	25.59	2.06	1.37	1.50						1	
M030-	$T_2 l^1$		3,722	52.39	17.05	30.56	61.37	27.00	11.63	1.31	1.37	0.96					1	1	
M030-H14	$T_2 l^1$		3,670	42.19	24.54	33.27	62.15	30.99	6.86	0.77	0.69	1.11							
M030-H16	$T_2 l^1$	MX gas field	3,589	44.30	32.98	22.72	69.55	20.45	10.00	0.59	0.54	1.10							
M030-H17	$T_2 l^1$		3,612	36.86	33.13	30.01	57.60	24.70	17.70	0.44	0.43	1.02						1	
M004-H9	$T_2 l^1$		3,717	42.07	33.01	24.92	66.32	22.01	11.66	0.63	0.43	1.45							
M004-H10	T_2l^1		3,270	40.38	34.18	25.44	66.03	22.53	11.44	0.57	0.41	1.38					I		
M91	$T_2 l^1$		2,676	33.34	22.25	44.41	51.06	40.79	8.15	0.65	0.59	1.11	-24.1	-24.6	-26.6	-24.2		-26.4	-21.3
M89	$T_2 l^1$		2,723	51.30	33.46	15.24	76.25	13.71	10.03	0.65	0.48	1.36	-24.5	-24.6	-27.0		-24.8		I
M017-H3	T_2l^1		3,200	56.79	30.11	13.10	77.83	11.73	10.43	0.75	0.61	1.24	-22.2	-21.9	-24.8	-22.3	-25.0	-26.4	-23.5
M032-H2	T_2l^1		3,431	46.62	35.79	17.59	73.93	15.78	10.29	0.58	0.40	1.45	-24.3	-24.7	-27.0	-23.7	-26.3	-25.6	-23.4
																(Co	ntinued or	the follov	/ing page)

		MCH	I	I	I	I
		nC_{7}				I
		CH	I	I	I	I
		MCP	I			I
		nC ₆	I	I	I	I
	°,VPDB)	3–MP	I	I	I	I
elds.	δ ¹³ C (%	2–MP	I	I	I	I
vY gas h	\mathbf{A}_1		1.05	0.77	0.53	0.48
om the MX and P	(3-MH+2.4 -DMP)/nC ₇		0.42	1.02	1.94	2.65
th natural gases fr	(2–MH+2.3 –DMP)/nC ₇		0.44	0.79	1.03	1.28
C ₆ -C ₇ LHS WI	(%	ΣDMCPs	15.07	5.16	5.36	21.82
ratios ol	area (nC_7	30.80	42.41	41.01	41.59
d isotope	C ₇ /GC	MCH	54.13	52.43	53.63	36.60
ters, and		Cyc	36.26	44.72	43.33	53.19
nent parame	llkanes/G((%)	Normal	32.08	13.29	11.30	8.78
i compo	C ₆₋₇ a area (lso	31.66	41.99	45.37	38.03
ontents, mair	Depth/m		4,575	4,640	4,750	4,660
ed) Kelative c	Gas pool				AY gas held (MX	structure)
ontinue	E L		$\epsilon_1 l$	$\epsilon_1 l$	$\epsilon_1 l$	$\epsilon_1 l$
ABLE Z (C	Well		M201-1	M12	M009-X1	M8

2-MH, 2-methylpentane; 3-MH, 3-methylpentane; 2.3-DMR, 2.3-dimethylpentane; 2.4-DMR, 2.4-dimethylpentane; nC, hexane; MCR, methylcyclopentane; CH, cyclohexane; nC, hexane; MCH $^{**}K_1 = (2-MH + 2.3-DMP)/(3-MH + 2.4-DMP).$

-23.4

-27.3

-23.9

-25.5

-26.7

-24.7

-24.6

1.27

0.47

0.60

11.98

34.27

53.76

38.93

24.16

36.91

4,574-4,608

 $\epsilon_1 l$

M201-2

"represent "not available".

5.1 Genetic identification of gas

On the $\delta^{13}C_1$ vs C_1/C_2+C_3) gas genetic diagram proposed by (Milkov et al., 2020) and the modified Bernard diagram (Figure 3A; Bernard et al., 1978), the gas samples mainly represent late-mature thermogenic gases. The $\delta^{13}C_{2-1}$ vs $\delta^{13}C_1$ plot (Figure 3B) also shows the characteristics of late-mature thermogenic gases (Milkov, 2021). Numerous case studies have indicated that the $\delta^{13}C_1$ in natural gas is primarily influenced by maturity (Stahl and Carey, 1975; Schoell, 1980; Dai, 1993). In contrast, the δ^{13} C values of heavier hydrocarbons such as C₂H₆ and C₃H₈ are less affected by maturity and mainly reflect the organic matter type of source rock. Based on the linear relationship between vitrinite reflectance (Ro) and the $\delta^{13}C_1$ of natural gas, empirical formulas for its variation have been deduced. Chen et al. (2021) comprehensively evaluated these formulas and pointed out that previous formulas often showed deviations when determining the maturity of gas source rocks. By combining actual geological conditions and a large amount of measured gas data, they revised the relationship between the $\delta^{13}C_1$ in oil-type gas and coal-derived gas and the vitrinite reflectance (Ro) of the source rock. The revised formulas are, for oil-type gas, $\delta^{13}C_1 = 25$ lgRo-42.5; for coal-derived gas, $\delta^{13}C_1 = 25$ lgRo-37.5. In this study, the above formulas were applied to estimate the maturity of gas source rocks in each production layer in the MX area of CRSB. The estimated maturities of gas from the Triassic and the $\mathcal{E}_1 l$ in the MX area are as follows, the maturity of $T_2 j^2$ gas ranges from 2.34% to 2.38% (average = 2.36%), that of $T_2 l^1$ gas ranges from 1.94% to 2.43% (average = 2.21%), and that of the $\epsilon_1 l$ gas ranges from 2.31% to 2.54% (average = 2.38%). According to the classification criteria for the thermal evolution stages of organic matter in source rocks, the gas source rocks of the Triassic and $\in_1 l$ in the MX area are mainly in the over-mature stage, with the gas maturities of $T_2 j^2$ and $\varepsilon_1 l$ being slightly higher than that of $T_2 l^1$. Correspondingly, the same conclusion can be drawn from the $\delta^{13}C_1$ vs δD_1 diagram (Figure 3D).

Based on the type of organic matter in source rocks, thermogenic gases can be classified into oil-type gas and coal-derived gas. The former is mainly generated from sapropelic organic matter, while the latter is primarily produced from humic organic matter (Dai et al., 1992). The data points of the samples fall within the range of shale-sourced rather than coal-sourced gases in the Milkov diagram (Figure 3B; Milkov, 2021) and follow the trend of gases generated from Type II kerogen in the modified Bernard diagram (Figure 3A; Bernard et al., 1978), exhibiting typical characteristics of oil-type gas. This shows significant differences from the coal-derived gas in the T_2x in the Zhongba area and the natural gas in the Feixianguan Formation sourced from the P_3l in the Longgang area of the Sichuan Basin. The hydrogen isotope composition of natural gas is influenced by the organic matter in the source rock, thermal maturity, and the conditions of the aqueous medium, and thus is commonly used to identify the origin of natural gas (Dai et al., 2012; Wang et al., 2015). Generally, the δD_1 values of oil-type gas are higher than those of coal-derived gas with the same or similar $\delta^{13}C_1$ values. The δD_1 values of $T_2 j^2$ and $T_2 l^1$ gas are generally heavier than -140‰, ranging from -144‰ to -120% (average = -131%). All samples show the characteristics of oil-type gas in the $\delta^{13}C_1$ vs δD_1 diagram (Figure 3D), and there

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TABLE 3 Geochemical data for the components and stable carbo

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References								Lıu et al. (2014)											Hu et al. (2013)				
δD/%,VSMOW	CH_4	-122	-122		-108	-116	1	-126	-123	-118	-123	-117	-116	-170	-170	-171	-171	-170	-171	-171	-170	-171	-173
	C_1-C_2	3.5	3.4	2.3	3.9	4.5			5.3	4.4	2.7	I		-11.0	-10.9	-11.0	-11.0	-11.1	-11.0	-11.3	-11.2	-10.0	-11.3
", VPDB	C ₂ H ₆	-37.3	-38.0	-37.4	-37.3	-36.9	-35.1	-41.8	-37.2	-36.2	-34.6	I		-24.0	-24.5	-24.4	-24.0	-24.4	-24.3	-24.3	-24.3	-24.8	-25.6
δ ¹³ C/%	CH_4	-33.8	-34.6	-35.1	-33.4	-32.4	-33.0	-37.0	-31.9	-31.8	-31.9	-31 3	-30.8	-35.0	-35.4	-35.4	-35.0	-35.5	-35.3	-35.6	-35.5	-34.8	-36.9
Dryness coefficient		0.9879	0.9827	0.9899	066.0	0.9899	0.9899	0.9832	0.9893	0.9920	0.9936	0.9929	0.9980	0.9190	0.920	0.9200	0.9180	0.9210	0.9200	0.9130	0.9200	0.8970	0.8850
	H_2S	0.08	0.04	0.01	0.05	0.14			0.12	0.42	0.40	0.34	0.77			I	I	I	I	I			
	N_2	3.26	0.79	0.98	0.67	0.52	0.37	4.58	0.76	0.67	0.43	1.61	0.68	0.63	0.53	0.21	0.91	0.28	0.46	0.49	0.27	0.28	0.03
ons/%	CO ₂	1.22	1.67	2.59	1.76	1.44	0.58	1.00	1.78	1.35	1.55	2.00	0.87	0.45	0.48	0.52	0.47	0.46	0.54	0.56	0.47	0.39	0.32
npositi	C_{3+}	0.10	0.42	0.09	0.07	0.11	0.12	0.34	0.09	0.05	0.02	0.03	0.01	2.20	2.07	2.15	2.23	2.09	2.14	2.46	2.11	3.53	5.02
ular cor	C_2H_6	1.06	1.27	0.88	0.84	0.88	0.88	1.24	0.95	0.73	0.61	0.66	0.19	5.81	5.70	5.75	5.79	5.75	5.75	6.10	5.77	6.53	6.36
Moleo	CH_4	94.36	95.82	95.46	96.66	97.05	98.05	92.72	96.41	97.20	97.39	96.17	97.47	90.36	90.80	90.90	90.19	91.00	90.53	89.80	90.82	87.86	87.82
Depth	E	5,194	4,588	4,977	5,171	3,850	2,660	3,744	5,035	4,388	4,806	5,046	4,035	2,602	2,408	2,628	2,510	2,366	2,578	2,446	2,400	2,268-2,361	2,422
Fm		C_2hl	C_2hl	C_2hl	C_2hl	C_2hl	C ₂ hl	C_2hl	C_2hl	C_2hl	C_2hl	C_2hl	C_2hl	$T_3 x^2$	$T_3 x^2$	$T_3 x^2$	$T_3 x^2$	$T_3 x^2$					
Well		Feng8	Tiandong9	Tiandong93	Yuedongl	W094	Xiang22	Jian32-1	Tiandong51	Guang19	Qili53	Oili28	Tieshan4	Z19	Z34	Z36	Z44	Z63	Z2	Z16	Z2	Z29	Z39
Gas field		GFC	LM	SPC	SPC	MLH	XGS	JN	WBT	SGP	TMC	TMC	TS					E	ZB				

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	E	Depth/m	Molec	ular con	npositid	%/suc			Dryness coefficient	δ ¹⁵ C/9	%。, VPDB	~	8D/%,VSMOW	References
			CH_4	C_2H_6	C ₃₊	CO ₂	N_2	H_2S		CH_4	C_2H_6	$C_{1}-C_{2}$	CH_4	
	P_2m	I	97.41	0.10	0.01	1.56	0.54	0.39	0666.0	-35.7	-33.4	-2.3	-136	
	P_2m	I	97.85	0.60	0.06	0.06	1.38	0.05	0.9930	-34.6	-32.5	-2.1	-137	
	P_2m	I	98.63	0.84	0.09	0.07	0.37	0.00	0.9910	-35.5	-33.4	-2.1	-135	Guo et al. (2024)
	P_2m	I	98.59	0.18	0.01	0	1.22	00.0	0.9980	-35.5	-35.2	I	I	
	I		91.19	5.11	2.41	0.85	0.44	0.00	0.9238	-31.2	-26.3	-4.9	-143	
	I		95.65	2.17	0.4	1.57	0.18	0.00	0.9738	-29.0	-27.6	-1.4	-190	
3	$T_{1}f^{3}$	6,348–6,365	96.32	1.54	0.34	1.36	0.40	0.00	0.9809	-29.3	-23.5	-5.8	-154	
_	$T_{1}f^{3}$	I	95.38	0.07	0.01	3.87	0.67	0.00	0.9992	-31.2	-27.3	-3.9	-141	
9	$T_{1}f^{3}$	I	95.25	0.2	0.03	3.9	0.62	0.00	0.9976	-29.0	-26.4	-2.6	-184	
5	$T_{1}f^{3}$	I	94.39	0.06	0.01	2.49	1.27	1.75	0.9993	-31.3	-26.6	-4.7	-128	
2	$P_2 ch$	6,261–6,828	92.35	60.0	0.00	5.67	0.28	1.61	0666.0	-31.6	-27.3	-4.3	-128	
8-1	$P_2 ch$	6,261–6,364	96.26	0.12	0.11	3.13	0.38	0.00	0.9976	-31.6	-27.0	-4.6	-124	
	$T_{1}f^{3}$	6,055–6,124	95.41	0.08	0.00	3.89	0.62	00.0	0.9992	-30.6	-24.3	-6.3	-150	Deng et al. (2018)
30	$P_2 ch$	6,238–6,346	88.40	0.06	0.1	7.68	0.39	3.37	0.9982	-30.5	-26.5	-4.0	-144	
32	$P_2 ch$	4,075–6,075	88.99	0.06	0.00	6.59	0.31	4.05	0.9993	-31.4	-25.2	-6.2	-136	
26	$P_2 ch$	6,480	88.42	0.06	0.00	7.39	1.97	2.16	0.9993	-30.0	-23.9	-6.1	-148	
	$T_{1}f^{3}$	5,536–5,796	93.77	0.06	1.39	4.63	0.00	0.15	0.9848	-30.9	-26.3	-4.6	-128	
	$T_{1}f^{3}$	4,781-4,889	94.13	0.06	0.01	4.34	0.57	0.89	0.9993	-31.4	-25.6	-5.8	-139	
	$T_{1}f^{3}$	5,170	90.71	0.08	0.53	6.32	0.51	1.85	0.9933	-31.4	I	-31.4	-130	
28	$P_2 ch$		91.22	0.06	0.01	7.8	0.21	0.70	0.9992	-30.8	-26.4	-4.4	-151	

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e components and stable carbon and hydrogen isotopes of natural gases for correlation in the Sichuan Basin. Molecular commonsitions /% DSMOW Baferances	CH_{A} $C_{3}H_{6}$ C_{3+} CO_{2} N_{2} $H_{3}S$ CH_{A} $C_{3}H_{6}$ $C_{1}-C_{2}$ CH_{4}	92.76 0.04 0.00 5.77 0.66 1.53 0.9996 -33.1 -29.3 -3.8 -137	90.47 0.04 0.00 7.52 1.00 0.86 0.9996 -32.9 -29.5 -3.4 -141	90.29 0.04 0.00 8.38 0.86 0.9996 -32.9 -28.6 -4.3 -139	92.75 0.05 0.00 4.49 0.88 1.99 0.9955 -33.9 -27.6 -6.3 -138	91.40 0.04 0.00 5.87 1.65 0.9996 -32.8 -28.3 -4.5 -147	90.19 0.04 0.00 8.3 0.73 1.49 0.9996 -33.1 -28.1 -5.0 -138	92.14 0.04 0.00 6.42 0.70 1.08 0.9996 -33.1 -27.6 -5.5 -139	91.22 0.04 0.00 6.35 1.36 1.05 0.9996 -32.3 -28.1 -4.2 -137	91.49 0.04 0.00 6.75 0.73 0.85 0.9996 -33.2 -28.8 -4.4 -136	90.04 0.03 0.00 8.15 0.81 1.05 0.997 -33.4 -28.2 -5.2 -144	82.65 0.04 0.00 14.19 2.12 0.97 0.995 -32.3 -27.8 -4.5 -137 Wet et al. (2015)	86.62 0.03 0.00 7.05 4.56 2.31 0.9997 -32.6 -28.0 -4.6 -149	91.21 0.03 6.41 1.72 0.81 0.9997 -33.6 -27.3 -6.3 -146	91.82 0.05 0.00 4.24 0.96 3.01 0.9995 -33.5 -28.8 -4.7 -141	96.80 0.14 0.00 2.26 0.63 0.9986 -32.4 -32.3 -0.1 -133	97.35 0.13 0.00 1.8 0.40 0.9987 -32.1 -33.6 1.5 -134	97.09 0.13 0.00 2.04 0.44 0.9987 -32.5 -32.4 -0.1 -133	95.98 0.13 0.00 2.53 0.55 0.9986 -33.4 134	
ar compositions /%	, H ₆ C ₃₊ CO ₂ N ₂ H	04 0.00 5.77 0.66 1	04 0.00 7.52 1.00 0.	04 0.00 8.38 0.8 0.	05 0.00 4.49 0.88 1.	04 0.00 5.87 1.65 0.	04 0.00 8.3 0.73 1.	04 0.00 6.42 0.70 1	04 0.00 6.35 1.36 1.	04 0.00 6.75 0.73 0.	03 0.00 8.15 0.81 1.	04 0.00 14.19 2.12 0.	03 0.00 7.05 4.56 2	03 0.00 6.41 1.72 0.	05 0.00 4.24 0.96 3	14 0.00 2.26 0.60 0.	13 0.00 1.8 0.69 0.	13 0.00 2.04 0.67 0.	13 0.00 2.53 0.72 0	
cal data for the components and Denth Molecula		92.76 0.	- 90.47 0.	- 90.29 0.	92.75 0.	5,109.5 91.40 0.	— 90.19 0.	- 92.14 0.	- 91.22 0.	— 91.49 0.	- 90.04 0.	- 82.65 0.	- 86.62 0.	— 91.21 0.	5,453 91.82 0.	- 96.80 0.	- 97.35 0.	4,885 97.09 0.	- 95.98 0.	
BLE 3 (Continued) Geochemic Cas field Well Em		MX12 $Z_2 dn^4$	MX13 $Z_2 dn^4$	MX17 $Z_2 dn^4$	MX11 $Z_2 dn^4$	MX8 $Z_2 dn^4$	GS3 $Z_2 dn^4$	GS2 $Z_2 dn^4$	GS1 Z ₂ dn ⁴	GS8 $Z_2 dn^4$	GS10 Z ₂ dn ⁴	$GS9 = Z_2 dn^2$	GS3 $Z_2 dn^2$	GS6 $Z_2 dn^2$	MX9 $Z_2 dn^2$	MX8 $\epsilon_1 l$	MX9 e_1l	MX11 e_1l	MX12 $\epsilon_1 l$	

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ICAM		Deptn/m	Moled	cular con	npositid	%/suc			Dryness coefficient	δ ¹³ C/9	%。, VPDB		8D/%,VSMOW	References
ICAM			CH₄	C_2H_6	C_{3+}	CO ₂	$^{\rm N}$	H_2S		CH_4	C_2H_6	$C_1 - C_2$	CH_4	
TEXTAT	$\epsilon_1 l$	I	95.21	0.27	0.00	3.93	0.28	0.16	0.9972	-33.5	-34.9	1.4	-132	
MX201	$\epsilon_1 l$	I	95.91	0.13	0.00	2.83	0.78	0.51	0.9986	-33.1	-33.0	-0.1	-133	
MX202	$\epsilon_1 l$	I	95.48	0.15	0.00	2.89	0.63	0.84	0.9984	-34.7	-35.3	0.6	-132	
MX204	$\epsilon_1 l$	I	96.63	0.13	0.00	2.06	0.71	0.40	0.9987	-32.6	-32.4	-0.2	-134	
MX205	$\epsilon_1 l$	Ι	95.30	0.2	0.00	3.18	0.42	0.73	0.9979	-33.2	-34.8	1.6	-132	
MX008-	HI $\epsilon_1 l$	I	95.15	0.14	0.00	3.34	0.70	0.52	0.9985	-32.2	-33.3	1.1	-136	
-600XW	X1 $\epsilon_1 l$	I	96.50	0.14	0.00	2.37	0.67	0.61	0.9986	-33.0	-33.3	0.3	-137	
MX13	$\epsilon_1 l$		95.44	0.13	00.0	1.65	0.70	0.50	0.9986	-32.7	-33.0	0.3	-132	

is a positive correlation between δD_1 and $\delta^{13}C_1$, indicating that as gas maturity increases, δD_1 is significantly affected and shows a gradually increasing trend. The salinity of water during source rock formation affects the hydrogen isotope composition of natural gas. Compared with fresh water, saltwater is rich in deuterium (D), which usually results in higher δD_1 values in natural gas generated in such environments (Schoell, 1980; Wang et al., 2015; Ni et al., 2019). The samples can be roughly divided into two parts based on δD_1 values, gas with $\delta D_1 > -132\%$ mainly comes from T_2j^2 , while gas with δD_1 < -132% is from \mathcal{C}_1l and T_2l^1 . Considering that the source rocks in these areas are generally in the high to over – mature stage, the increase in water salinity during source rock deposition rather than maturity may be the reason for the higher δD_1 values in gas. The wide distribution of data points for T_2l^1 gas may be due to the complexity of the gas source rocks or the existence of mixing effect.

Given that the C₃H₆ content in certain samples is below the instrument's detection limit, registering as 0%, the $\ln (C_2/C_3)$ vs $\ln (C_1/C_2)$ plot is inapplicable for differentiating kerogen-cracking gas from oil-cracking gas in this study. On the C_1/C_n vs $\delta^{13}C_1$ diagram (Figure 3C), the $\delta^{13}C_1$ values of all samples are less than -30‰, indicating that they belong to the category of oil-cracking gas. These samples can be compared with natural gas from other regions that has been confirmed as oil-cracking gas, such as the Z_2 dn gas in the GST-MX area and the $\varepsilon_1 l$ gas in the MX area (Wei et al., 2015), the C_2hl gas in eastern Sichuan (Liu et al., 2014), and the P_2m gas in the HWC and WJB gas fields in northwestern Sichuan (Guo et al., 2024). They differ from the coal-derived gas in the $P_2ch-T_1f^3$ Fms in the LG gas field that displays characteristics of kerogen-cracking gas. In addition, the relative contents of non-hydrocarbon gases CO2 and N2 can also be used to identify the formation pathways of natural gas (Liu et al., 2012), which is related to the high content of nitrogen-containing compounds in argillaceous source rocks. During the thermal evolution of argillaceous source rocks, denitrification occurs, resulting in a higher N2 content in kerogen-cracking gas than in oil-cracking gas. According to the CO2-N2 relationship diagram proposed by Li et al. (2022), the boundary of the N_2 content between oil-cracking gas and kerogen-cracking gas in the Sichuan Basin is approximately 4%. Samples with an N2 content higher than 4% in natural gas can generally be considered as kerogen-cracking gas. Therefore, all the natural gas in the Triassic marine strata of the MX area is oil-cracking gas. However, according to the LHs discrimination diagrams (Figures 4A, B), most $T_1 l^1$ samples have relatively low MCH/nC7 ratios, and the data points fall into the mixed - gas region of oil - cracking gas and kerogen-cracking gas. It is believed that the mixing of gas with a relatively high MCH content in the gas of the $T_1 l^1$ causes deviations in the results.

On the $\delta^{13}C_{2-1}$ vs $\delta^{13}C_1$ plot (Figure 3B), some gases from the Triassic and the C_1l Fm are located in the reversal region ($\delta^{13}C_1 > \delta^{13}C_2$). Considering the distribution of several potential source rocks (Figure 1) and the high maturity of the natural gas, there are several reasonable explanations for the observed phenomenon of $\delta^{13}C_1 > \delta^{13}C_2$. (1) Mixing of hydrocarbon gases generated by thermal cracking of oils from different source rocks. In the CRSB, there are three potential source rocks, the C_1q marine shale, the Upper Permian marine shale, mudstone, and carbonate rocks are mainly composed of sapropelic organic matter and may thus

represent "not available"



FIGURE 3

Relationship diagram between (A) $\delta^{13}C_1 \text{ vs. } C_1/(C_2+C_3)$. Abbreviations: SM–secondarymicrobial, EMT–early mature thermogenic gas, LMT–late mature thermogenic gas [adapted from Bernard et al. (1978); Milkov et al. (2020)] (B) $\Delta\delta^{13}C_2 - \delta^{13}C_1$ vs $\delta^{13}C_1$ [adapted from Milkov (2021)], (C) $C_1/C_n \text{ vs. } \delta^{13}C_1$. [adapted from Clayton (1991)] and (D) $\delta^{13}C_1 \text{ vs} \delta^{13}C_1$ [adapted from Wang et al. (2015)] of $T_2 j^2$, $T_2 l^1$ gas in the MX gas field and $\mathcal{E}_1 l$ gas in the AY gas field, Sichuan Basin (Data source same as Table 3).



Cross-plot of MCH/nC₇ versus MCH/CH (A) and MCH/nC₇ versus (2–MH+3–MH)/nC₆ (B) of $T_2 j^2$, $T_2 l^1$ gas in the MX gas field and $\mathcal{C}_1 l$ gas in the AY gas field [modified from Hu et al. (2005)].

contribute to the accumulated oils that generate secondary-cracking gases. Although existing data suggest that the accumulated oils in the $T_2 l^1$ gas reservoir are mainly sourced from the Upper Permian source rocks (Wang et al., 1998), the contribution of other source rocks (such as the Lower Triassic marine shale) cannot be excluded. (2) Mixing of secondary-cracking gases and primary-cracking gases. As mentioned earlier, all samples are secondary-cracking gases (Figure 3C), while this explanation is not applicable to the Triassic gas in CRSB. (3) Mixing of gases from thermal cracking of previously-accumulated oils at different maturity levels may occur within the reservoir. However, Triassic dry gas reservoirs did not evolve from paleo oil reservoirs within the stratigraphic system. Thus, the gases likely accumulated instantaneously. Nevertheless, it cannot be ruled out that the mixture of gases generated by thermal cracking of accumulated oils at different maturity levels in the Lower Cambrian migrated upward and retained their reversed characteristics.

5.2 Gas-source correlation

The distribution of various sets of source rocks has been discussed in the geological setting. The $\epsilon_1 q$ Fm, $P_2 l$ Fm and $T_2 l^3$ Fm are three sets of source rocks that may contribute to the natural gas in the Triassic marine strata. The $\mathcal{E}_1 q$ source rock is controlled by the Deyang-Anyue rift trough and has a widespread distribution (Figure 2B). It has strong hydrocarbon - generating potential and is one of the highest-quality source rocks in the Sichuan Basin. Its gas-generation intensity reaches 20×10^8 - 200×10^8 m³/km² (Zhang et al., 2024). However, the gypsum layers developed in the overlying Cambrian Gaotai Formation ($\mathcal{E}_1 g$) and $\mathcal{E}_1 l$ are distributed in the eastern and southern Sichuan Basin, and they do not play a capping role for the hydrocarbon–generation of the $\in_1 q$ in the MX area of CRSB (Xu et al., 2016). In addition, the multi-stage strike-slip faults developed in the CRSB serve as "highways" for oil and gas migration. These faults have been active in multiple stages from the Sinian to the Cretaceous. The Indosinian period was the main stage of oil - generation from the Lower Cambrian shale, and traps along the migration pathways may have captured the oil and gas generated in historical periods (Shuai et al., 2022; Lu et al., 2024).

The P₂*l* source rock has sedimentary centers in both central and northeastern Sichuan Basin. The maximum sedimentary thickness can reach 120–140 m. The superior source rock conditions have laid a material foundation for hydrocarbon generation in the central Sichuan region. However, according to the burial history of Well GS 17, the P₂*l* source rock in the CRSB entered the high–maturity stage (geo–temperature>180°C) during the Cretaceous period (Figure 5; Yang et al., 2020). Subsequently, under the influence of the uplift of the Indosinian movement, the temperature dropped rapidly. Therefore, it did not experience a large scale dry gas generation stage in history. As can be seen from the hydrocarbon–expulsion intensity map of the P₂*l* Fm at the end of the Triassic, the hydrocarbon–expulsion amount is less than 50 × 10⁸ t/km² south of Suining in the CRSB (Zhang et al., 2022).

As for the Triassic T_2l^3 marl, it is a set of lagoonal facies source rocks explored in recent years. The average TOC of the T_2l^3 cuttings in Well CT1 is 0.77%, the organic matter type is Type II, and the measured Ro is 1.59–1.63%, indicating a high mature stage (Yu, 2022; Wang et al., 2023). Currently, only the self–generated and self–stored reservoir of the T_2l^3 marl in Well CT1 has been discovered, and no contribution to conventional gas reservoirs has been found yet.

The $T_2 j^2$ and $T_2 l^1$ gases are oil-type gases and specifically oil-cracking gases. The $\delta^{13}C_{kerogen}$ of $P_2 l$ ranges from -24‰ to -23‰, and the organic matter type is II₂-III (Huang et al., 2024). According to the carbon isotope fractionation between oil-cracking ethane and kerogen is larger than that between kerogen-cracking ethane and kerogen, the carbon isotope of oil-cracking ethane in oil-cracking gas is generally 3‰-5‰ lighter than that of kerogen (Dai et al., 2005). Consequently, the carbon isotope of ethane in $P_2 l$ gas should range from -29‰ to -26‰, which is inconsistent with the ethane characteristics of the gas in the MX Gas Field. Moreover, the $P_2 l$ gas in the CRSB has not entered the over mature stage, while the dryness coefficients of $T_2 j^2$ and $T_2 l^1$ gases are very high. Therefore, the $P_2 l$ source rock is unlikely to be the main source rock in this area. Thus, $\mathcal{E}_1 q$ is the most important source rock for the Triassic gas reservoirs in the CRSB.

By comparing the characteristics of $\delta^{13}C_1$ and $\delta^{13}C_2$ (Figure 3A, B), the Triassic gases are found to closely resemble the P_2m gases from the WJB and HWC in western Sichuan, and the C_1l gas. Previous research indicates that these gases share a common origin from the C_1q source rock. Specifically, the P_2m gases are oil–cracking gases generated during the early phase of the C_1q 's hydrocarbon generation process, while the C_1l gas represents shale gas released in the late stage of the C_1l maturation (Shuai et al., 2022; Guo et al., 2024). As illustrated in the $\delta^{13}C_{2-1}$ vs. $\delta^{13}C_1$ diagram (Figure 3B), in contrast to oil – cracking gases, the late – stage shale gas in the C_1l Fm exhibits distinct characteristics. It has heavier methane isotopes, reversed carbon isotope ratios between methane and ethane ($\delta^{13}C_1 > \delta^{13}C_2$), and a notably high dryness coefficient (>0.9995), as shown in Figure 3C.

LHs are a crucial component of oil and natural gas, generally referring to C5-C10 compounds with a boiling point less than 200°C. They contain abundant geochemical information and are often used for origin identification (Hu et al., 2007; Dai et al., 2016), organic matter type determination (Hill et al., 2007), and oil-gas source correlation. Sapropelic parent materials are rich in normal alkanes, while humic parent materials are abundant in isoparaffins, naphthenes, and aromatics (Leythaeuser et al., 1979; Snowdon and Powell, 1982). The triangular diagram of C₆₋₇ normal alkanes, isoparaffins, and naphthenes is commonly used to characterize the genetic types of natural gas. In the triangular diagram of C₇ compounds, MCH is derived from the lignin, cellulose, and sugars of higher plants and has stable thermodynamic properties. DMCP comes from the cyclic lipid compounds in sterols and terpenoids of aquatic organisms, and its abundant presence is an important characteristic of oil-type gas. The parent sources of nC7 are complex, mainly including bacteria, algae, and the chain-like complexes of higher plants. However, it is quite sensitive to maturity and will undergo extensive cracking at high maturity levels (Dai et al., 1992; Odden et al., 1998; Whiticar, 1999; Mango, 2000; Hu et al., 2007). When natural gas reaches the over mature stage, such as the shale gas in the S_1l in southern Sichuan, light hydrocarbons undergo intense thermal cracking. The normal alkanes that originally reflect the sapropelic parent material source are destroyed, and the contents of compounds with strong thermochemical stability, such as naphthenes and aromatics, increase significantly. This distorts



FIGURE 5

Burial and thermal histories, and hydrocarbon charging stages from Well GS17 in CRSB [Burial history and thermal evolution history are modified from Yang et al. (2020)].



the identification results of LHs, causing them to fall into the coal-derived gas region (Dai et al., 2016; Figures 6A, B).

In the genetic identification chart of the LHs triangular diagram, the gases of $T_2 j^2$ and $C_1 l$ show a high degree of kinship correlation and also have a MCH advantage (Figures 7A–D), falling into the coal–derived gas region. However, the relative contents of normal and isoparaffins in $T_2 l^1$ gas are relatively high (Figure 7E–H), and the majority of samples fall into the identification region of oil–type gas. Whiticar and Snowdon (1999) proposed that the individual carbon isotopes of LHs from similar organic sources are highly correlated. In this study, the carbon isotope values of seven LHs compounds, namely, 2–MP, 3–MP, nC₆, MCP, CH, nC₇, and MCH, were compared (Table 2; Figure 8B) to observe whether there were differences in gas sources. Ignoring the isotope differences among different LHs compounds, the individual carbon isotope fingerprint diagram of LHs indicates that the T₂l¹ LHs should be from the same gas source, with their $\delta^{13}C_{MCP}$ >–24‰ and $\delta^{13}C_{MCP}$ > $\delta^{13}C_{CH}$. In contrast, the T₂l² and $\epsilon_1 l$ gases exhibit the characteristics of $\delta^{13}C_{CH}$



FIGURE 7

FIGURE 7 Gas chromatogram of LHs from natural gas from the $T_2 j^2$, $T_2 l^1$ in MX gas field and the $\mathcal{E}_1 l$ in AY gas field. In the MX gas field (2–MP, 2–methylpentane; 3–MP, 3–methylpentane; nC₆, hexane; MCP, methylcyclopentane; Bz, benzene; CH, cyclohexane; 2–MH, 2–methylhexane; 3–MH, 3–methylpentane; Ω_2 nC₇, heptane; MCH, methylcyclohexane; Tol, toluene). (A) M201, 4,630 m, C_1l . (B) MX009–X1, 4,750 m, C_1l . (C) MS1, 3,128.9 –3,180 m, T_2l^2 . (D) M005-H9, T₂/². (E) M017-H3, 3,200 m, T₂/¹. (F) M91, 2,676, T₂/¹. (G) M89, 2,723.2 m, T₂/¹. (H) M030-H14, 3,670 m, T₂/¹.



>-24% and $\delta^{13}C_{MCP} < \delta^{13}C_{CH}$. Based on this, it can be judged that the gas sources of $T_2 j^2$ and $T_2 l^1$ are different, and both $T_2 j^2$ and $\mathcal{C}_1 l$ gases are derived from the shale gas of $\mathcal{C}_1 q$.

The Mango coefficient K₁, proposed by (Mango, 1997), demonstrates a consistent K value within the same category of oil and gas, yet varies across different types. This characteristic renders it a reliable criterion for gas–source correlation. When examining the diagrams of $(2-MH + 2.3-DMP)/nC_7$ versus $(3-MH + 2.4-DMP)/nC_7$ (Figures 4A, B), distinct correlations emerge. For the T₂ l^2 and $C_1 l$ gas samples, the correlation equation is y = 1.01x-0.07, with a high correlation coefficient of $R^2 = 0.9865$, and the average K₁ value of the natural gas is 1.11. In contrast, for the T₂ l^1 gas, the correlation relationship is described by y = 1.18x-0.225, with a correlation coefficient of $R^2 = 0.9384$, and the average K₁ value of the natural gas is 1.28. These differences in correlation equations and K₁ values can potentially provide insights into the unique characteristics and origins of each gas sample (Figure 8A).

Based on the above evidence of LHs, it is considered that the sources of $T_2 j^2$ and $T_2 l^1$ natural gases are not completely the same. Although both gases are derived from the $C_1 q$, the gas in $T_2 l^1$ has the admixture of other oil-type gases, and the maturity of this mixed gas is lower than that of the shale gas from the $C_1 q$ Fm. However, considering the characteristics of CH₄ and $C_2 H_6$ in natural gas, $T_2 l^1$ gas is significantly different from the gas from the $P_2 l$ Fm (LG) and is more similar to the oil-cracking gas and shale gas from the $C_1 q$. It is believed to be formed by the mixing of oil-cracking gas with lower maturity and shale gas. This also well explains the reversal phenomenon observed in the gas from the $T_2 l^1$.

6 Conclusion

1. The natural gas components in the Triassic gas reservoirs of the MX gas field in central Sichuan are mainly alkane gases. The gas is extremly dry (dryness coefficients ranges from 0.9970 to 0.9985) with a low content of non–alkane gases, mainly being $\rm CO_2$ and $\rm N_2$. All the gas is of oil–type gas and belongs to crude oil–cracking gas.

- 2. The $\mathcal{C}_1 q$ Fm is the principal source rock, generating different types of natural gas. It includes both oil–cracking gas (reservoir gas) and shale gas (unaffected by TSR). These two types of gas exhibit diverse geochemical characteristics. The oil–cracking gas is characterized by relatively light $\delta^{13}C_1$ (<-35‰), a relatively low dryness coefficient (<0.9995), and a normal sequence of carbon isotopes ($\delta^{13}C_1 < \delta^{13}C_2$). The shale gas has heavy $\delta^{13}C_1$ (>-35‰), a high dryness coefficient (>0.9995), and shows carbon isotopes reversal ($\delta^{13}C_1 > \delta^{13}C_2$).
- 3. The shale gas, with a higher thermal evolution degree, has low abundances of normal and iso–alkanes and a higher proportion of cycloalkanes. Its has the feature of individual carbon isotopes with $\delta^{13}C_{CH}$ >–24‰ and $\delta^{13}C_{MCP}$ < $\delta^{13}C_{CH}$, while the oil–cracking gas shows $\delta^{13}C_{MCP}$ >–24‰ and $\delta^{13}C_{MCP}$ > $\delta^{13}C_{CH}$.
- 4. The $T_2 j^2$ gas comes from the shale gas of the $\mathcal{E}_1 q$ Fm. The $T_2 l^1$ gas is a mixture of oil–cracking gas and shale gas from the $\mathcal{E}_1 q$ Fm, and this mixture is the main cause of the carbon isotope reversal in the $T_2 l^1$ gas.

Data availability statement

The original contributions presented in the study are included in the article/supplementary material, further inquiries can be directed to the corresponding authors.

Author contributions

CG: Conceptualization, Formal Analysis, Funding acquisition, Investigation, Methodology, Resources, Visualization, Writing-original draft, Writing-review and editing. GH: Data curation, Funding acquisition, Investigation, Methodology, Project administration, Resources, Writing-review and editing. JL: Methodology, Writing-review and editing. XT: Data curation, Investigation, Resources, Supervision, Writing-review and editing. LT: Data curation, Methodology, Software, Supervision, Visualization, Writing-review and editing. JG: Formal Analysis, Methodology, Writing-review and editing. HC: Data curation, Resources, Writing-review and editing. ZL: Methodology, Writing-review and editing. XL: Funding acquisition, Supervision, Writing-review and editing.

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Conflict of interest

Authors CG, GH, JL, LT, JG, HC, and ZL were employed by PetroChina Research Institute of Petroleum Exploration and Development. Author XT was employed by PetroChina Southwest Oil & Gas field Company.

The remaining author declares that the research was conducted in the absence of any commercial or financial relationships that could be construed as a potential conflict of interest.

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